

Pattern recognition of images  
in nuclear emulsion and solid detectors  
experiments

on PAVICOM facility

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Laboratory of elementary particles

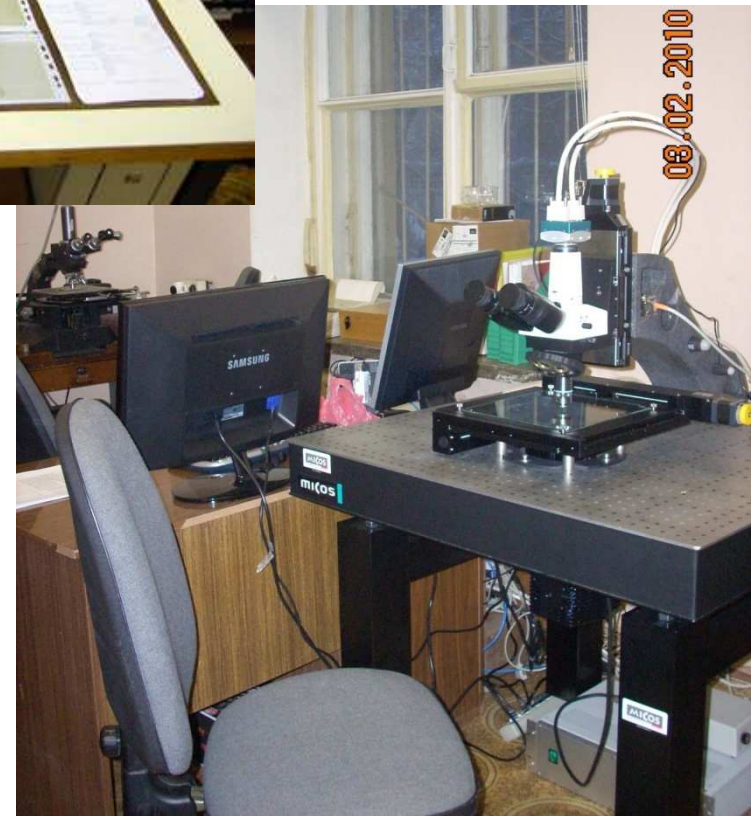
PAVICOM 2  
(X=12 cm, Y=10 cm)



PAVICOM 1  
(X=80 cm, Y=40 cm)



PAVICOM 3  
(OPERA)

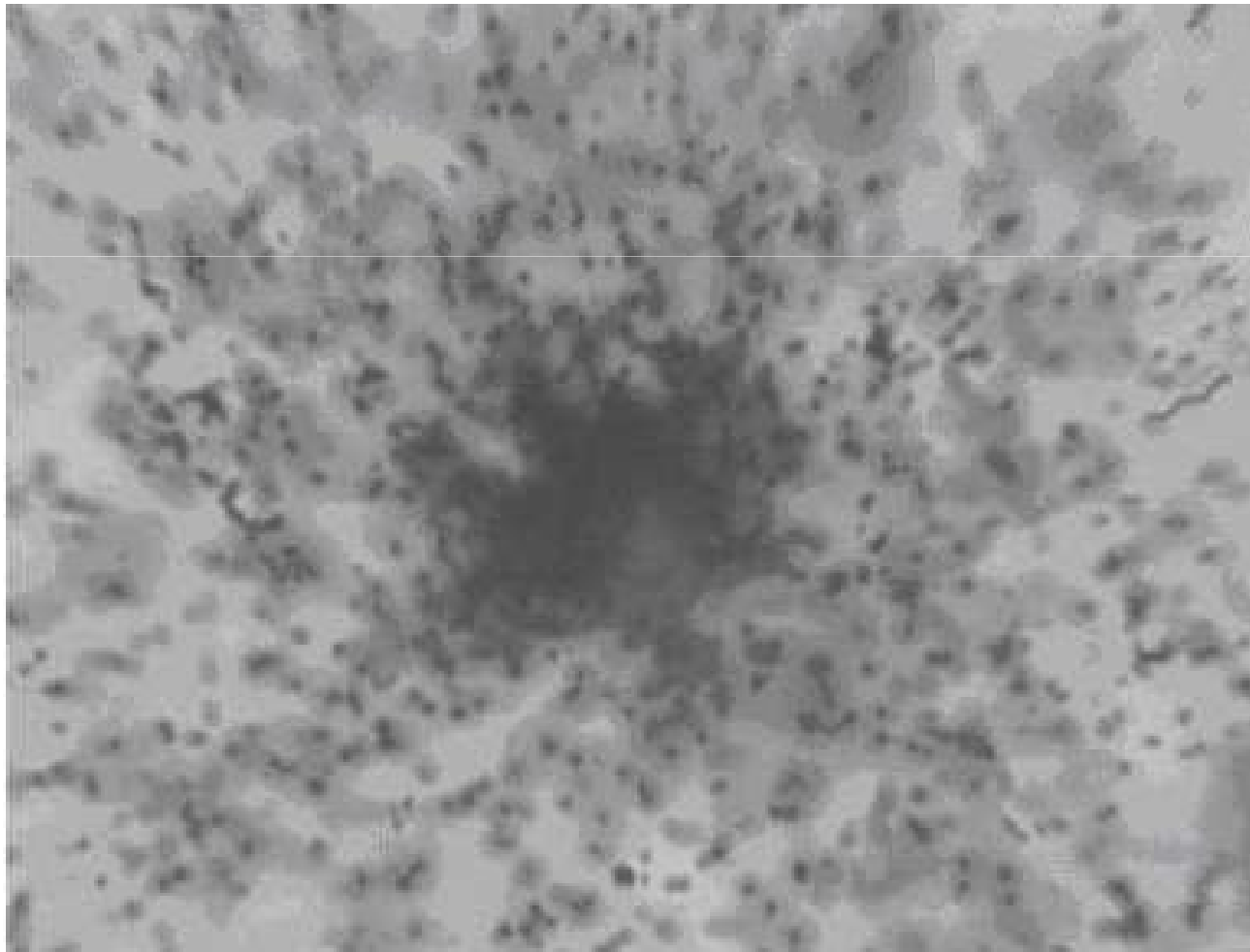


EMU-15.

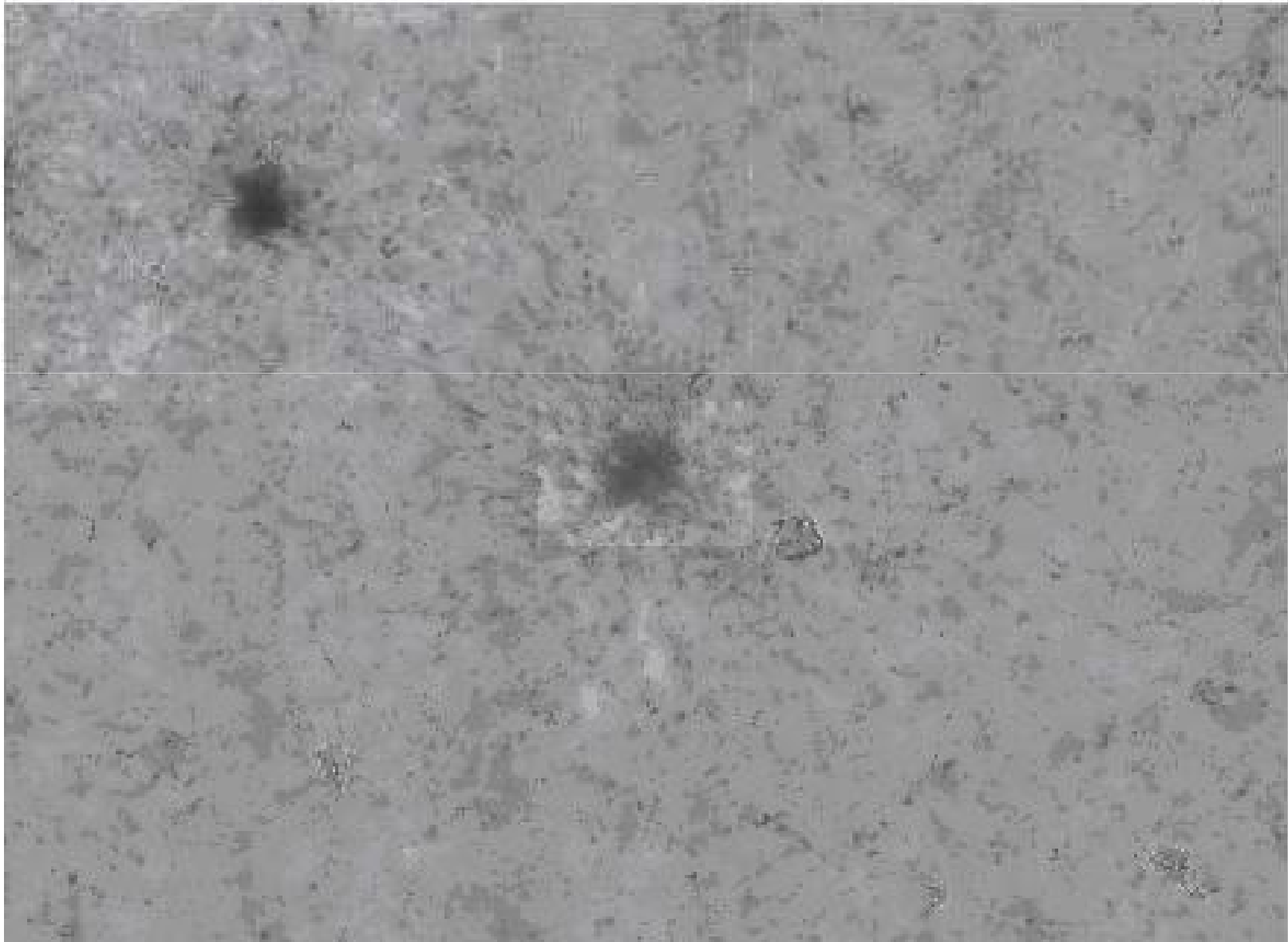
Pb+Pb(fixed) 158 GeV/A (CERN SPS).

Multiplicity is about 1000 particles.

Central field of view

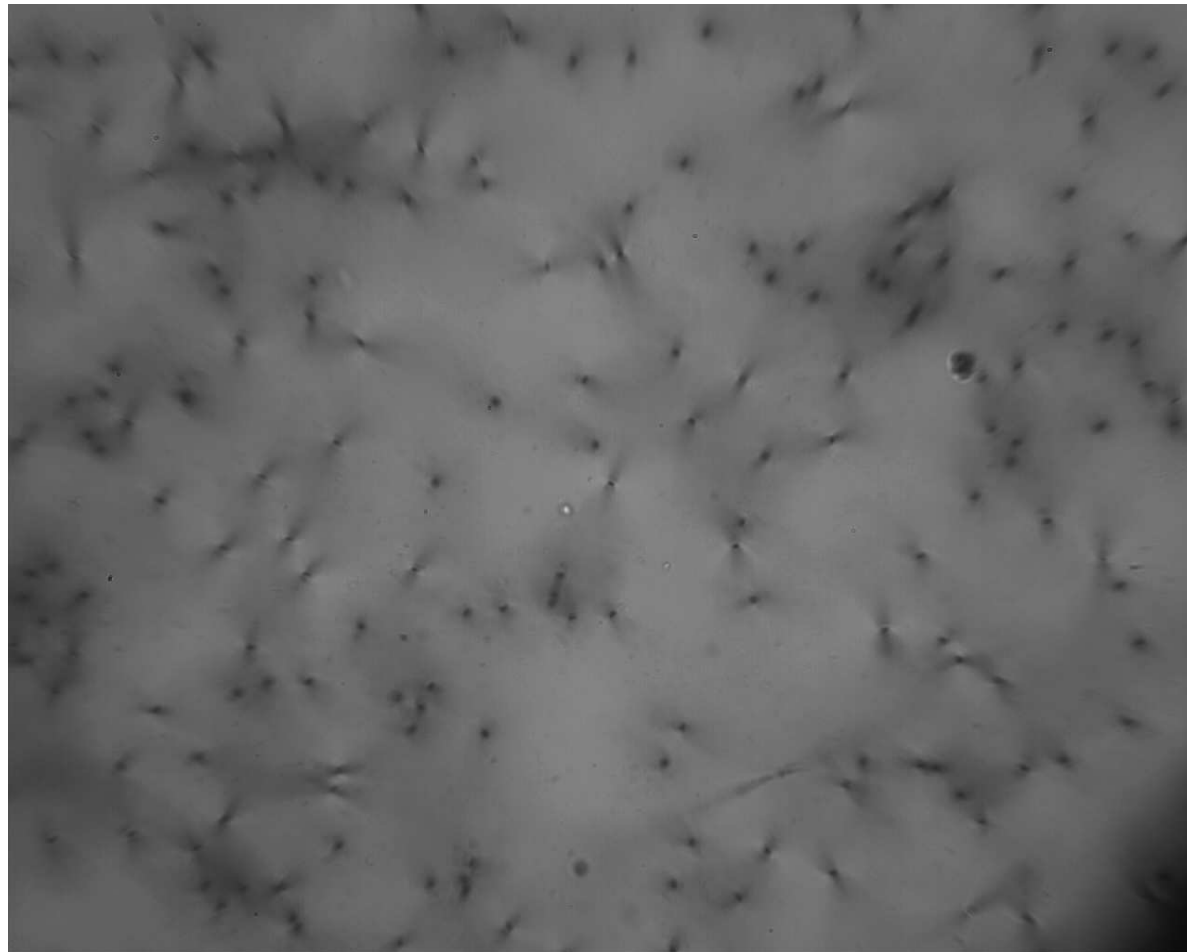


25 fields of view  
(track reconstruction, rapidity distribution)  
60x80 mcm



# Interaction of ${}^6\text{He}$ ( $E=60-20 \text{ MeV}$ ) in emulsion

(tracks reconstruction, looking for of vertexes, kinematics)  
320x280 mcm

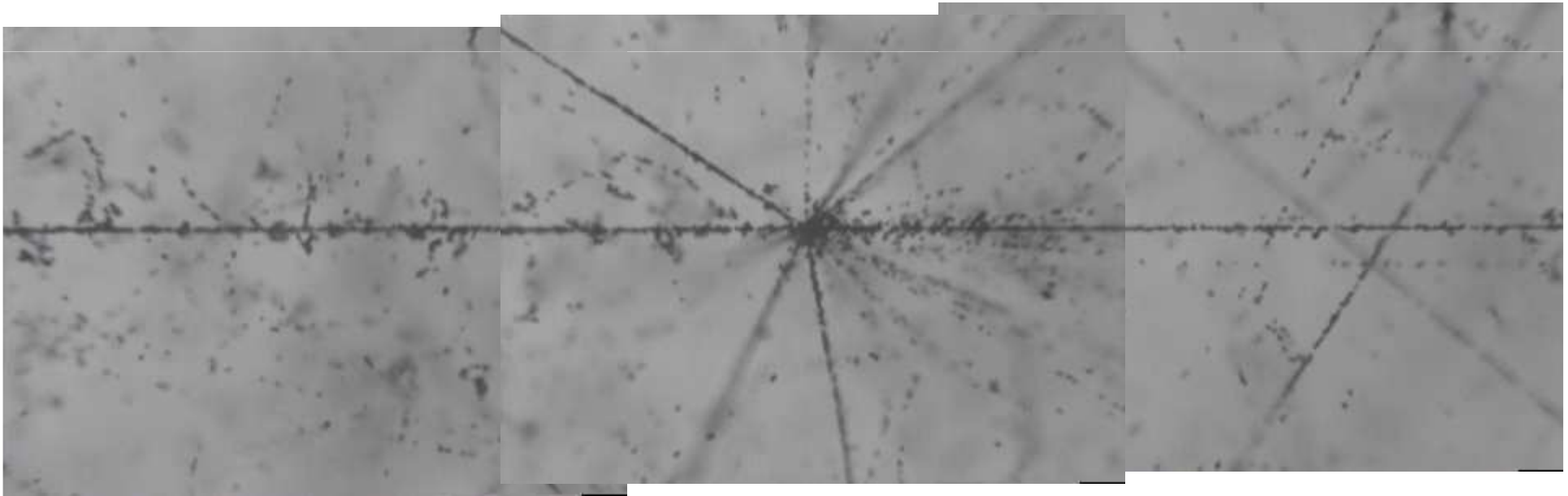


# Trace of relativistic nucleus in thick emulsion

(thickness is 600  $\mu\text{m}$ )

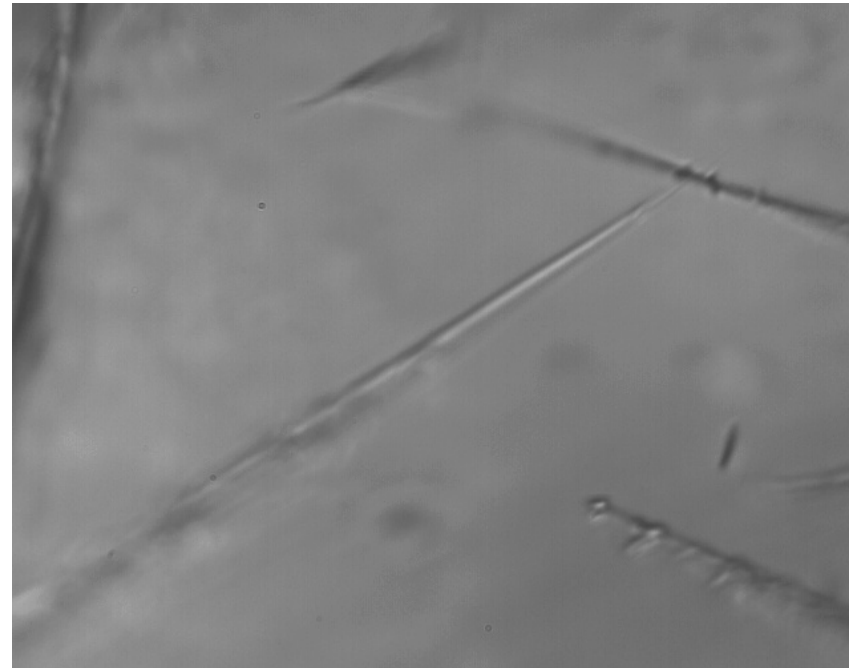
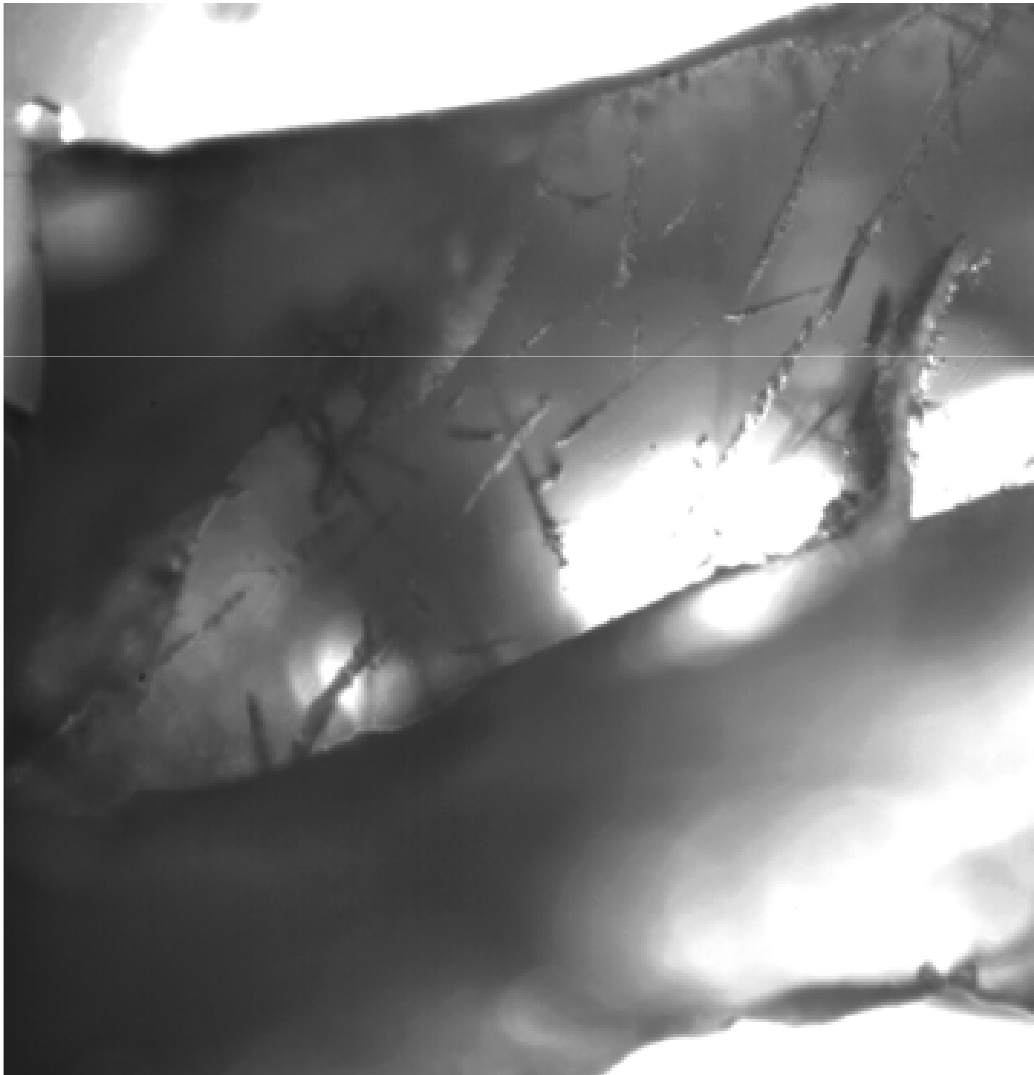
(definition of track characteristics and charge  
of relativistic nuclei)

magnification is 40x



# OLIMPIYA

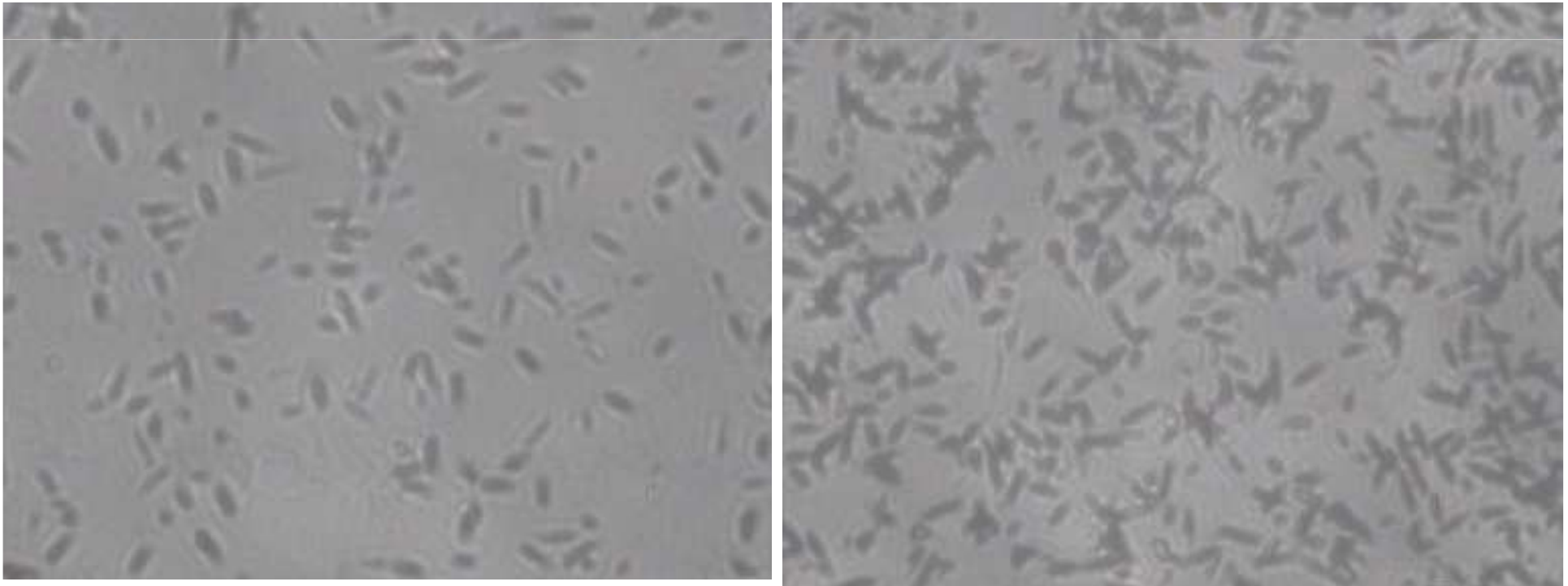
Looking for of super heavy nuclei traces  
in olivine from meteorites  
(track reconstruction, charge evaluation )



# ENERGY+TRANSMUTATION

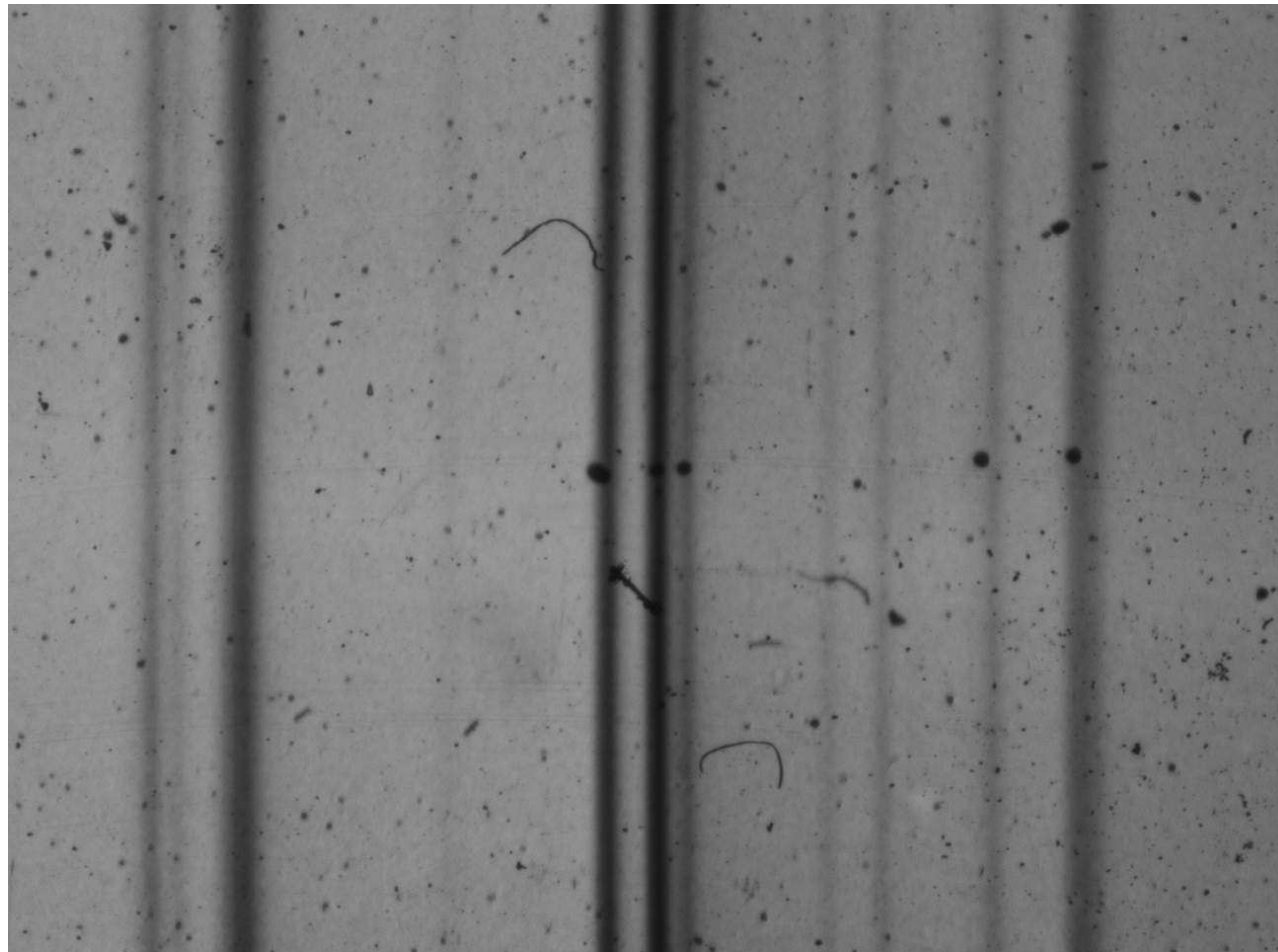
(evaluation of neutron flux using plastic detectors)

magnification is 20x

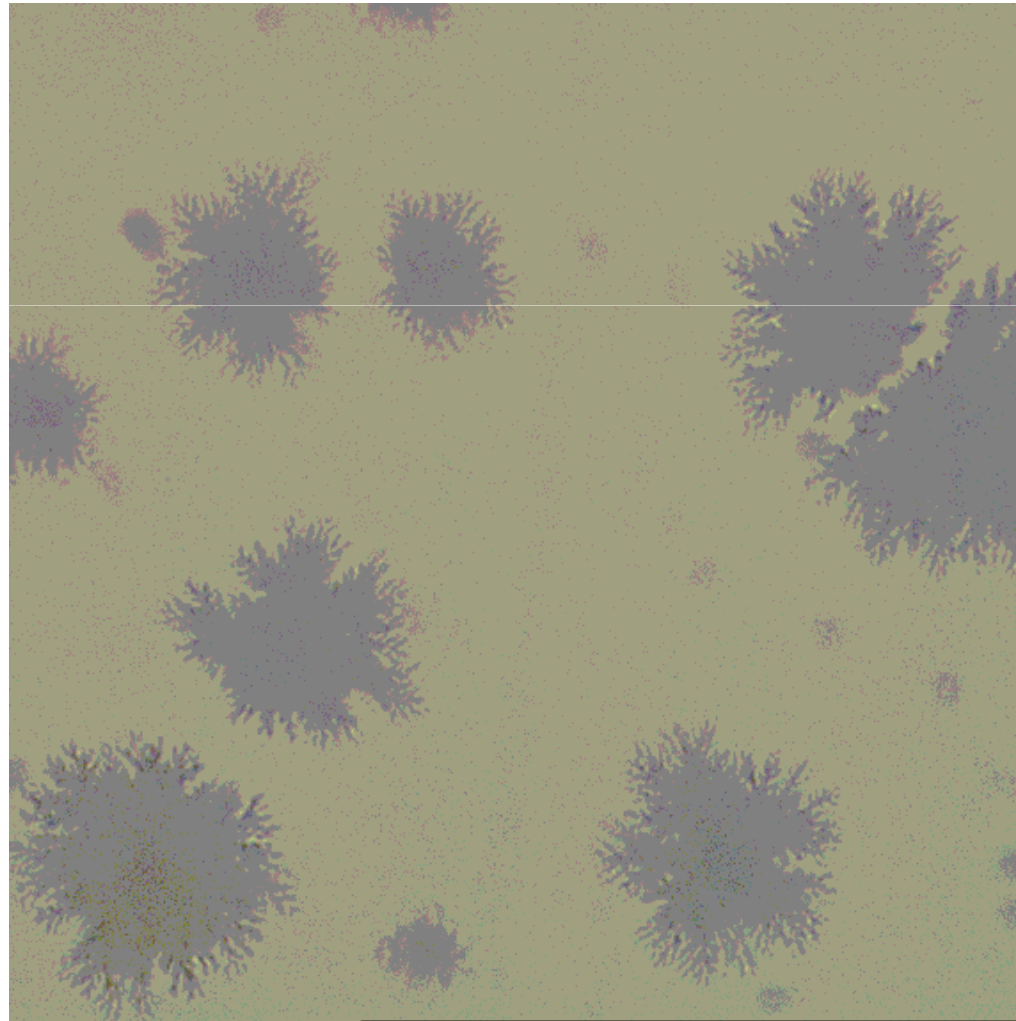




Investigation of inner conversion electrons  
of lanthanide nuclei  
(line identification, looking for new lines)  
magnification is 8x



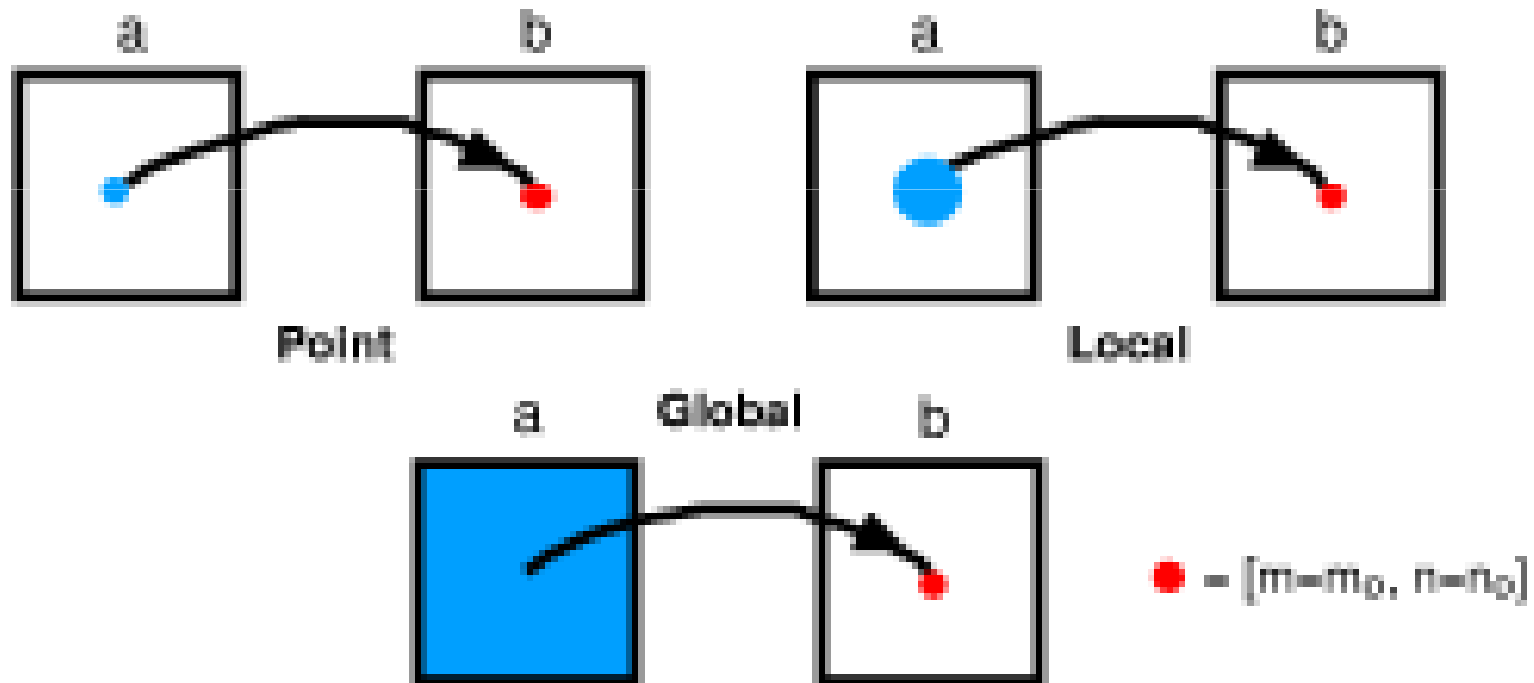
Traces of alpha particles in plastic after  
electro-chemical etching  
(calculation of tracks number)  
500x500 mcm



## The list of image processing stages

1. Preliminary transformation (filtering, logical operations, histogram analysis and go on).
2. Clustering (extraction, characteristics calculations).
3. Division of clusters into classes (tracking and other).
4. Informative analysis according to physical task.

# Three kinds of preliminary image transformation



# Linear filters:

$$P'_{mn} = \sum_{i=0}^{N-1} \sum_{j=0}^{N-1} P_{mn} \cdot h_{m-i, n-j}$$

## Smoothing

$$h_{circ}[j,k] = \frac{1}{21} \begin{bmatrix} 0 & 1 & 1 & 1 & 0 \\ 1 & 1 & 1 & 1 & 1 \\ 1 & 1 & 1 & 1 & 1 \\ 1 & 1 & 1 & 1 & 1 \\ 0 & 1 & 1 & 1 & 0 \end{bmatrix} \cdot \frac{1}{81} \begin{bmatrix} 1 & 2 & 3 & 2 & 1 \\ 2 & 4 & 6 & 4 & 2 \\ 3 & 6 & 9 & 6 & 3 \\ 2 & 4 & 6 & 4 & 2 \\ 1 & 2 & 3 & 2 & 1 \end{bmatrix}$$

## Laplacian

$$[h] = \begin{vmatrix} -1 & -2 & -3 & -2 & -1 \\ -2 & 2 & 5 & 2 & -2 \\ -3 & 5 & 20 & 5 & -3 \\ -2 & 2 & 5 & 2 & -2 \\ -1 & -2 & -3 & -2 & -1 \end{vmatrix}$$

## Gradient

### 1. Usual

$$[h_x] = \frac{1}{3} \begin{bmatrix} 1 & 0 & -1 \\ 1 & 0 & -1 \\ 1 & 0 & -1 \end{bmatrix}$$

$$[h_y] = \frac{1}{3} \begin{bmatrix} 1 & 1 & 1 \\ 0 & 0 & 0 \\ -1 & -1 & -1 \end{bmatrix}$$

### 2. Sobel's

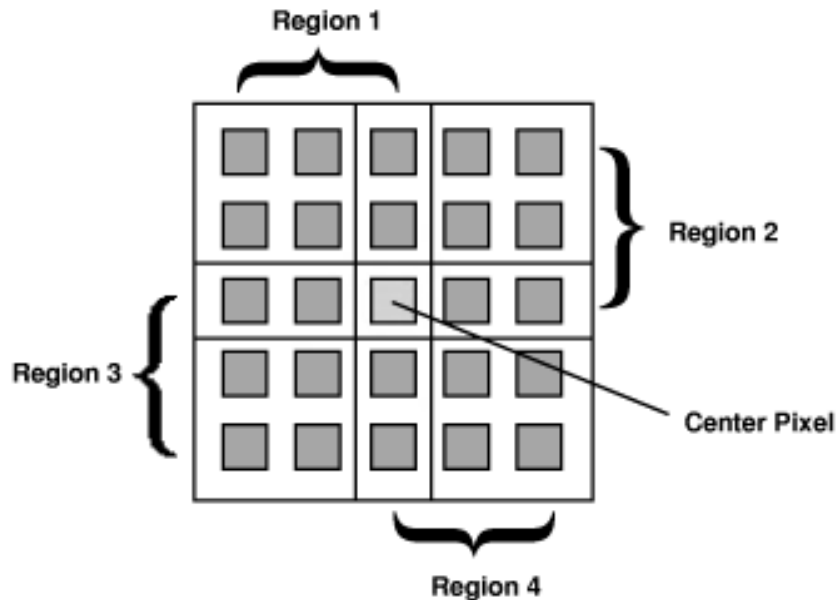
$$[h_x] = \frac{1}{4} \begin{bmatrix} 1 & 0 & -1 \\ 2 & 0 & -2 \\ 1 & 0 & -1 \end{bmatrix}$$

$$[h_y] = \frac{1}{4} \begin{bmatrix} 1 & 2 & 1 \\ 0 & 0 & 0 \\ -1 & -2 & -1 \end{bmatrix}$$

# Nonlinear transformation

$$T_{ij}^{\text{new}} = T_{ij}^{\text{old}} + K_{\text{br}} \cdot F_{\text{br}}(T_{ij}^{\text{old}}) \quad \text{Brightness,}$$

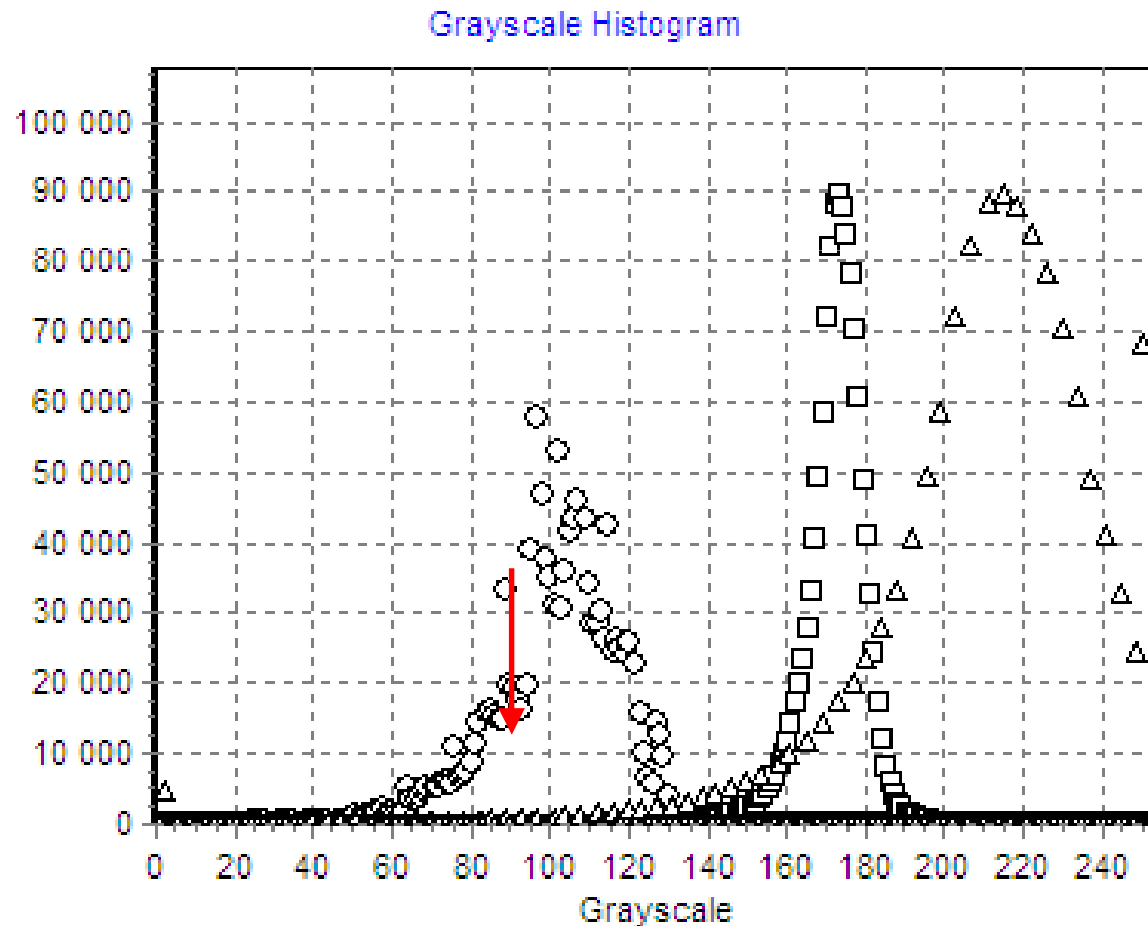
$$T_{ij}^{\text{new}} = T_{ij}^{\text{old}} + K_{\text{con}} \cdot F_{\text{con}}(T_{ij}^{\text{old}} - T^{\text{fix}}) \quad \text{Contrast}$$



Kuwahara's  
nonlinear filter

## Brightness histogram ( ${}^6\text{He}$ ):

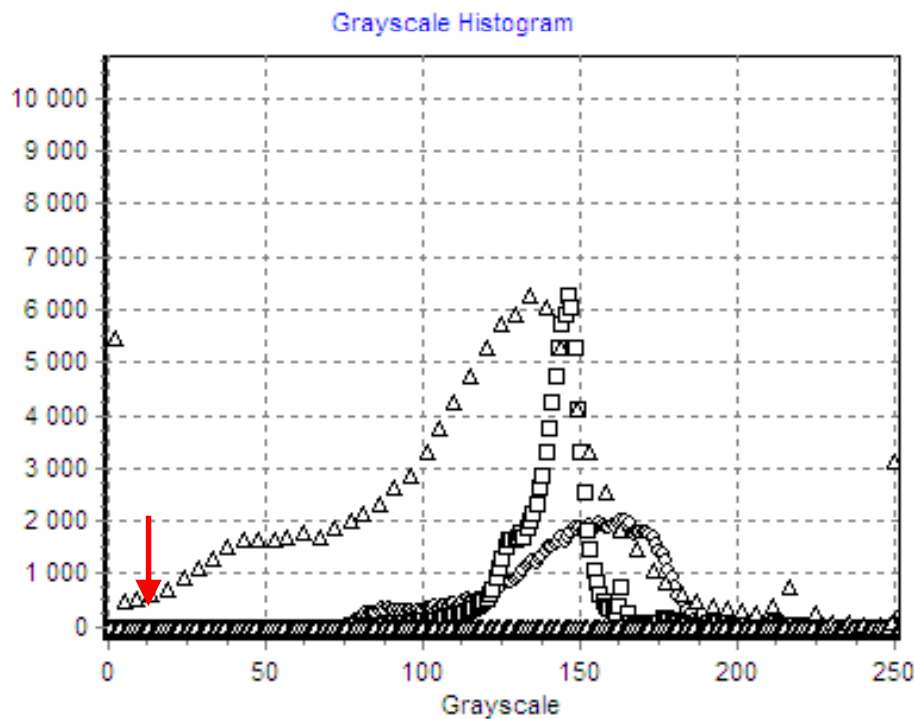
- circle - initial image,
- rectangle - after filtering,
- triangle - after filtering and contrast.



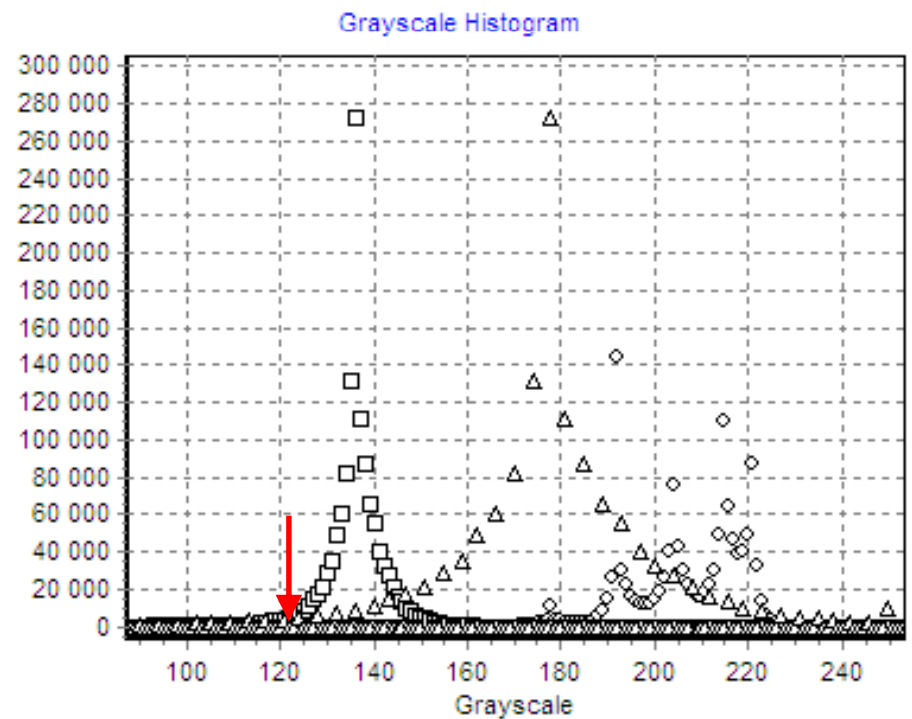
# Brightness histogram

- circle - initial image,
- rectangle - after filtering,
- triangle - after filtering and contrast.

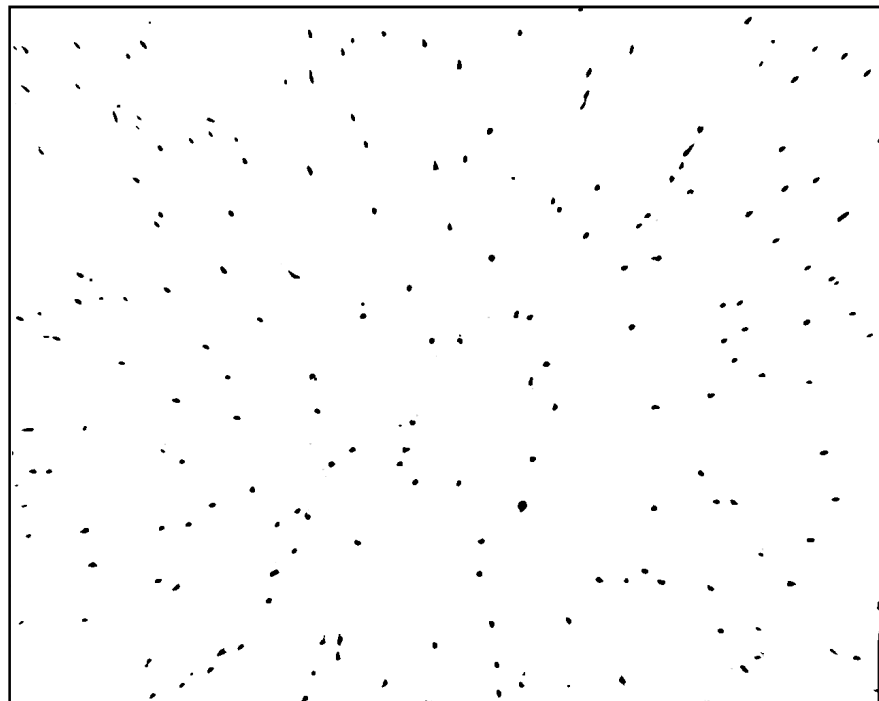
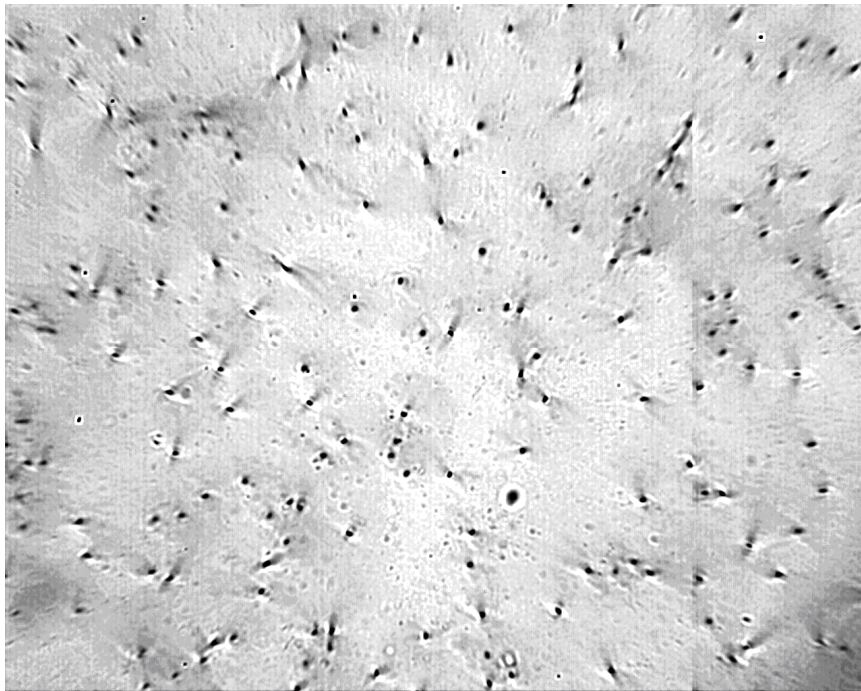
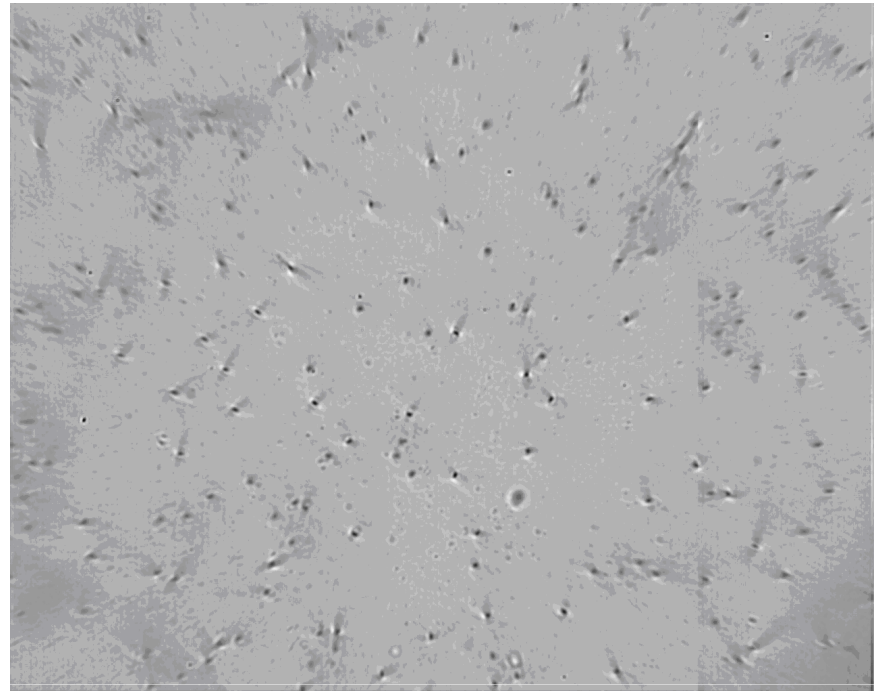
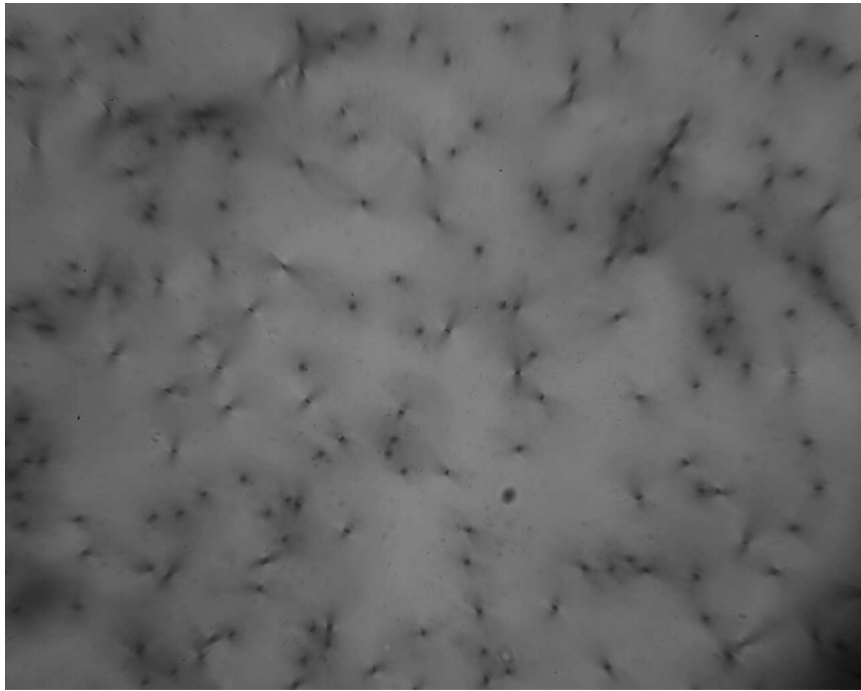
## EMU-15



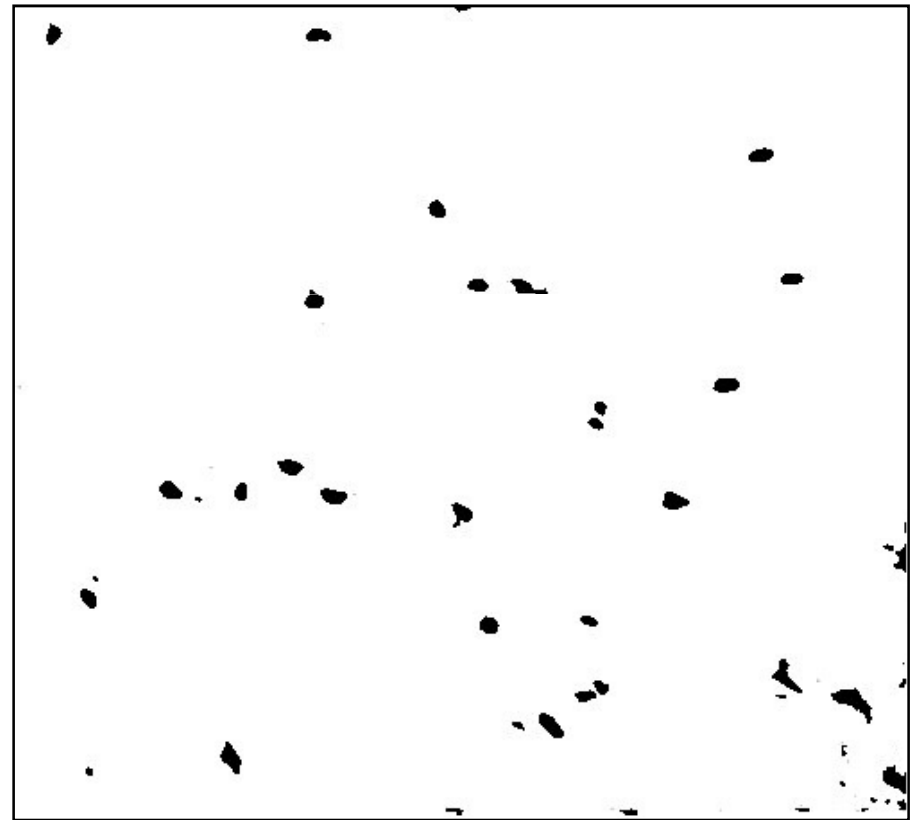
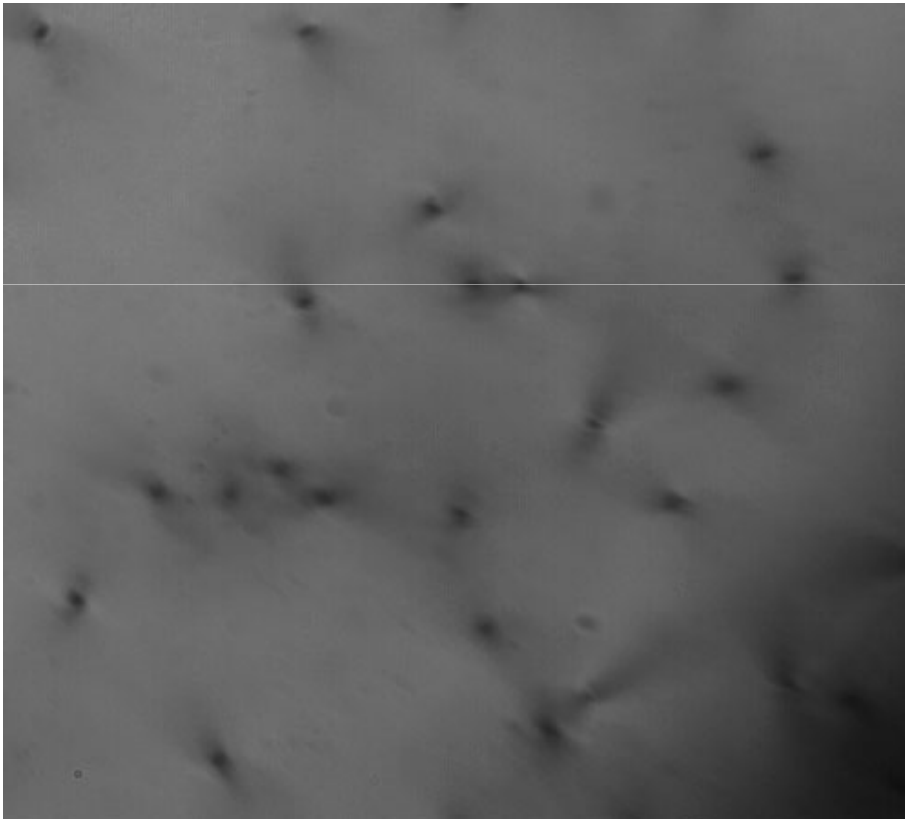
## OPERA







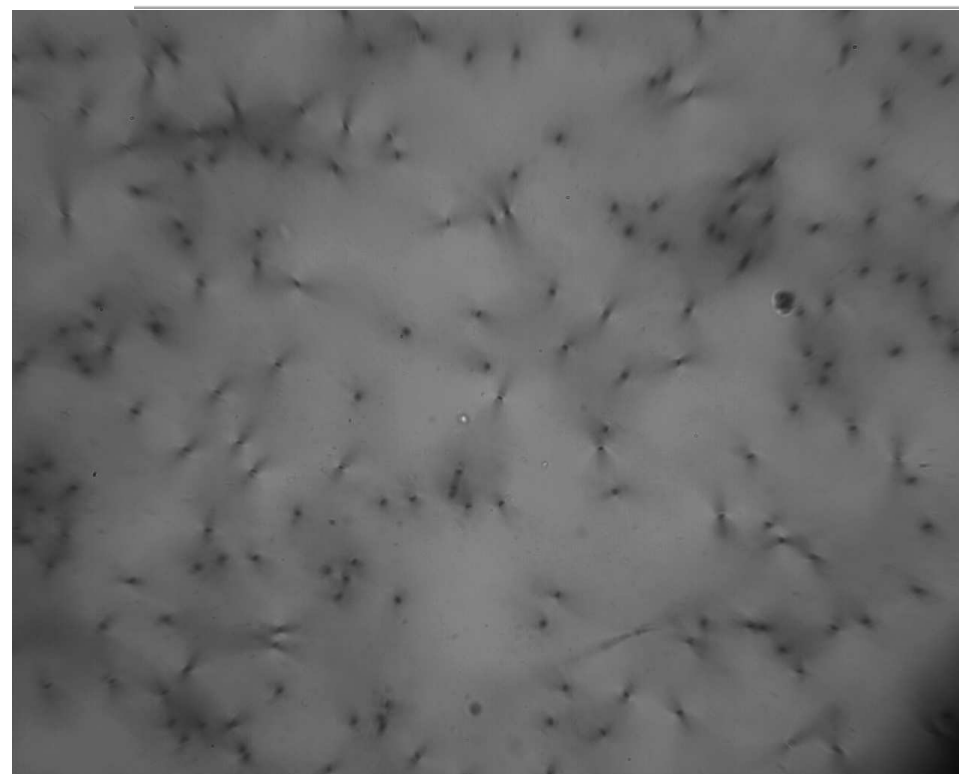
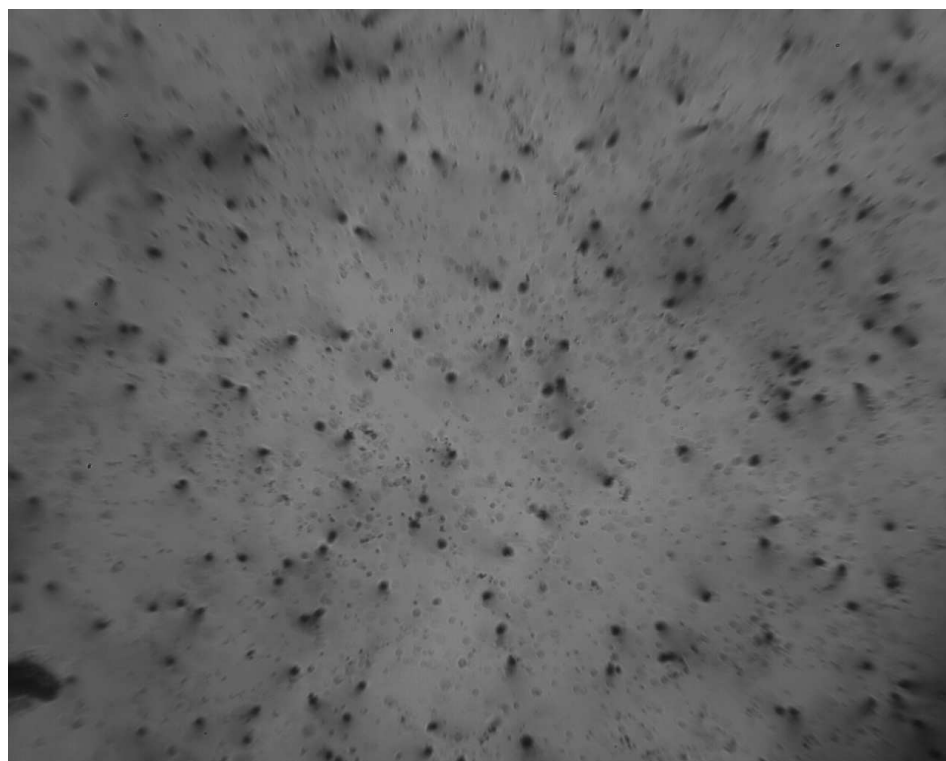
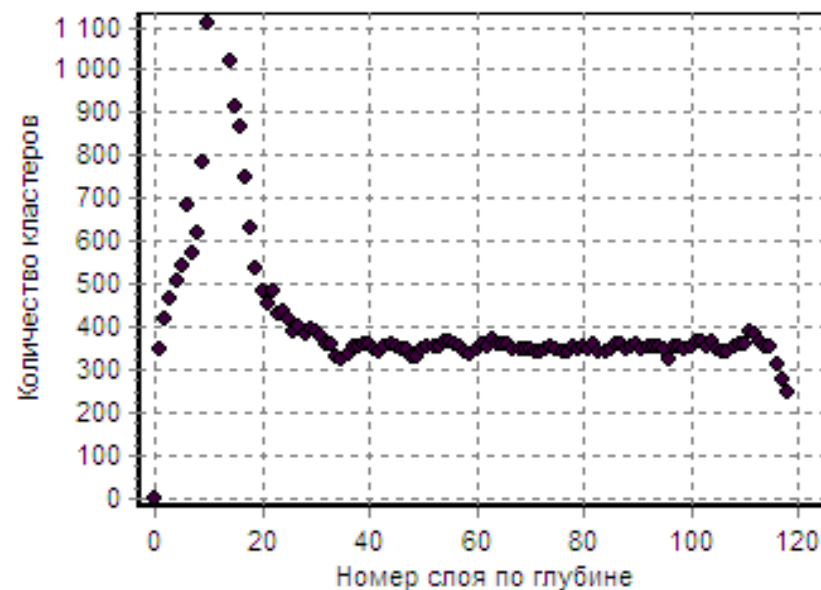
# Clusters extraction from dark background



Large thickness emulsion (600 mcm)

(different image properties is a result of irregular development)

=> Variable threshold and track characteristics

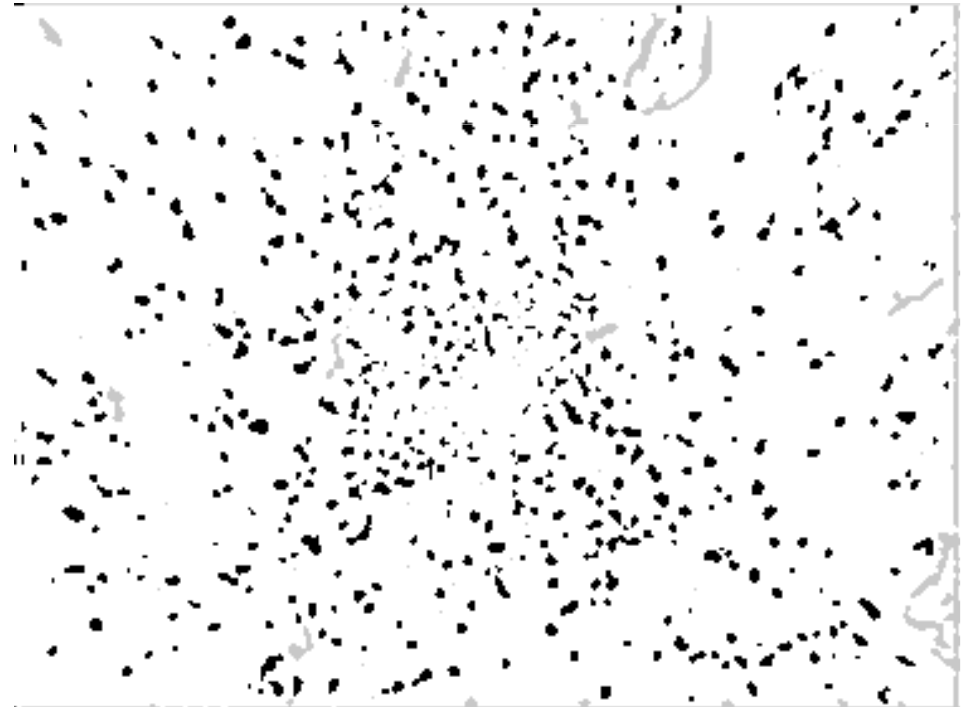
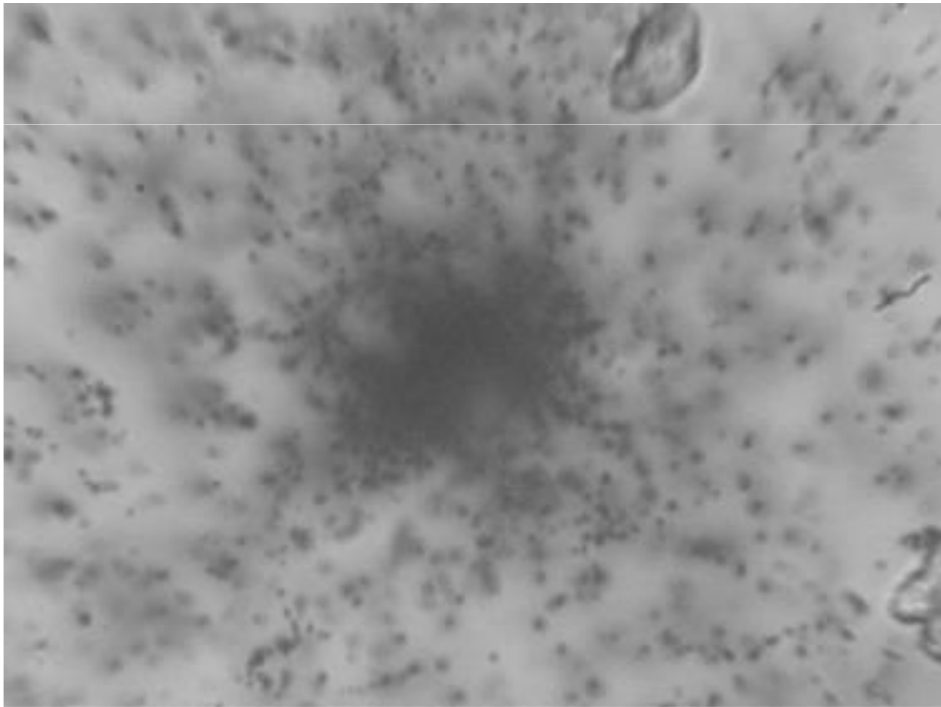


# The result of clustering.

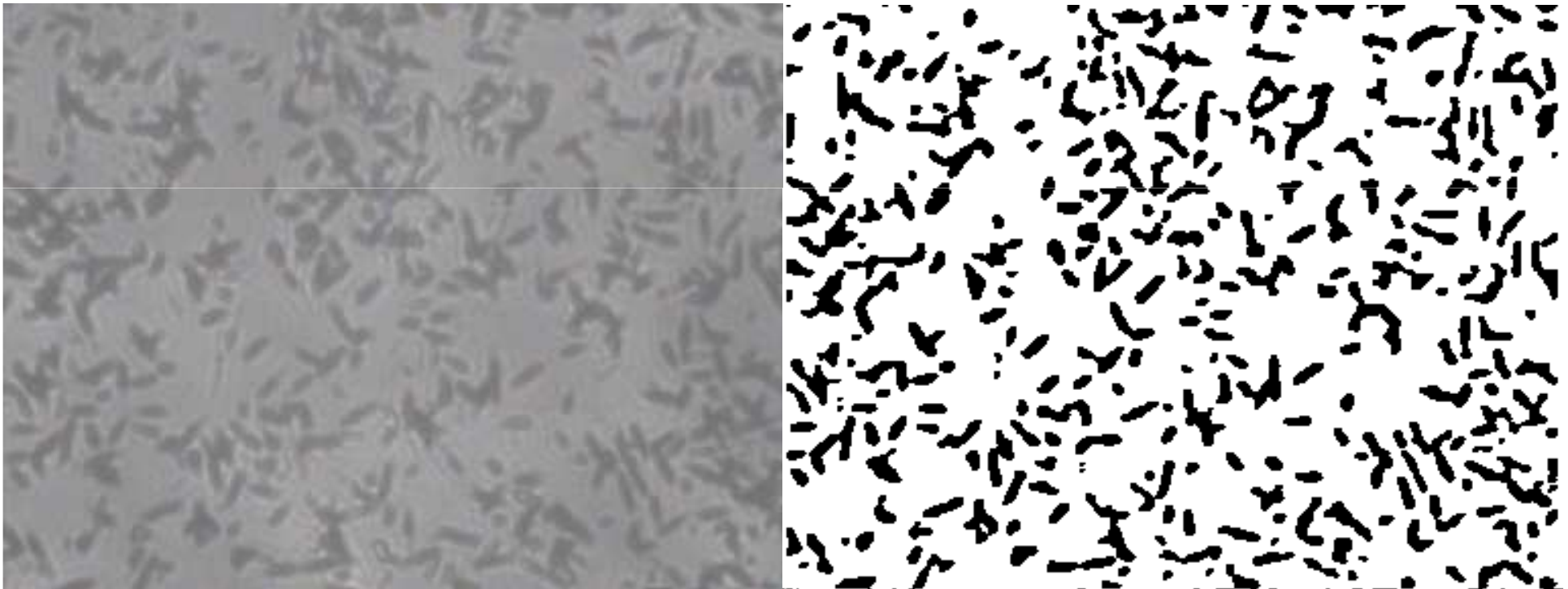
(projection of all CM of clusters from 120 layers)



## Clustering of EMU-15 image

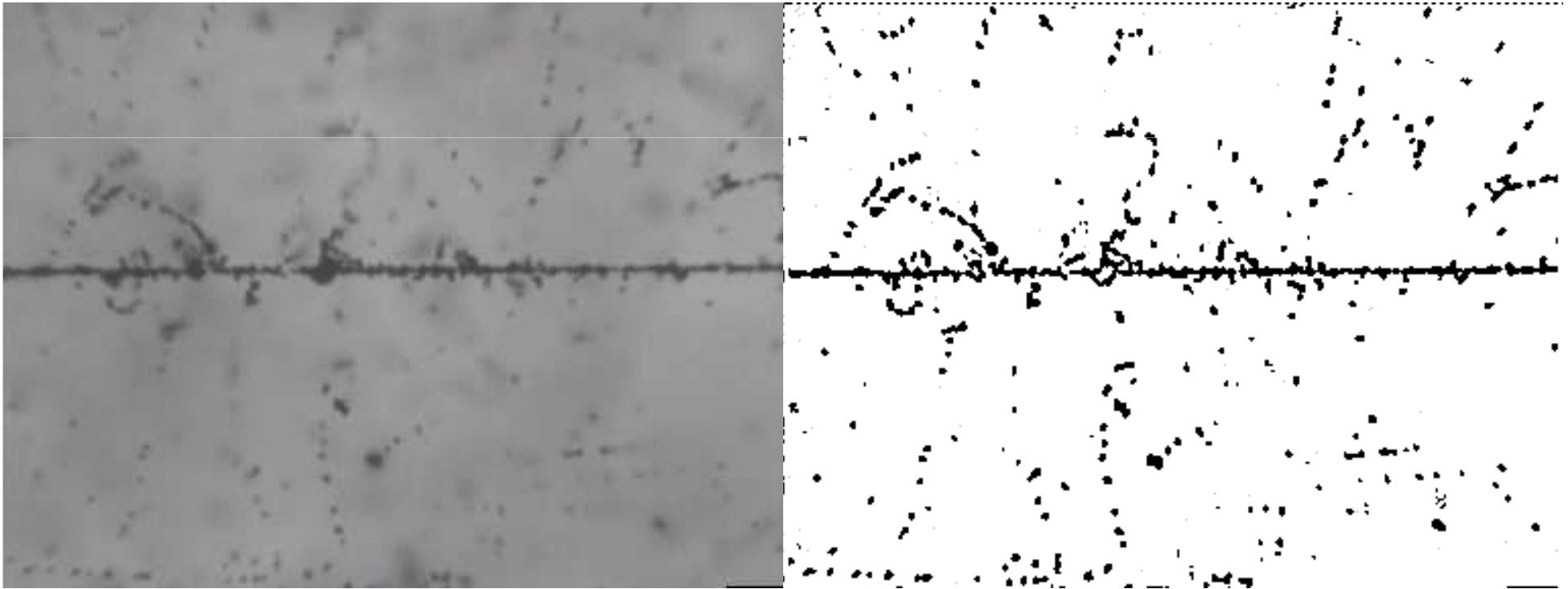


## Clustering of Energy+Transmutation image

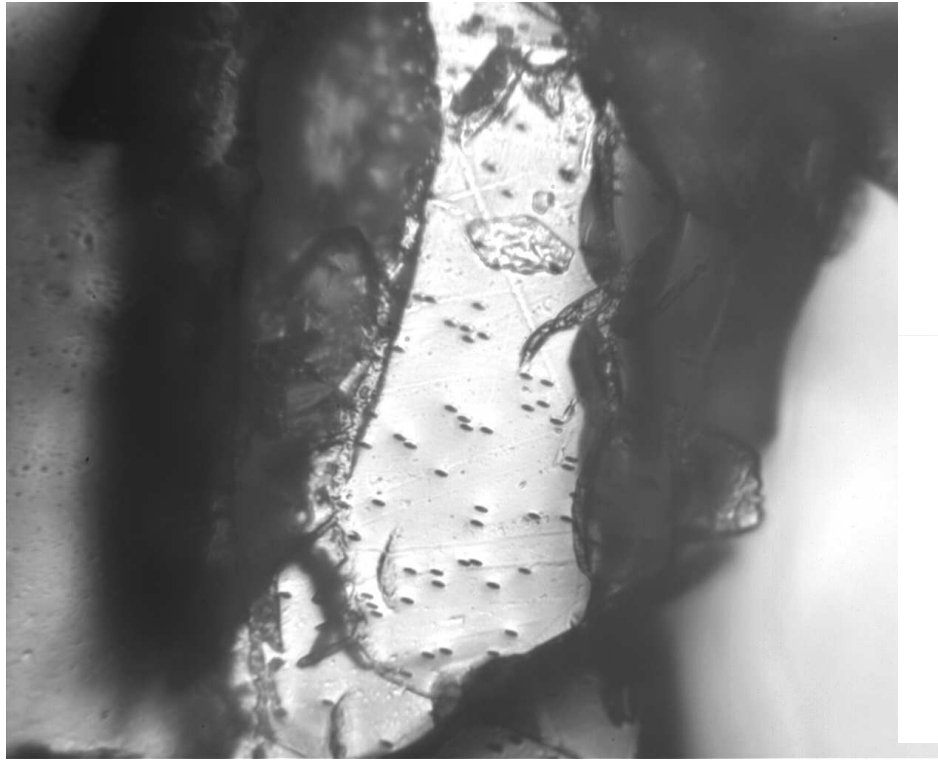


(many tracks intersect each other)

# Clustering of longitudinal track in thick emulsion



# Particularity of olivine from meteorites processing



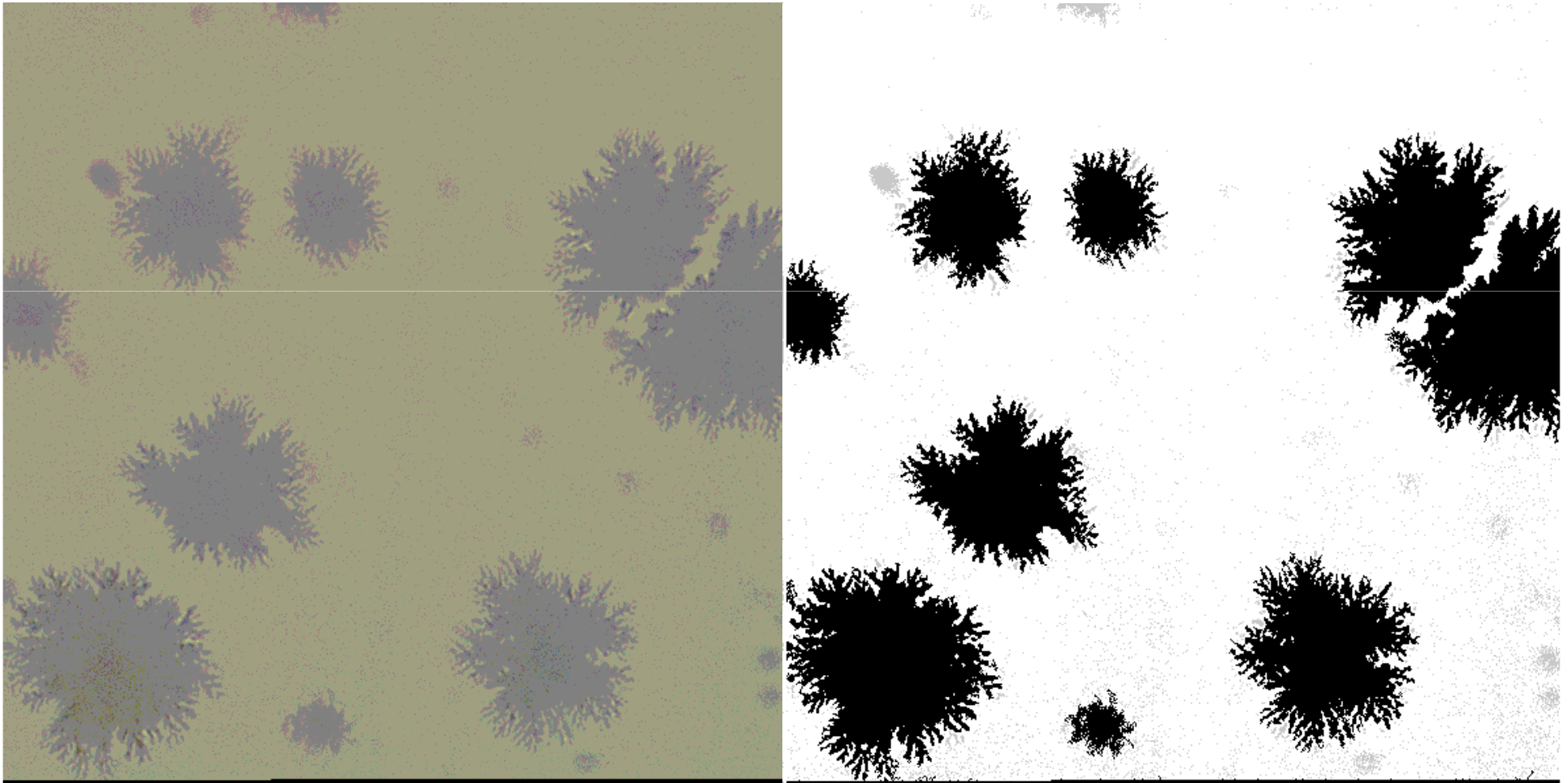
Extraction of processing area when there is large noise area



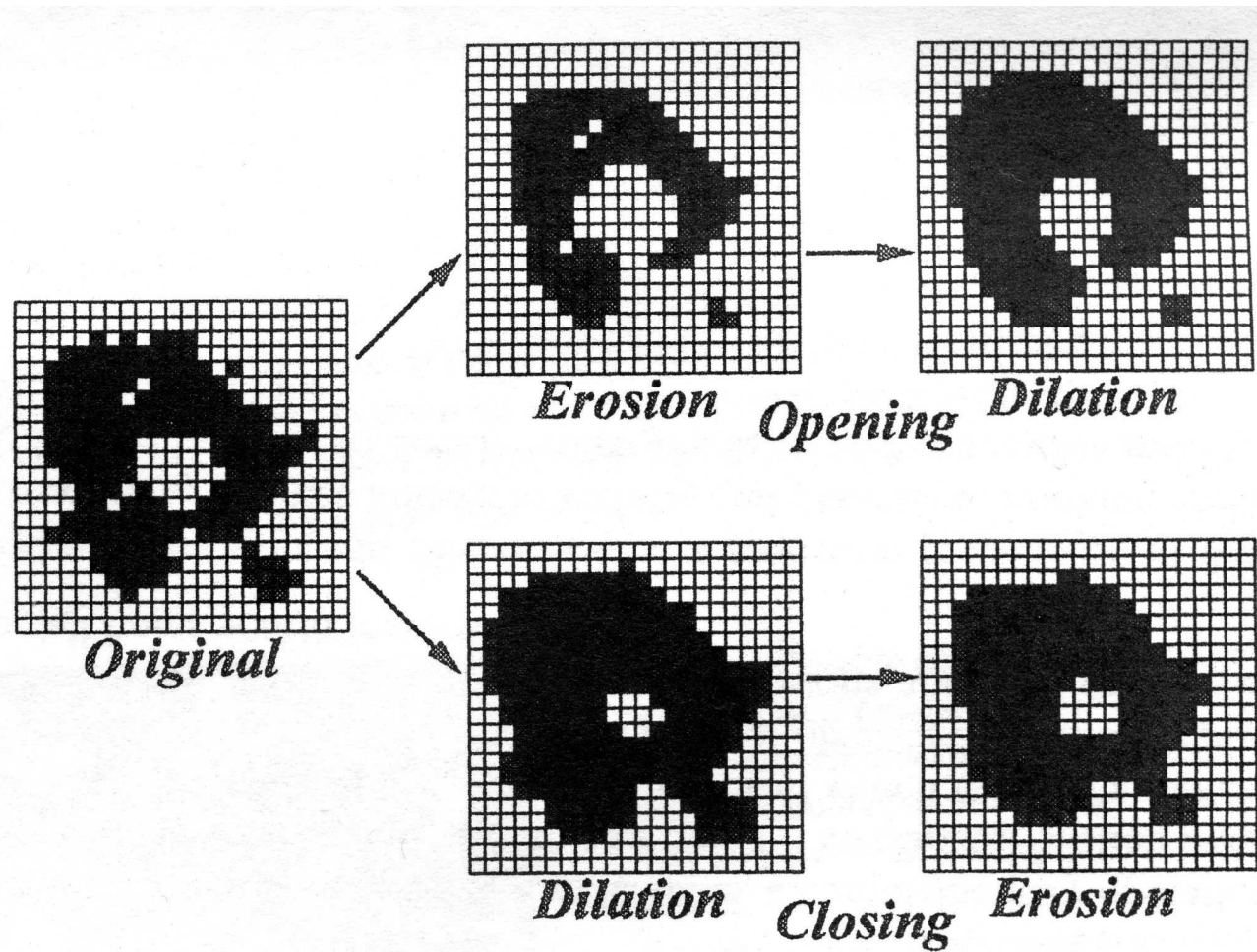
# Extraction of complicated track (heavy nucleus trace in olivine)



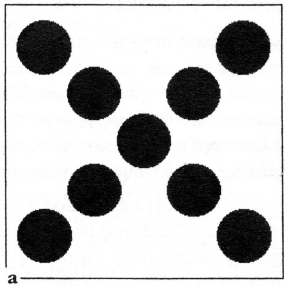
# Clustering of tracks having very complicated form



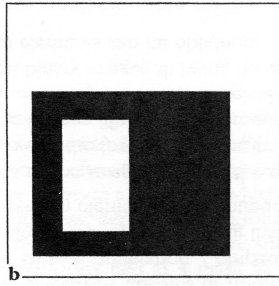
## Additional correction of clusters form



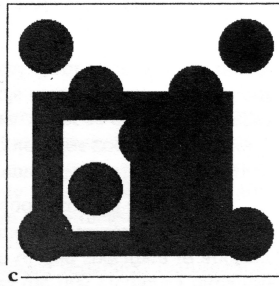
# Binary (logical) operations with images



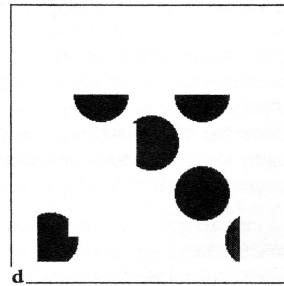
A



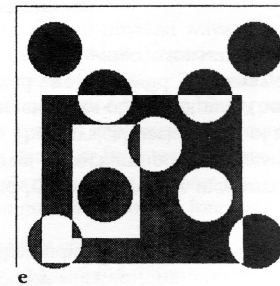
B



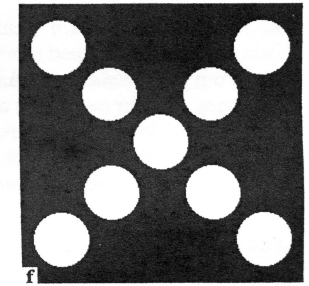
A Or B



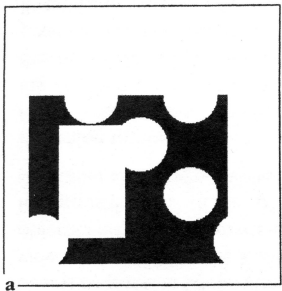
A And B



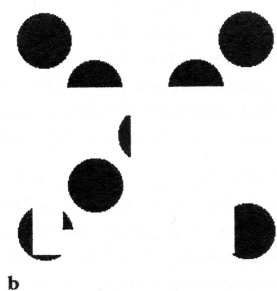
A exOr B



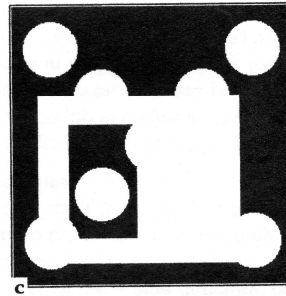
Not A



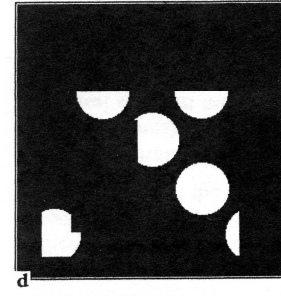
Not A) And B



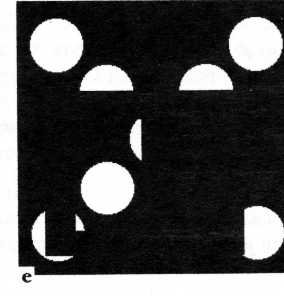
A And (Not B)



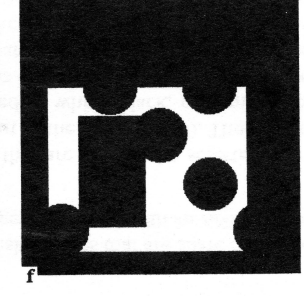
(Not A) And (Not B)



Not(A And B)

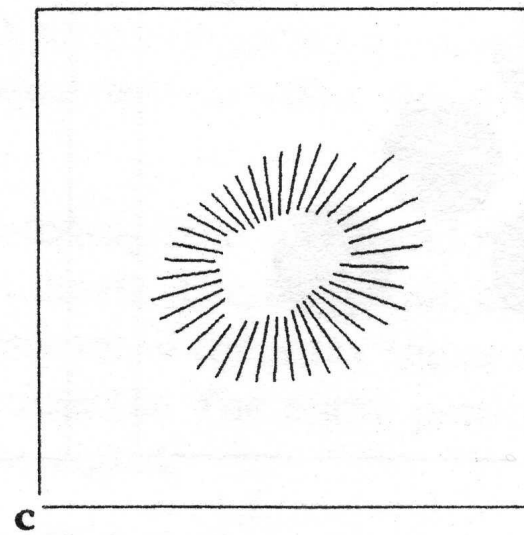
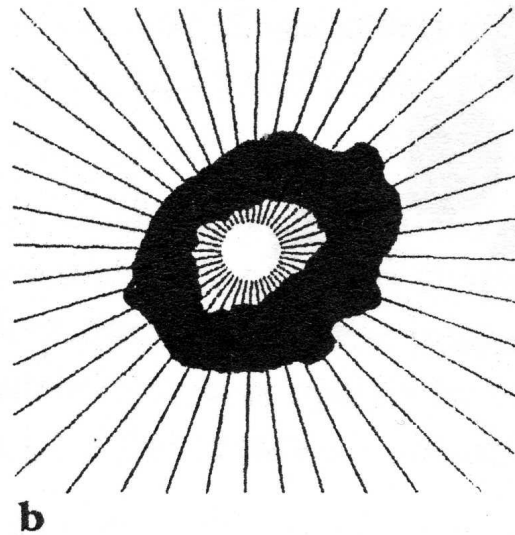
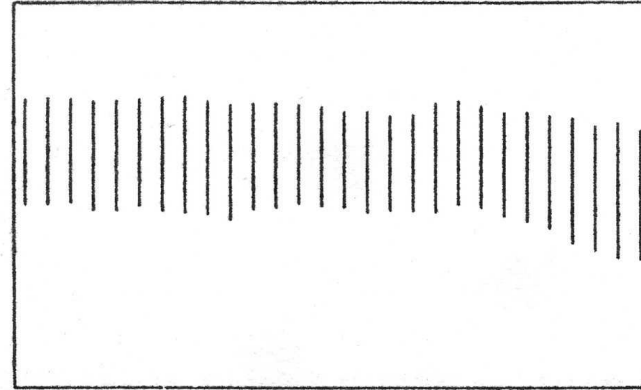
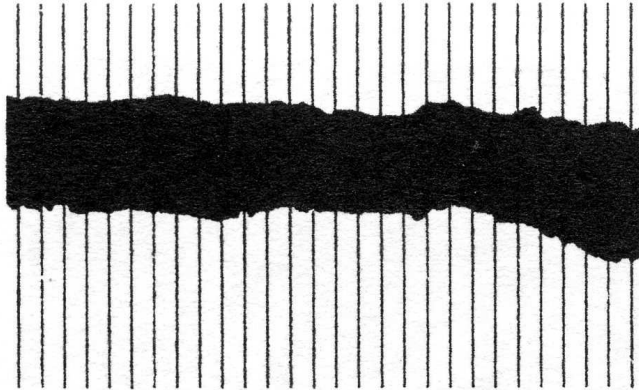


(Not A) Or B



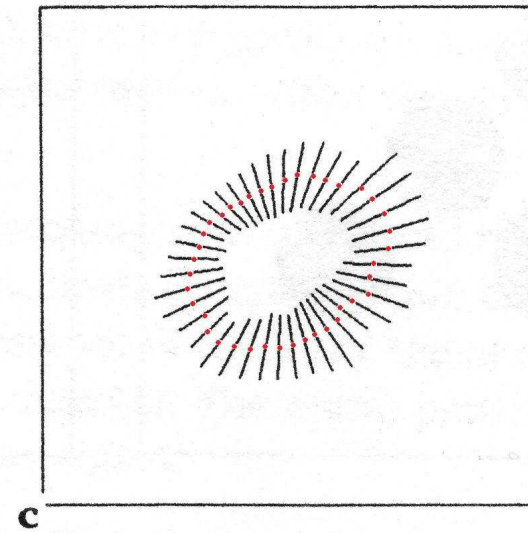
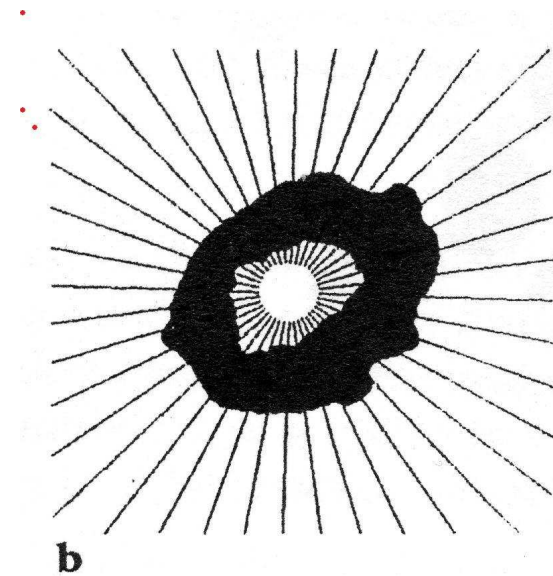
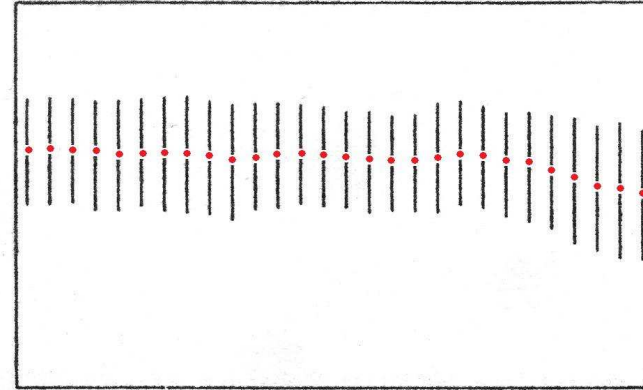
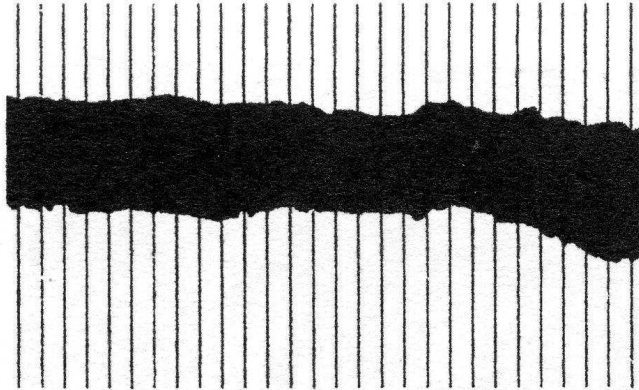
A And (Not B)

# Use of logical operations to define geometrical characteristics



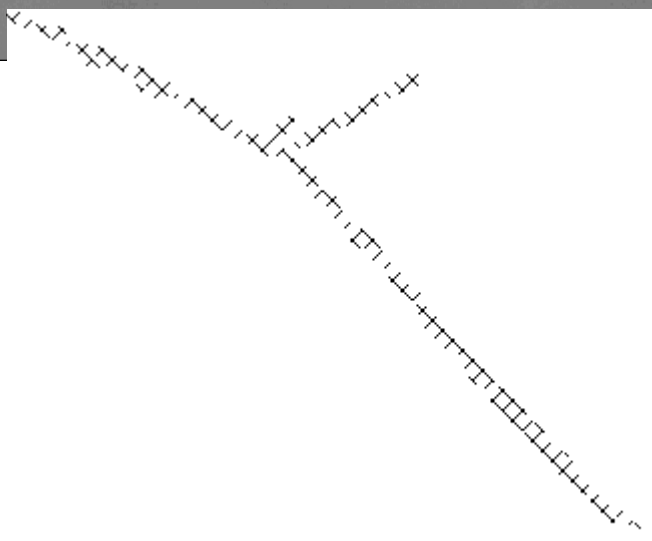
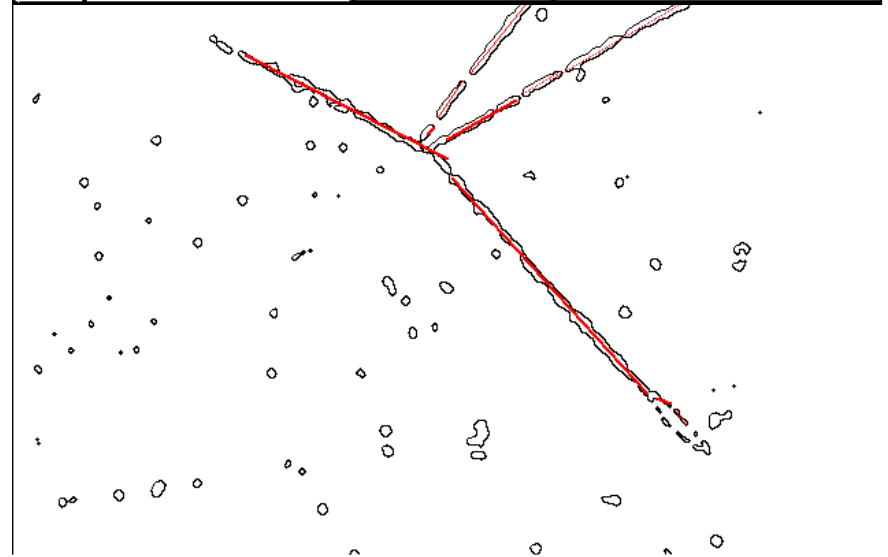
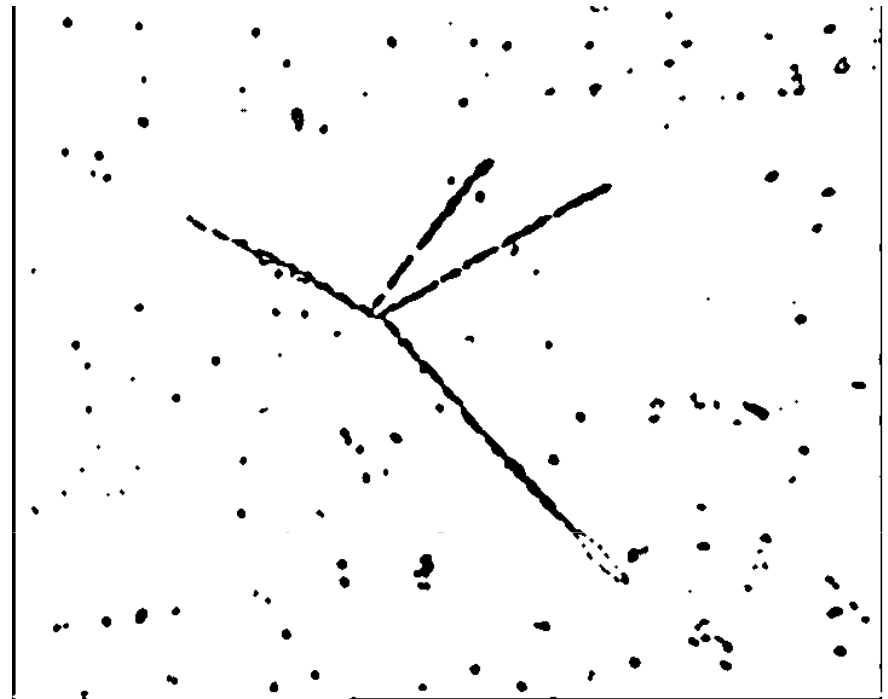
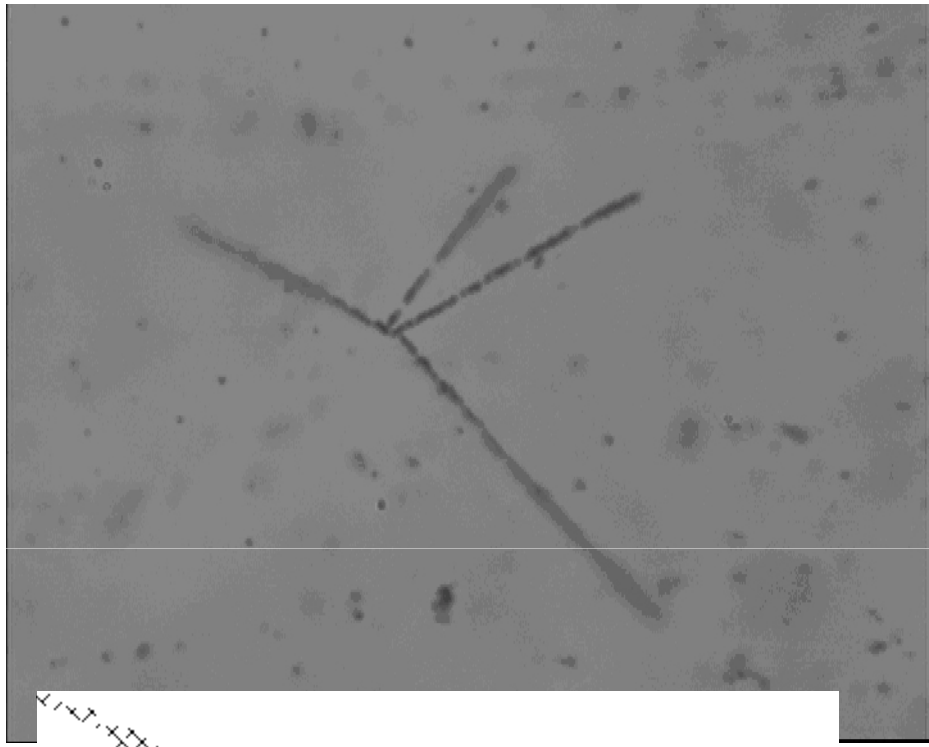
В более сложных случаях используются другие методы.

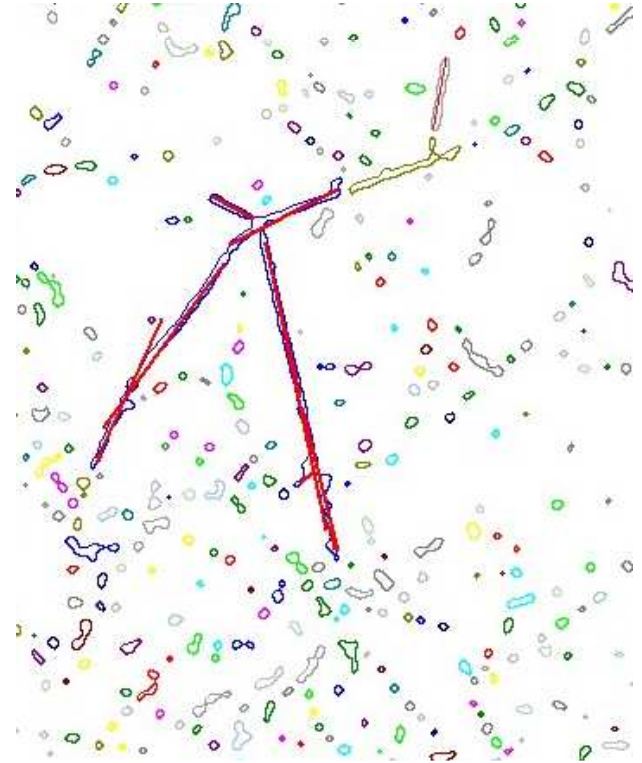
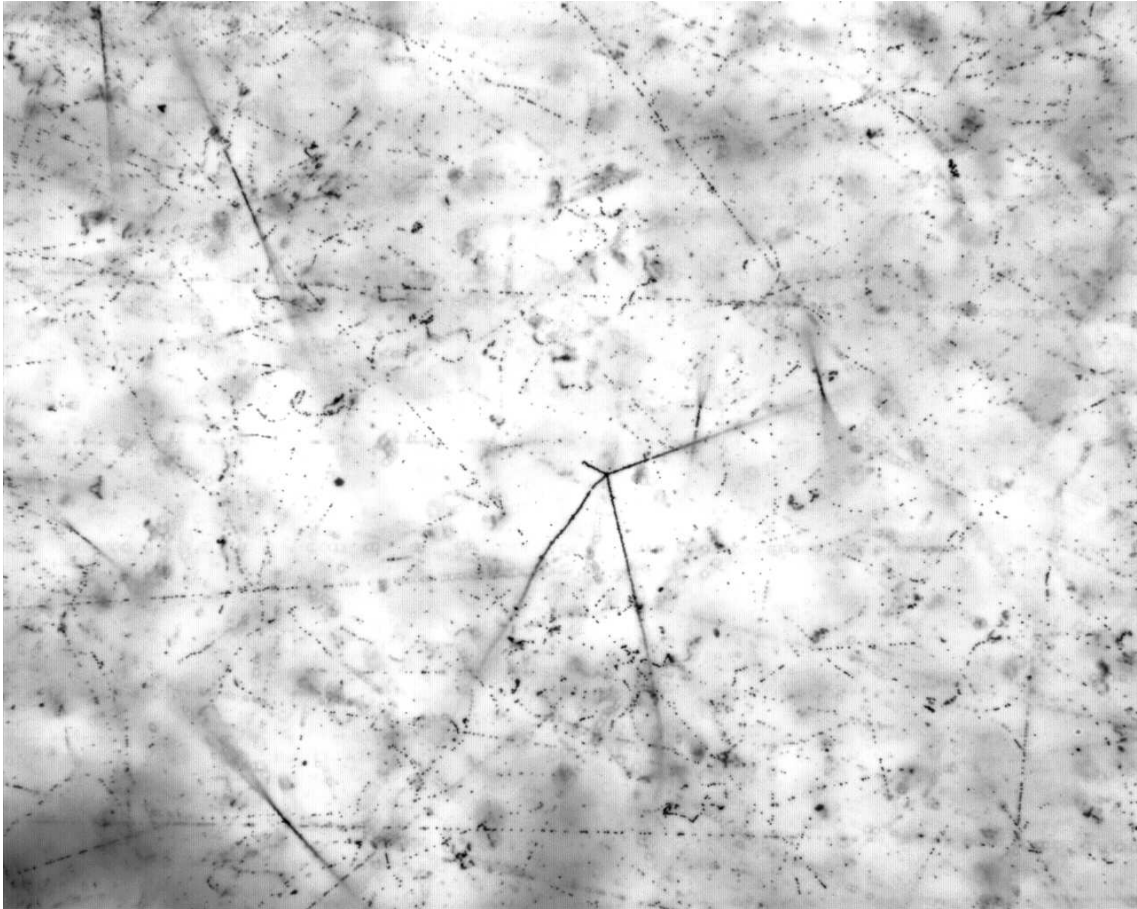
# Use of logical operations to define geometrical characteristics (axis, skeleton ...)



В более сложных случаях используются другие методы.

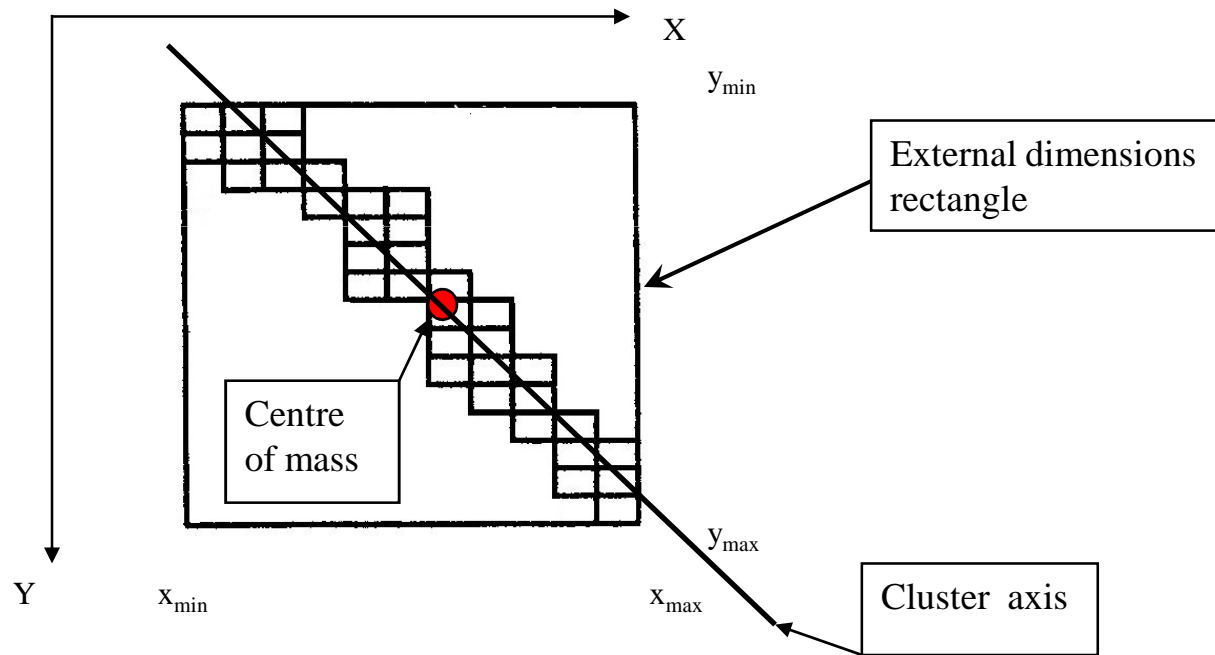
# The Extraction of a nuclear star rays using auxiliary grid





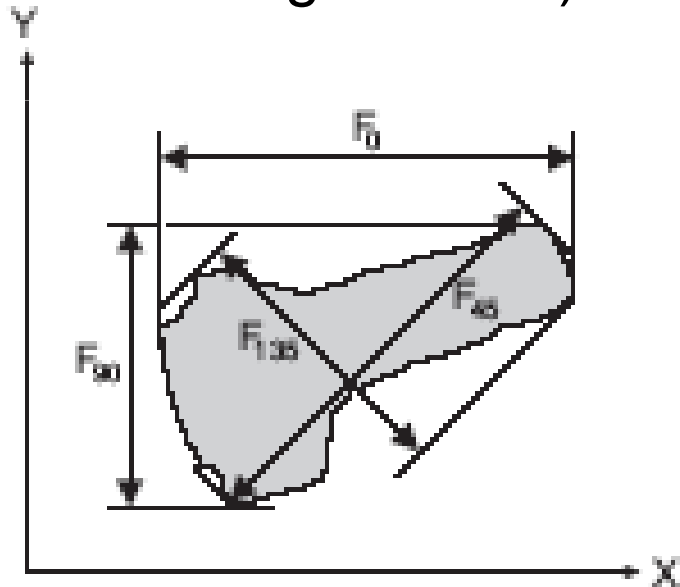


# Cluster's axis, length and width

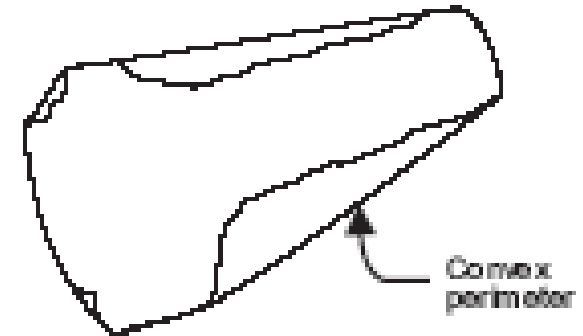
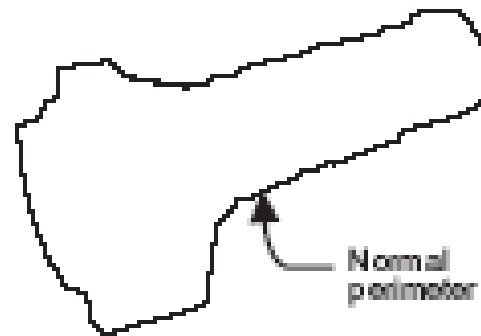


# Other geometrical characteristics

Ferret's diameters  
(distance between  
two tangent lines)



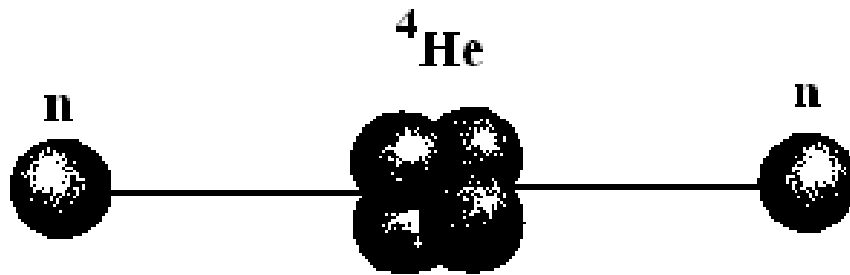
Usual and convex perimeter



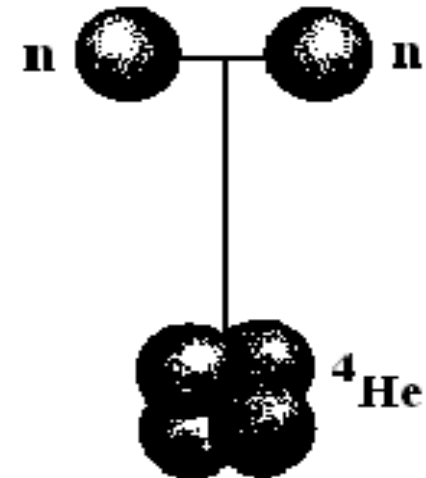
## The investigation of ${}^6\text{He}$ interaction in emulsion

It is possible the existence of two configurations of additional neutrons

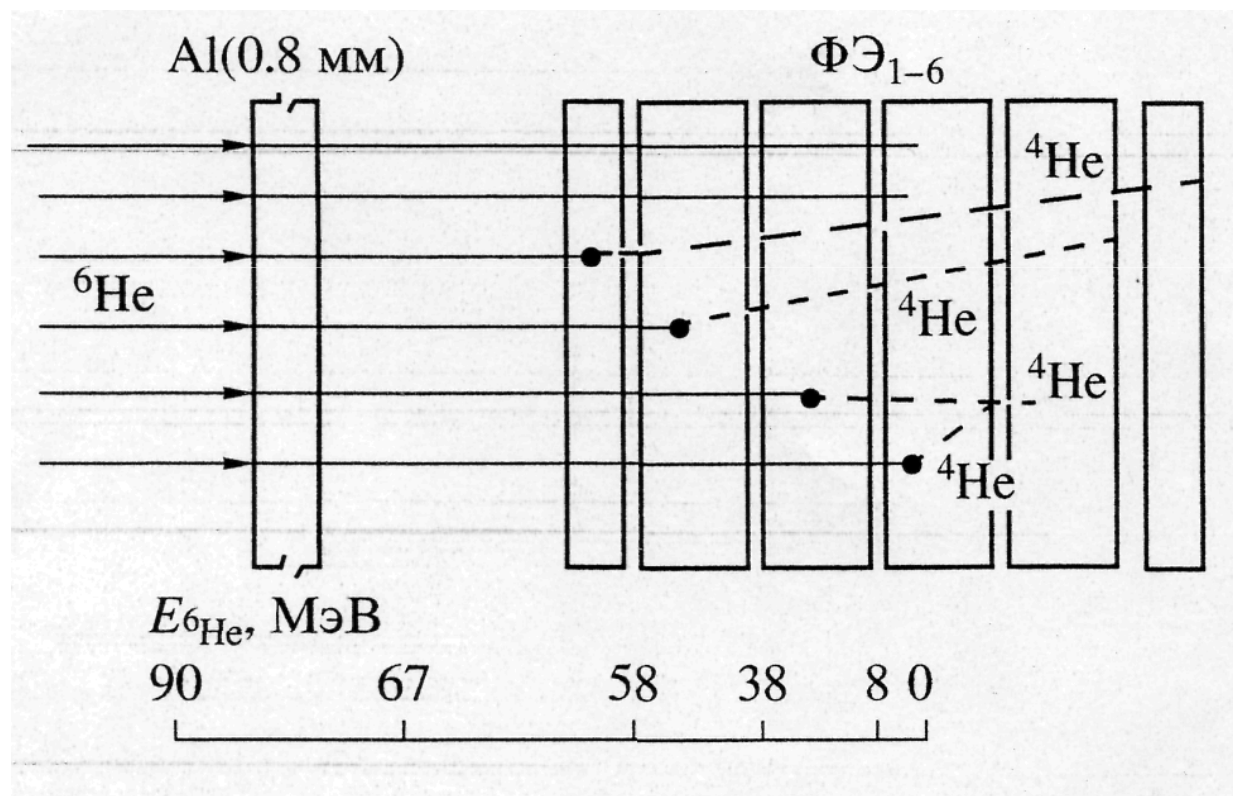
cigar-like



dineutron



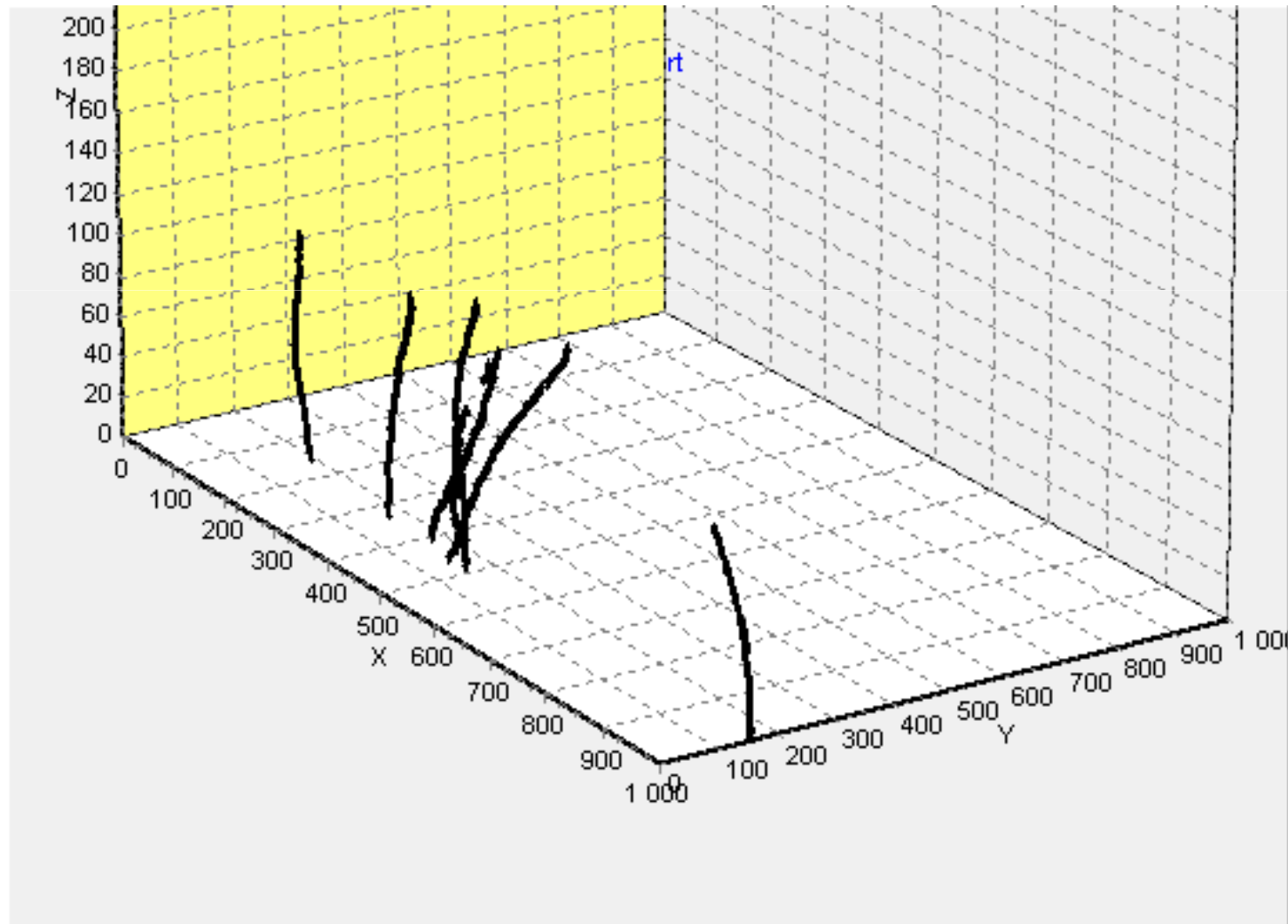
# The schema of experiment



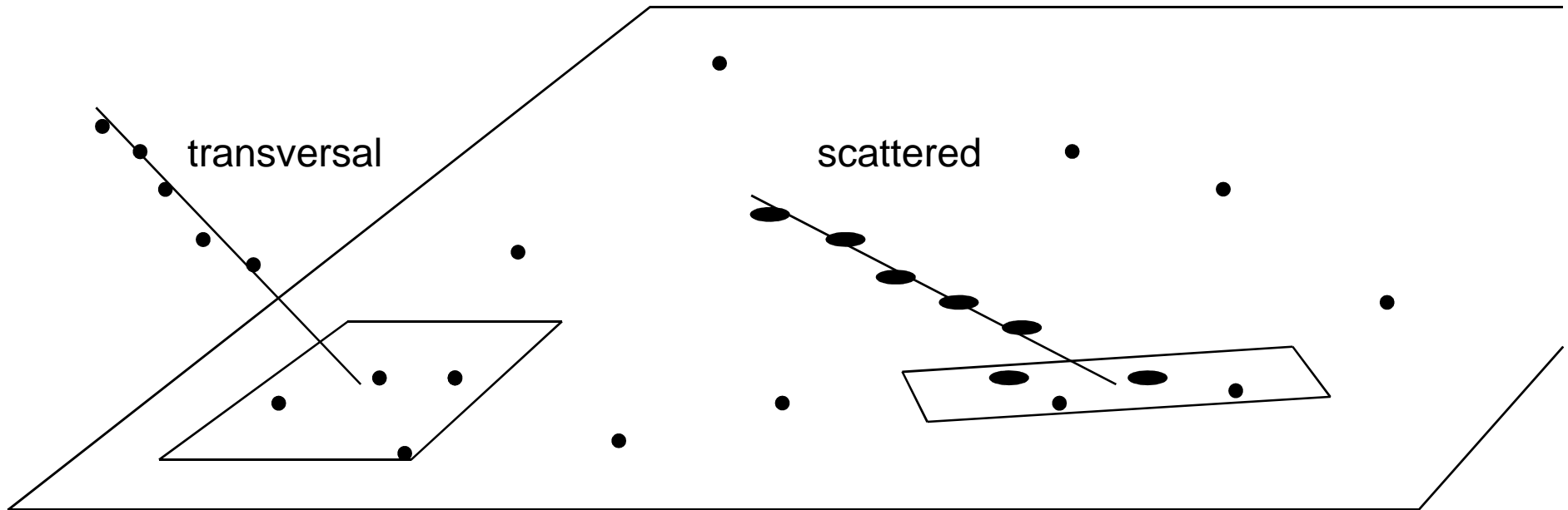
# Analysis of interactions

- Clusters alignment in three dimension space (tracking)
- Search of vertexes
- Geometrical reconstruction of interaction
- Reconstruction of kinematics

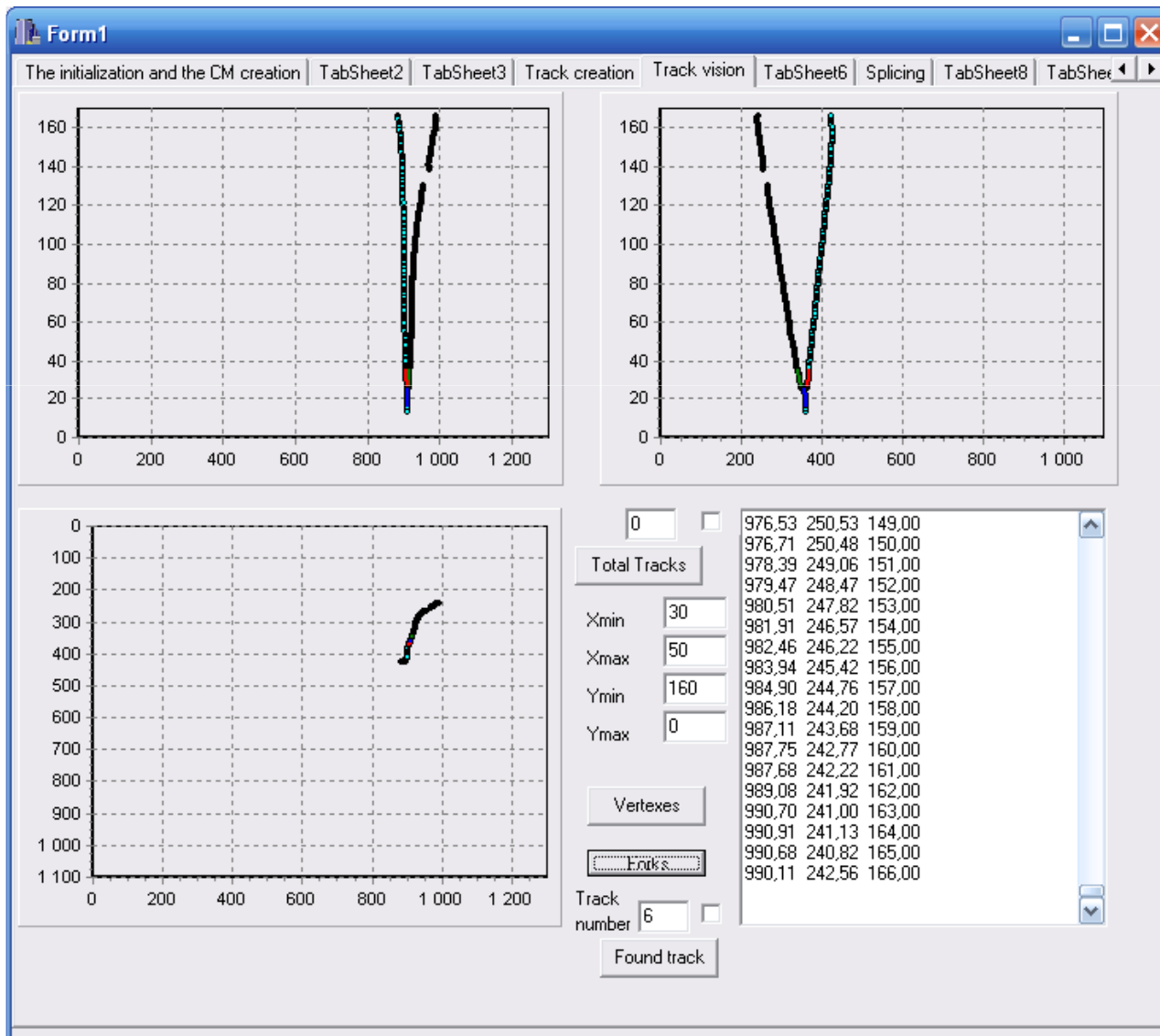
Most part of tracks are bent due to distortion



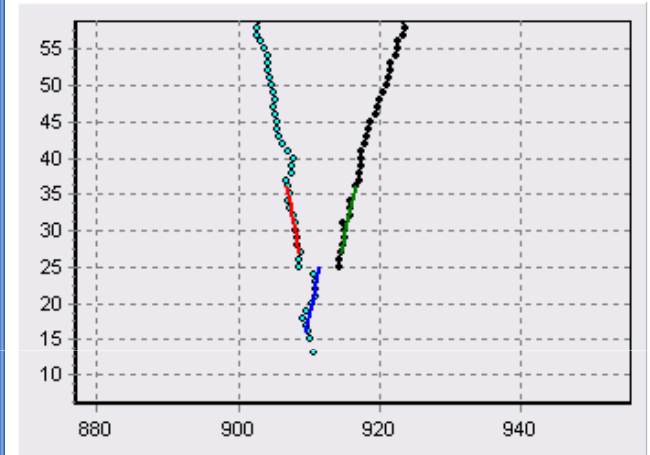
# Algorithm of track prolongation (sliding window)



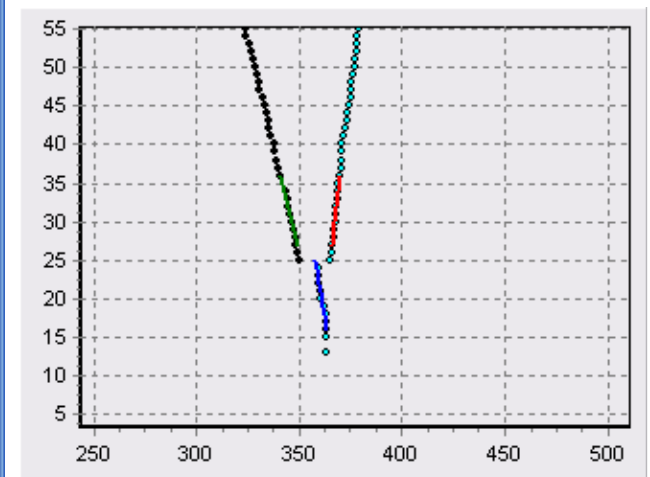
# One vertex of ${}^6\text{He}+p\Rightarrow\text{He}+(pnn)$



XZ

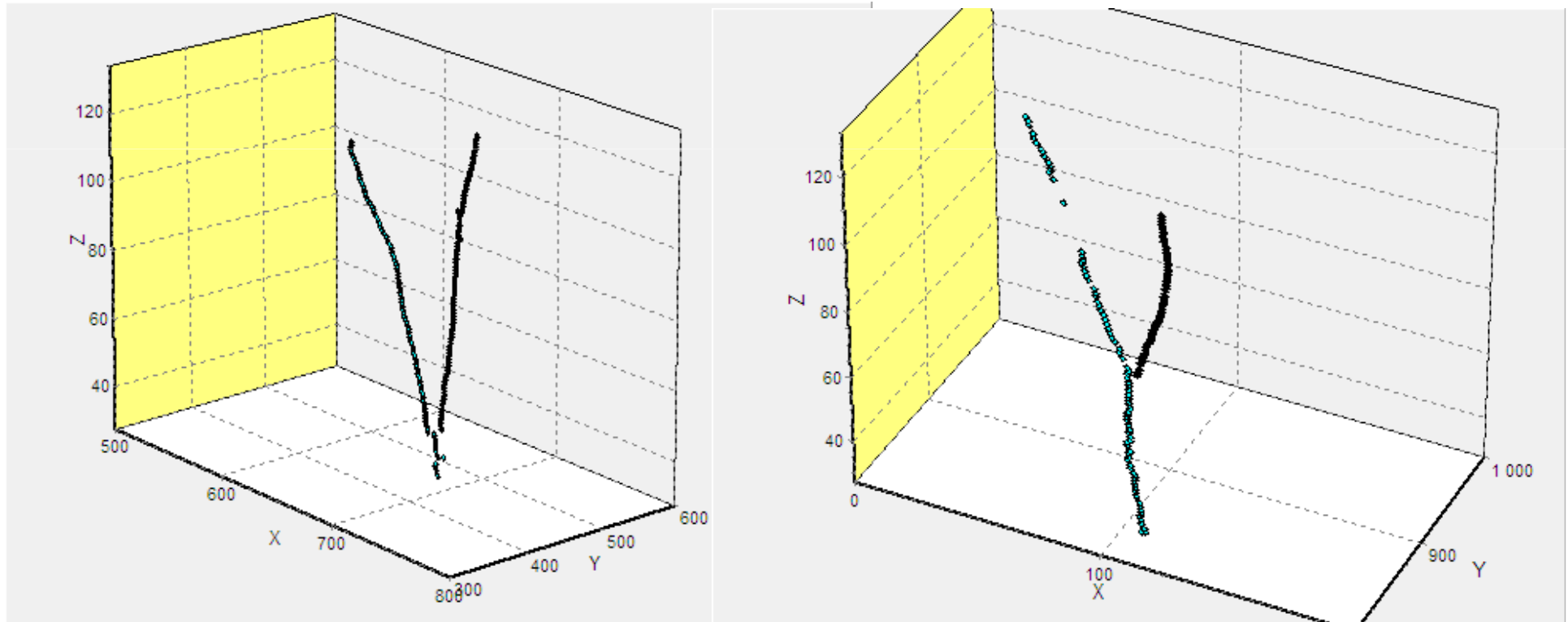


YZ

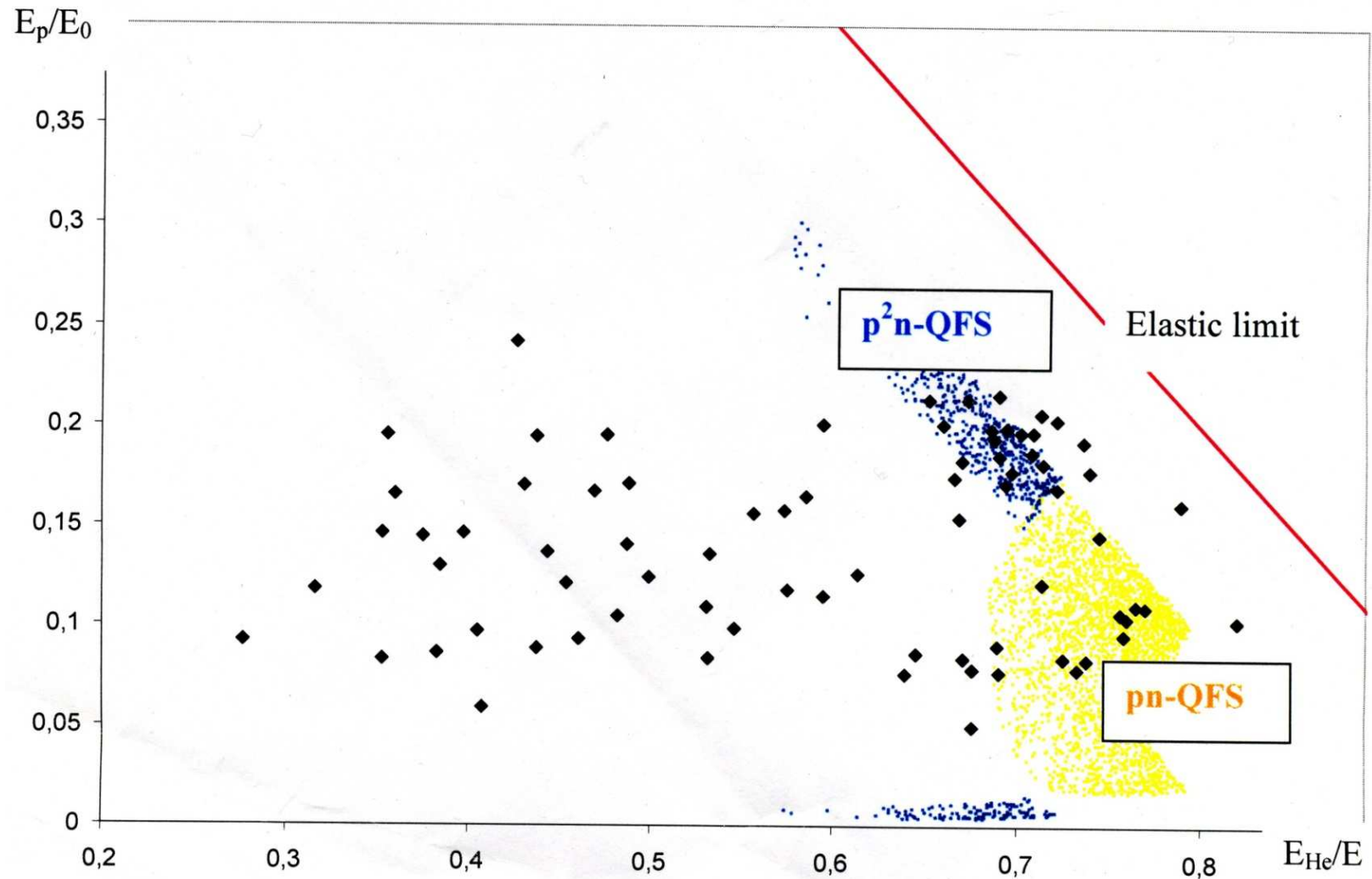




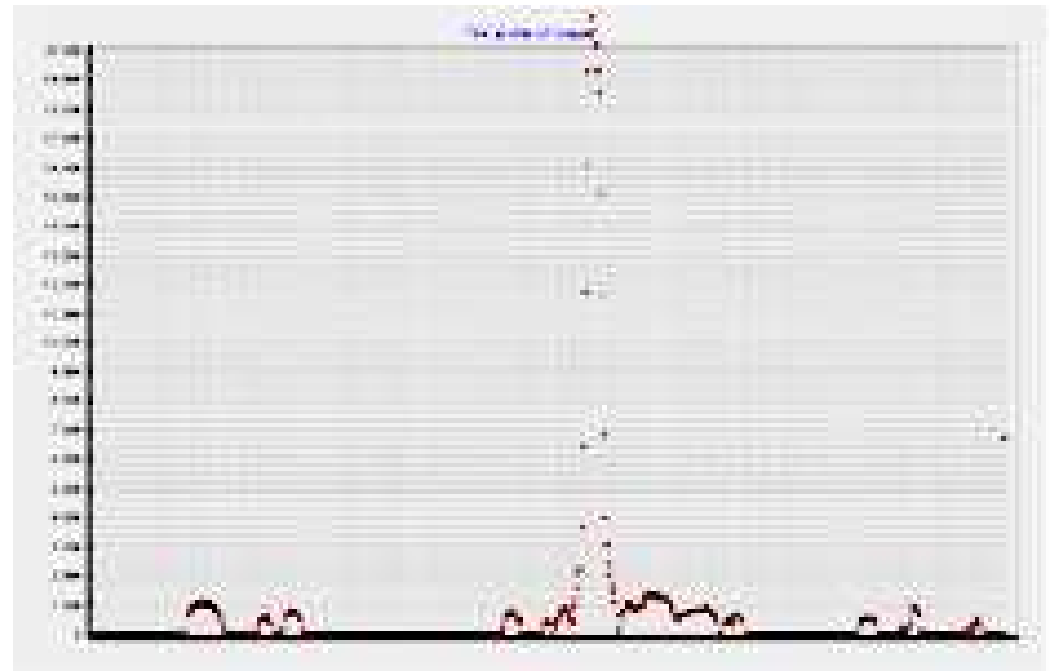
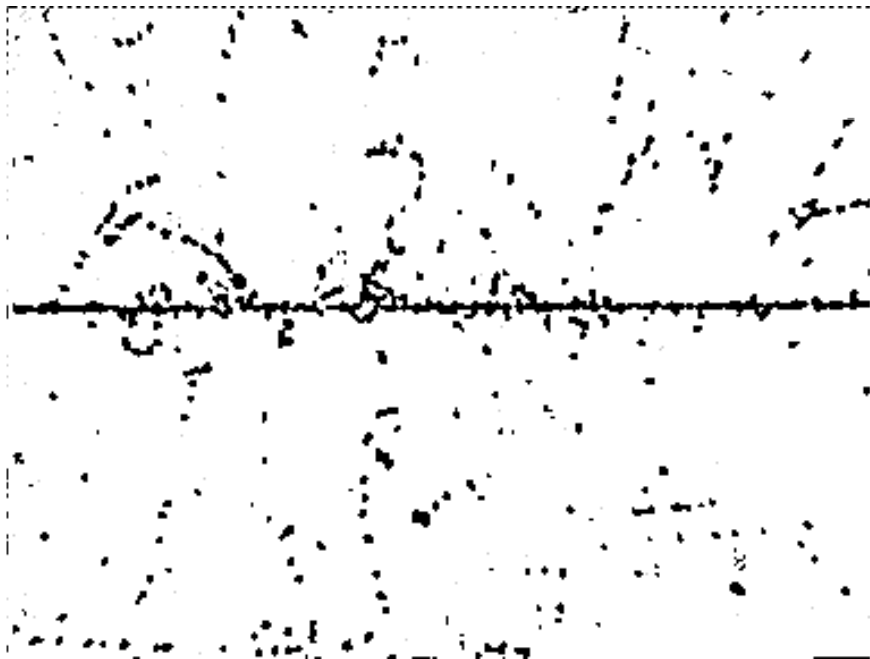
# Three dimensional representation



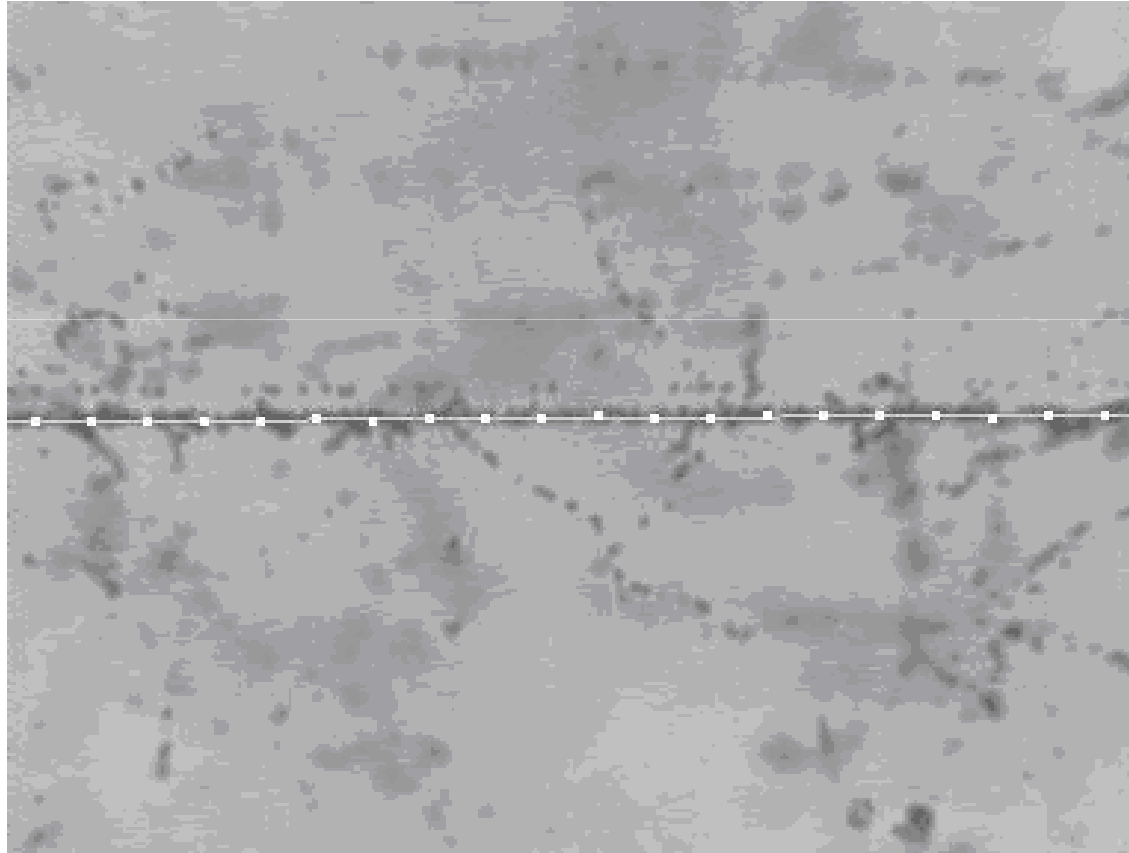
# Kinematical areas of reactions



# The method of automatic charge measurement (relativistic nuclei in thick emulsion)

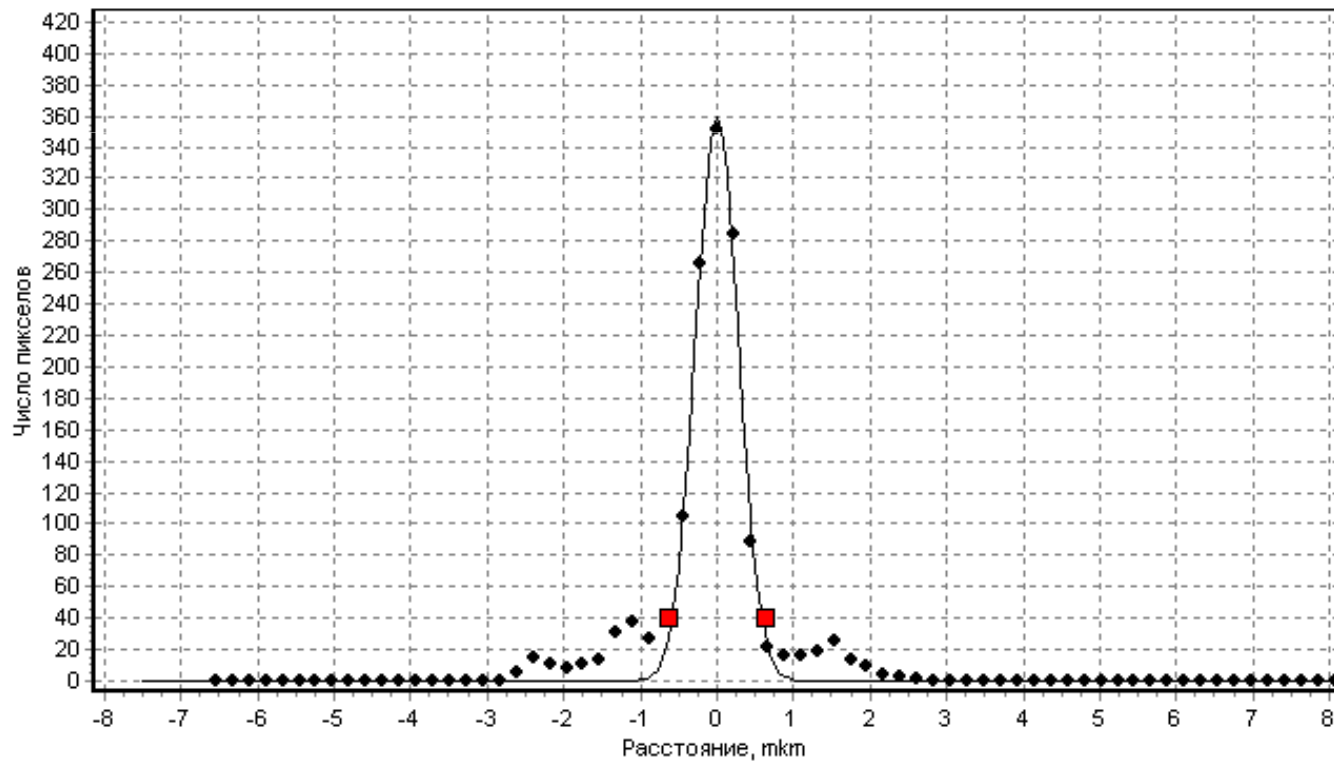


The axis creating  
(20 auxiliary vertical strips)



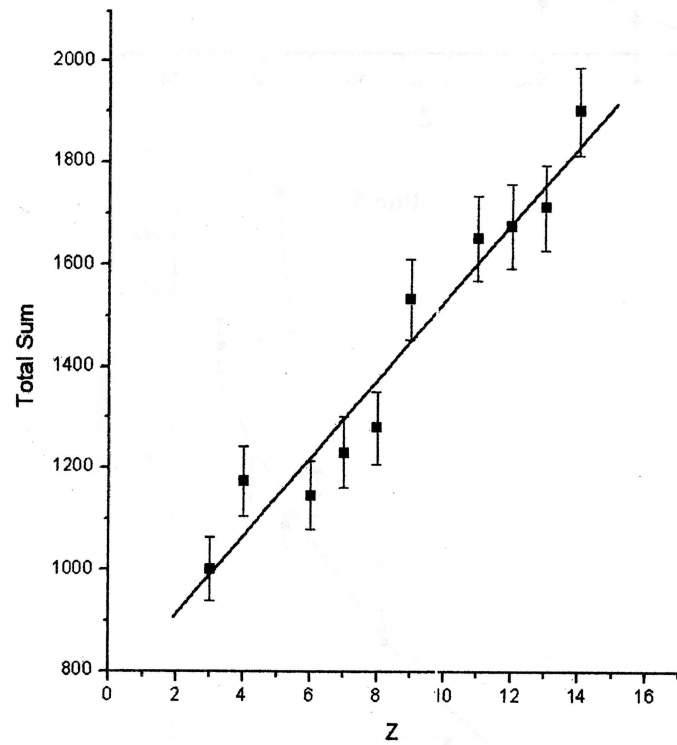
# Dependence of the pixels number on their distance to axis

(only those clusters belong to track which are crossed by axis)

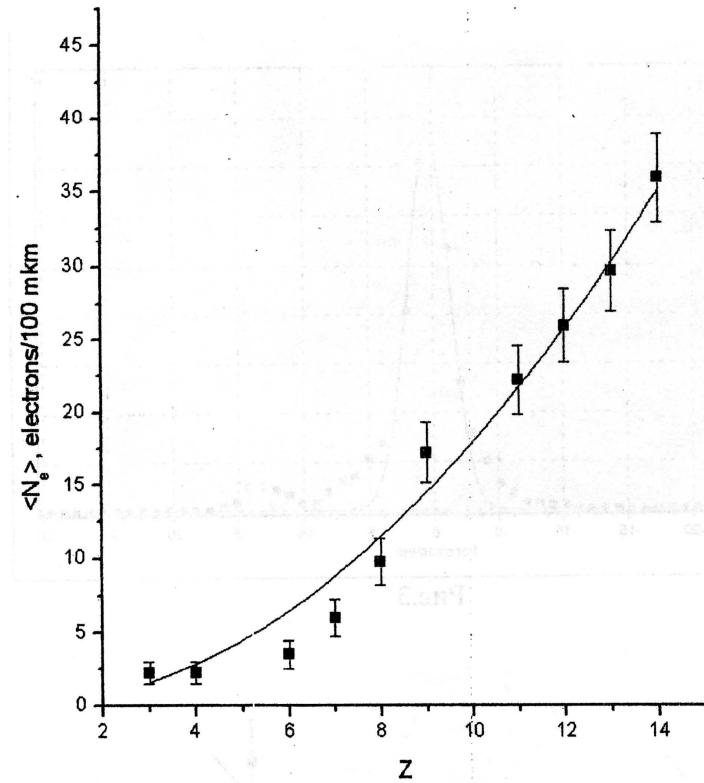


# Calibrating curves

*Total number of pixels of track*

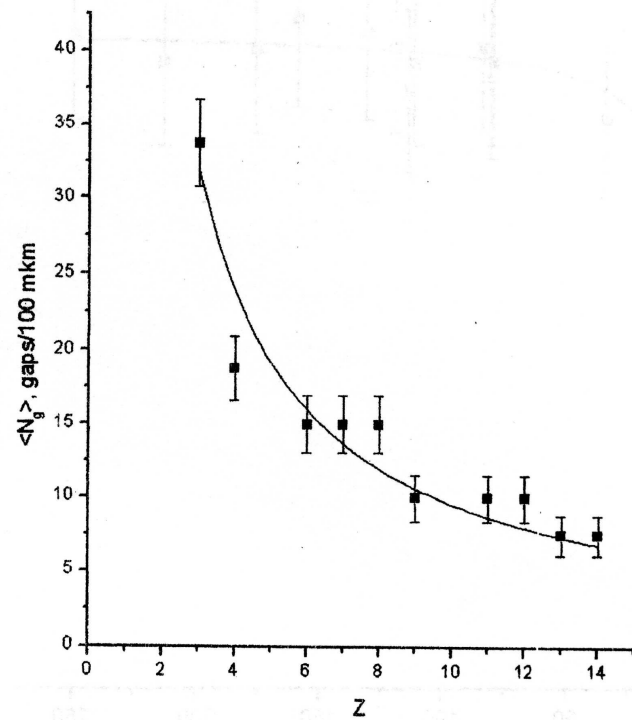


*Number of delta electrons*

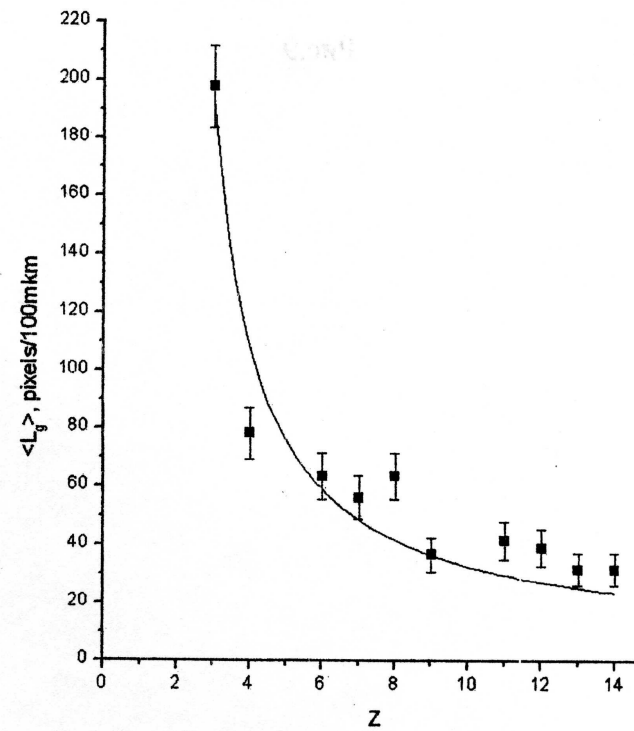


# Calibrating curves (light nuclei)

*Number of gaps between track clusters*



*Total length of gaps*



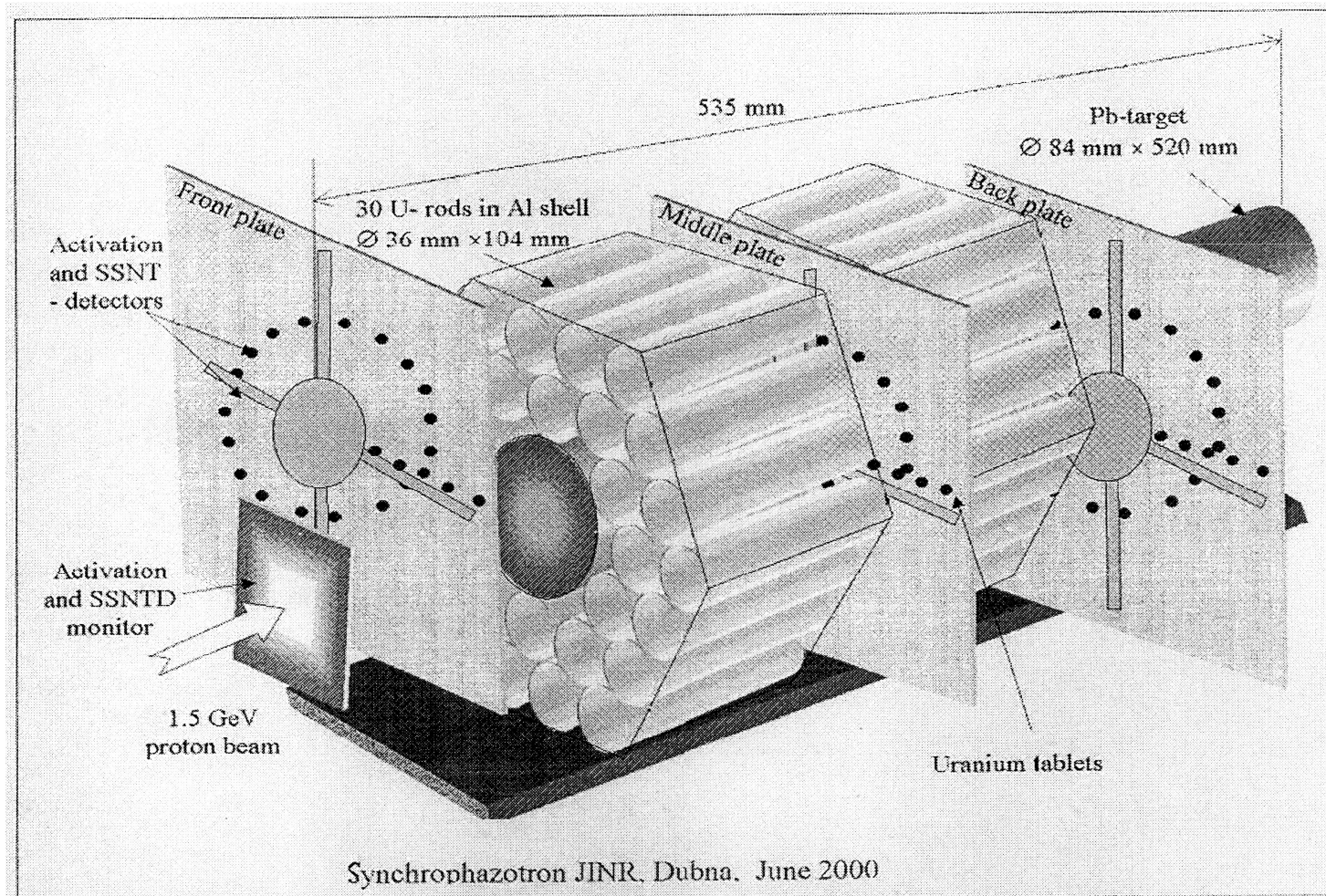
## Comparison of handle and automatic measurements of charges

<b>Event number</b>	<b>Handle</b>		<b>Automatic</b>	
	<b>projectile</b>	<b>fragment</b>	<b>projectile</b>	<b>fragment</b>
<b>174-43/153 90-39</b>	<b>12</b>	<b>10</b>	<b>12.2±0.4</b>	<b>10.2±0.3</b>
<b>174-43/174 80-48</b>		<b>5</b>		<b>4.9±0.3</b>
<b>174-43/180 67-52</b>	<b>12</b>		<b>12.2±0.4</b>	<b>10.2±0.3</b>
<b>174-43/88 44-55</b>	<b>12</b>	<b>6</b>	<b>12.2±0.4</b>	<b>6.6±0.5</b>
<b>174-43/169 75-48</b>	<b>10</b>		<b>9.7±0.6</b>	<b>4.2±0.9</b>
<b>174-43/129 48-39</b>	<b>10</b>	<b>6</b>	<b>9.3±0.6</b>	<b>5.9±0.4</b>
<b>174-43/59 08-68</b>	<b>14</b>	<b>13</b>	<b>14±0.5</b>	<b>13.3±0.5</b>



# Processing of “Energy-Transmutation” experiment data

Initializing of uranium decay with neutrons produced in inelastic proton interaction with lead target.  
Measurement of neutron flux from decay process.



Neutron flux is connected with number of radiator nuclei fragments.

=> Measurement of fragments traces in plastic detector

The problem: many tracks cross each other when flux is large

Solution: use fuzzy logic approach

Usual logic:

$a=b \Rightarrow$  Yes (true) or No (falls)

Fuzzy logic:

With probability  $P$   $a=b$

With probability  $1-P$   $a \neq b$

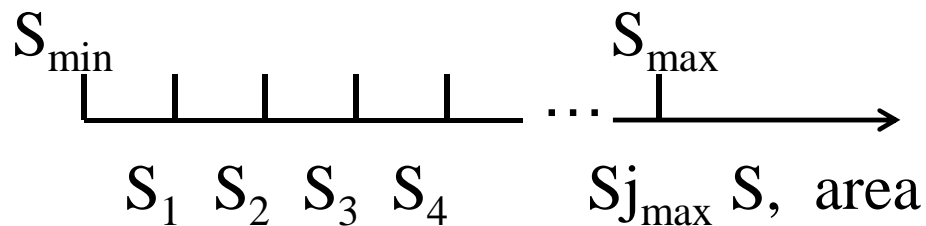


# Use probable description of crossing tracks number

Variable is the cluster area  $S_{CL}$ .

But the  $S_{CL}$  can corresponds to different tracks number

=> Probability matrix  $W$  is defined :



$$W(N, S_j), \quad S_j \in (S_{\min}, S_{\max}), \quad 1 \leq j \leq j^{\max}; 1 \leq N \leq N^{\max}$$

$$\hat{N}(S_j) = \sum_{N=1}^{N^{\max}} N \cdot W(N, S_j); \quad \bar{N} = \sum_{j=1}^{j^{\max}} n_j \cdot \hat{N}(S_j)$$

# The dialog window for training procedure

Form1

Choice of File | Conditions | Filtering | Contrast | Probability matrix | Clustering


N<sup>o</sup> Number of clusters  
Content of group 5 11

Filling of probability matrix

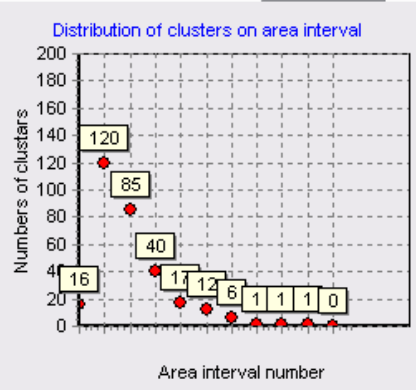
Number of crossing clusters  
 1  7  
 2  8  
 3  9  
 4  10  
 5  11  
 6  12

Number of Cluster 1  
Cluster area 203

Begin  
Continue



Distribution of clusters on area interval



Area interval number	Numbers of clusters
0	16
1	120
2	85
3	40
4	1
5	1
6	12
7	6
8	1
9	1
10	1
11	0

The initial probability matrix.

10	50	100	150	200	250	300	350	400	450	500	600	800	1000	1800
1.1	1.4	1.9	2.6	3.8	4.6	5.4	5.8	6.4	7.4	9.0	9.9	10.7	11.7	11.9

Distr.: NumbClusters = 283 NumbTracks = 554  
Intensity:  $554 / (80 \times 60 \text{ mkm}^2) = 1.158 \times 10^7 \text{ neutron/cm}^2$

Result.bmp Save of picture

# The dialog window for training procedure

Form1

Choice of File | Conditions | Filtering | Contrast | Probability matrix | Clustering

N<sup>o</sup> Number of clusters  
Content of group 5 11

Filling of probability matrix

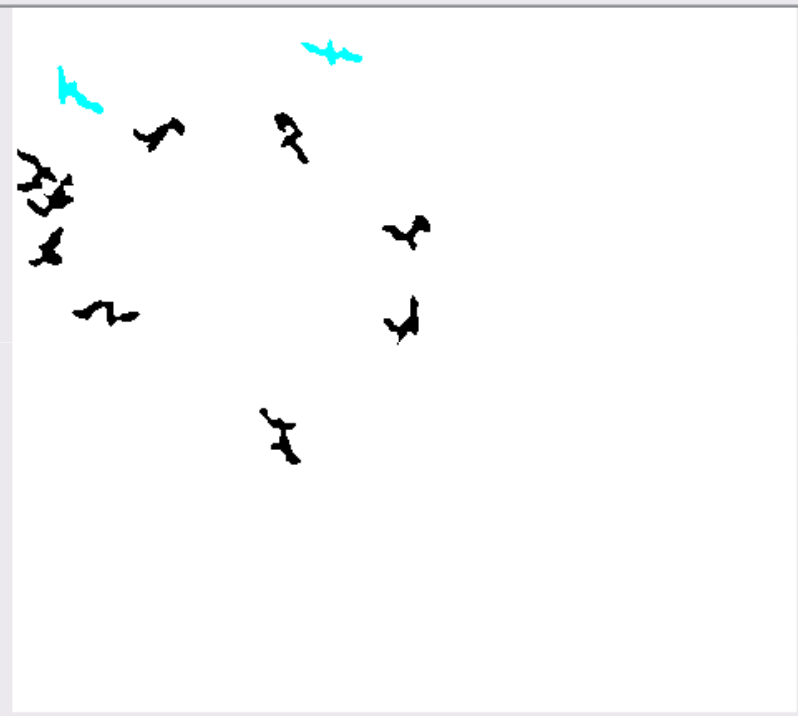
Number of crossing clusters  
 1  7  
 2  8  
 3  9  
 4  10  
 5  11  
 6  12

Number of Cluster 2

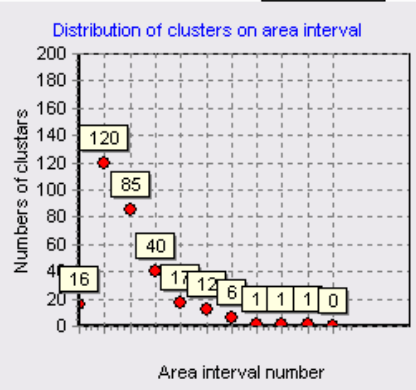
Cluster area 241

Begin

Continue



Distribution of clusters on area interval



Area interval number	Numbers of clusters
10	16
50	120
100	85
150	40
200	1
250	1
300	12
350	6
400	1
450	1
500	1
550	0

Result.bmp Save of picture

The initial probability matrix.

	10	50	100	150	200	250	300	350	400	450	500	600	800	1000	1800
1,2	1,9	2,6	3,6	4,2	4,6	5,4	5,8	6,4	7,4	9,0	9,9	10,7	11,7	11,9	

# The dialog window for training procedure

Form1

Choice of File | Conditions | Filtering | Contrast | Probability matrix | Clustering

N<sup>o</sup> Number of clusters  
Content of group 5 11

Filling of probability matrix

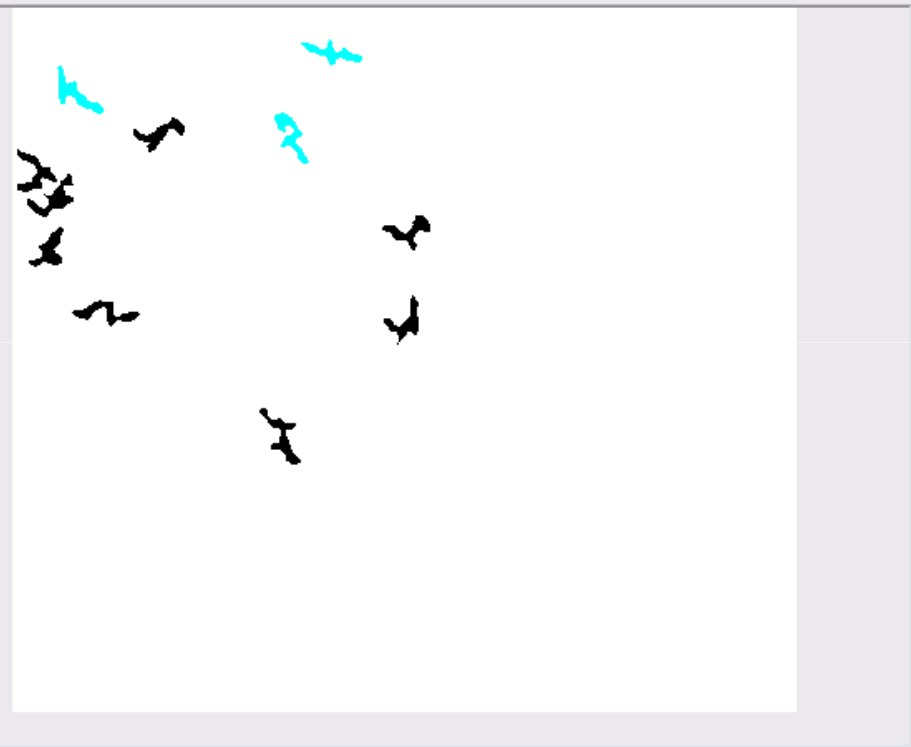
Number of crossing clusters  
 1  7  
 2  8  
 3  9  
 4  10  
 5  11  
 6  12

Number of Cluster 3

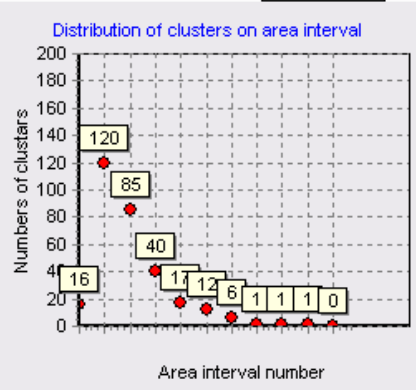
Cluster area 220

Begin

Continue



Distribution of clusters on area interval

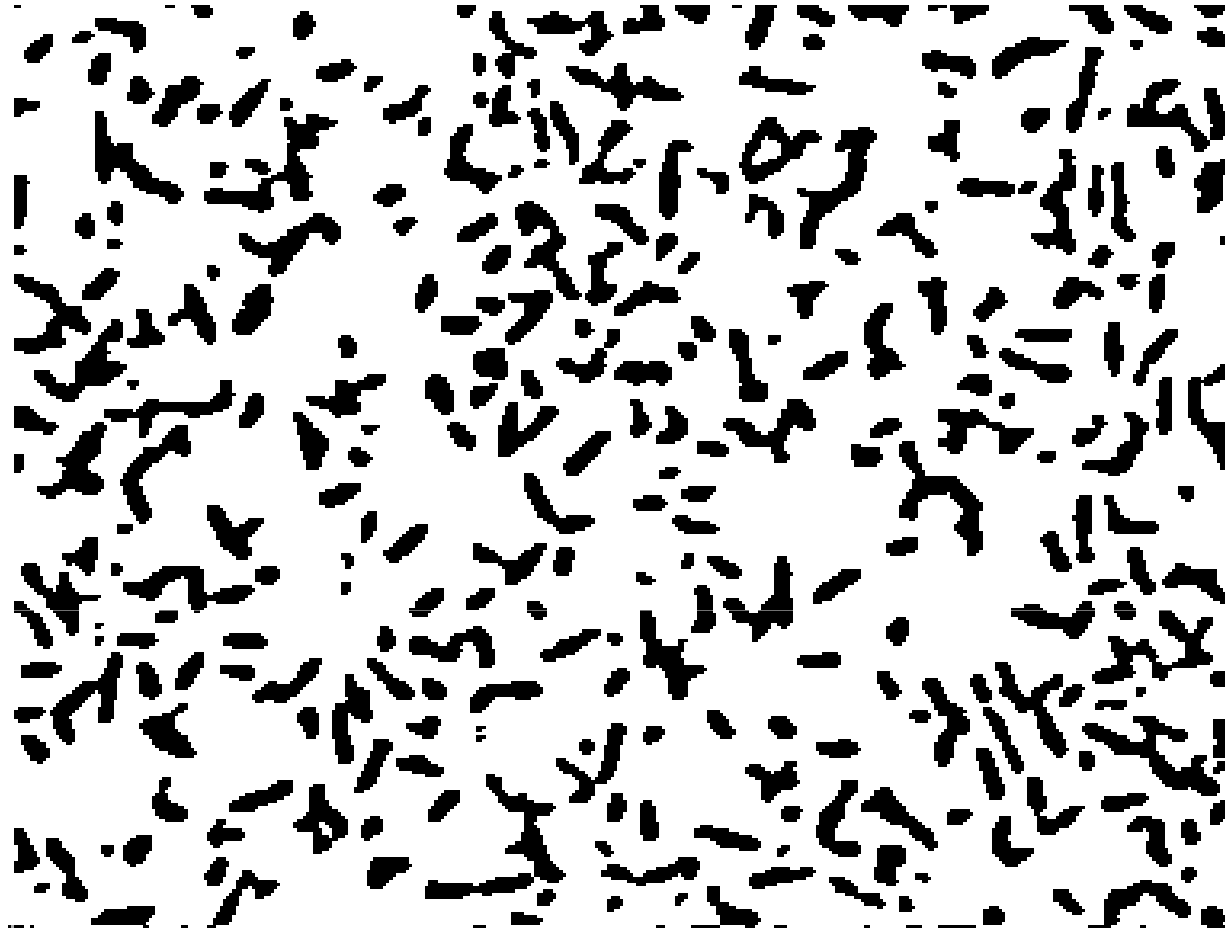


Area interval number	Numbers of clusters
10	16
50	120
100	85
150	40
200	1
250	1
300	6
350	1
400	1
450	1
500	0

The initial probability matrix.

	10	50	100	150	200	250	300	350	400	450	500	600	800	1000	1800
1,2	1,9	2,6	3,6	4,2	4,6	5,4	5,8	6,4	7,4	9,0	9,9	10,7	11,7	11,9	

Result.bmp Save of picture



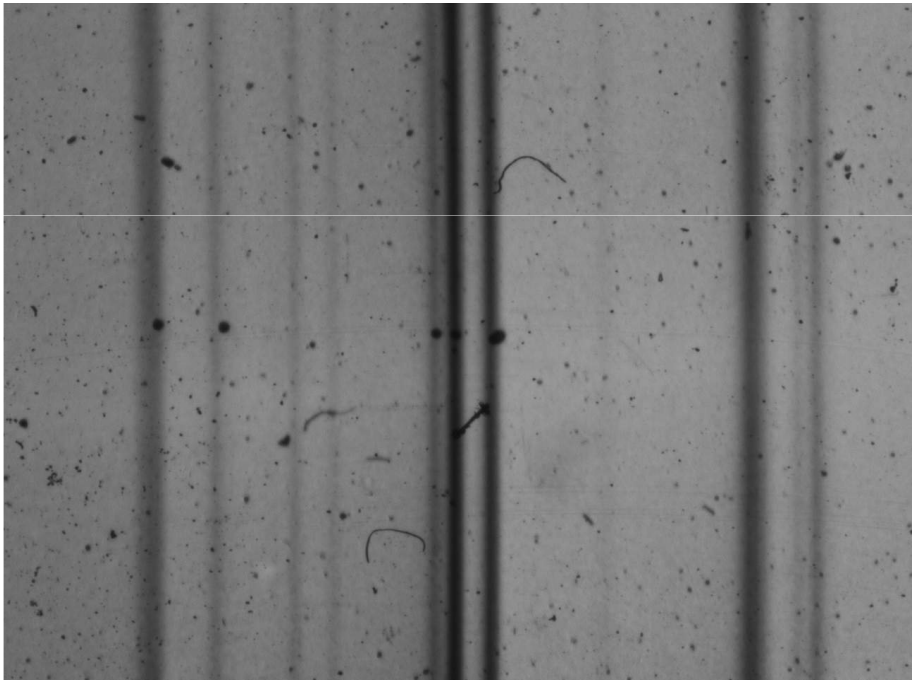
Result:

There are **283 clusters**. It corresponds to **554 tracks**.

==> The **flux is  $\sim 1,15 \cdot 10^7$**  fragments/cm<sup>2</sup>.

# Investigation of spectra of inner-conversion electrons

(lanthanides isotopes of Er, Ho, Dy, Gd ...).

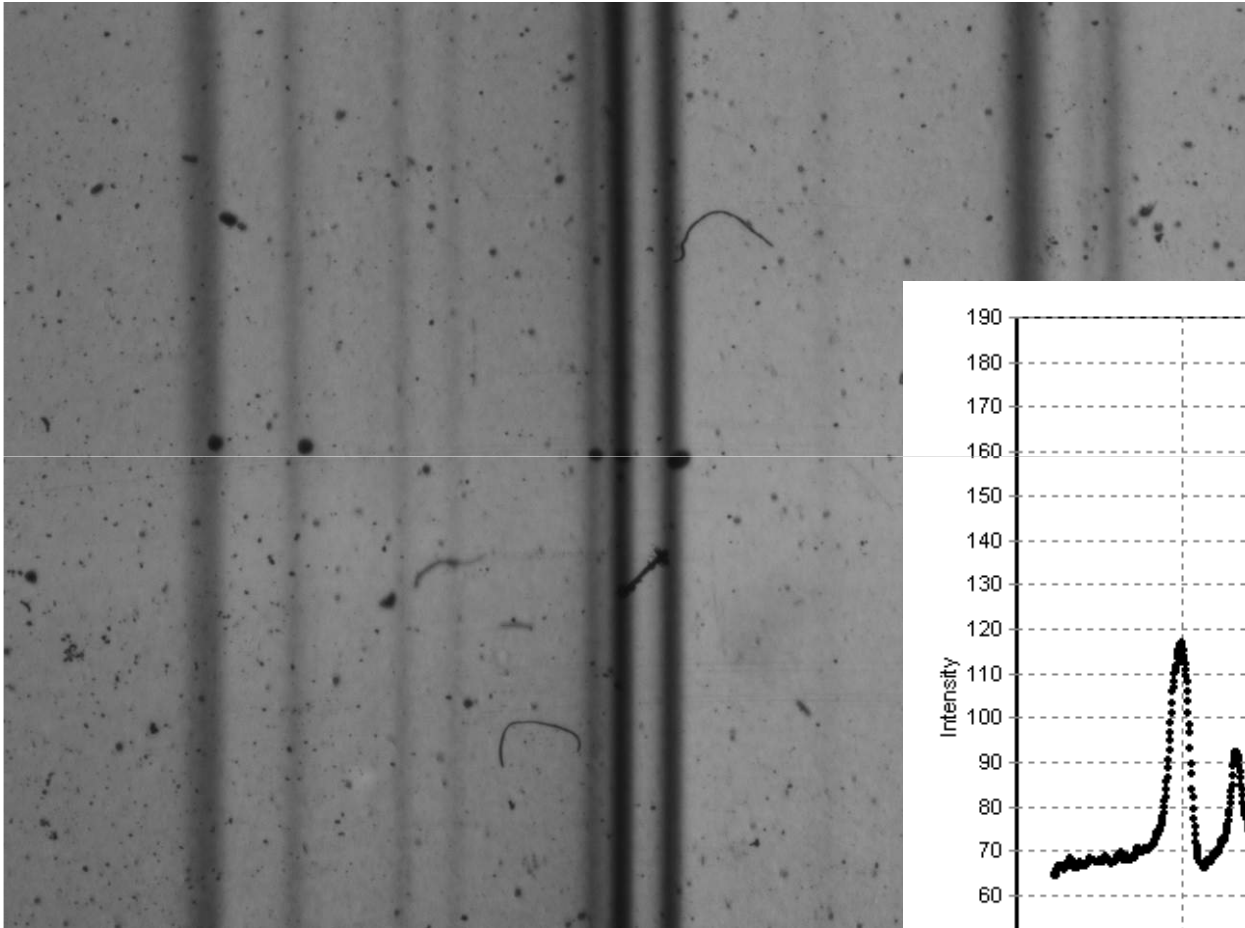


1. Preliminary filtering is not used because it is necessary to compare brightness of lines but filters change it.
2. The sum of vertical pixels amplifies useful signal and reduces noise.
3. Noise: spots, scratches, non-uniform of illumination.

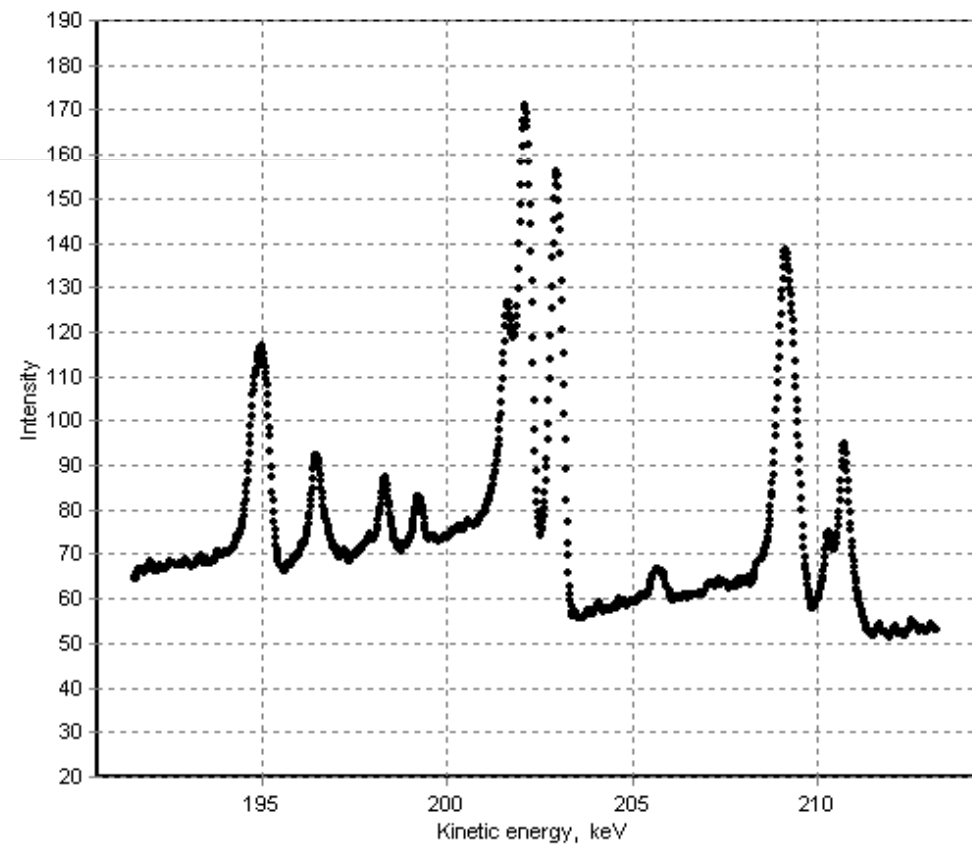


# Investigation of spectra of inner-conversion electrons

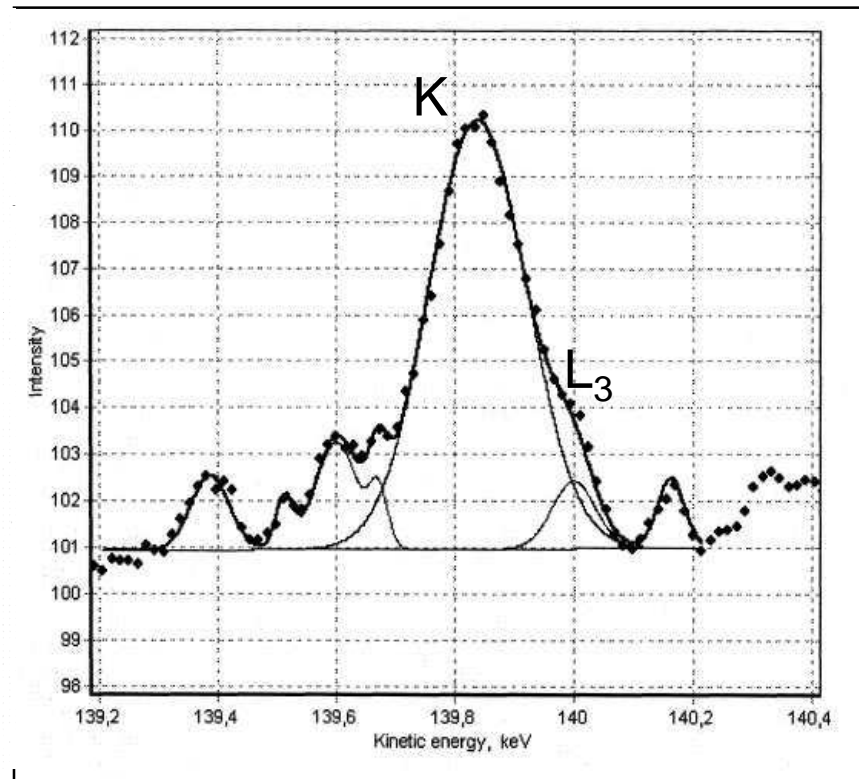
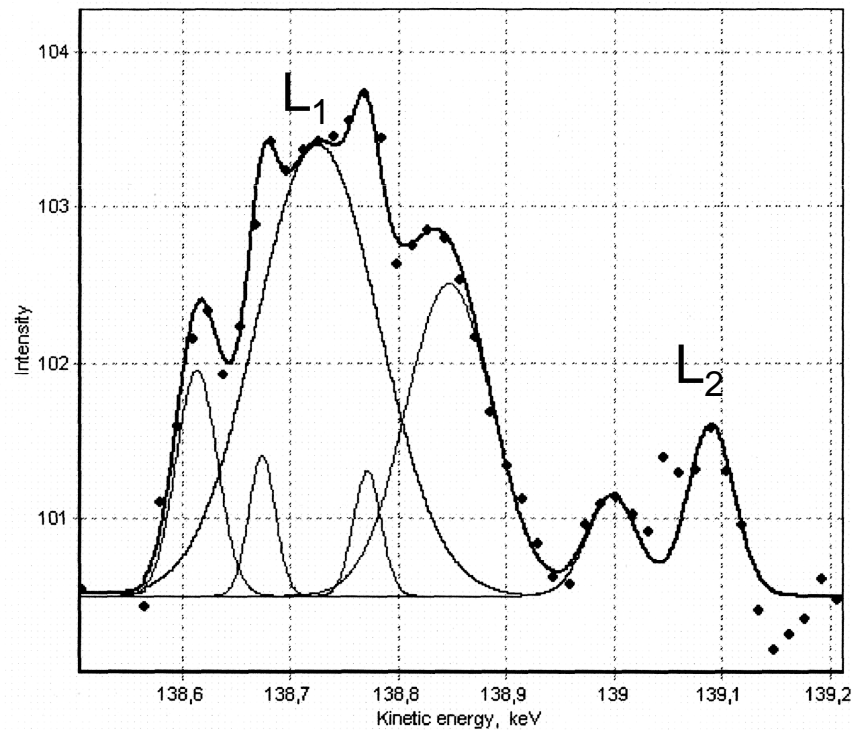
(isotopes of Er, Ho, Dy, Gd ...).



There are 40 fields of view of such kind

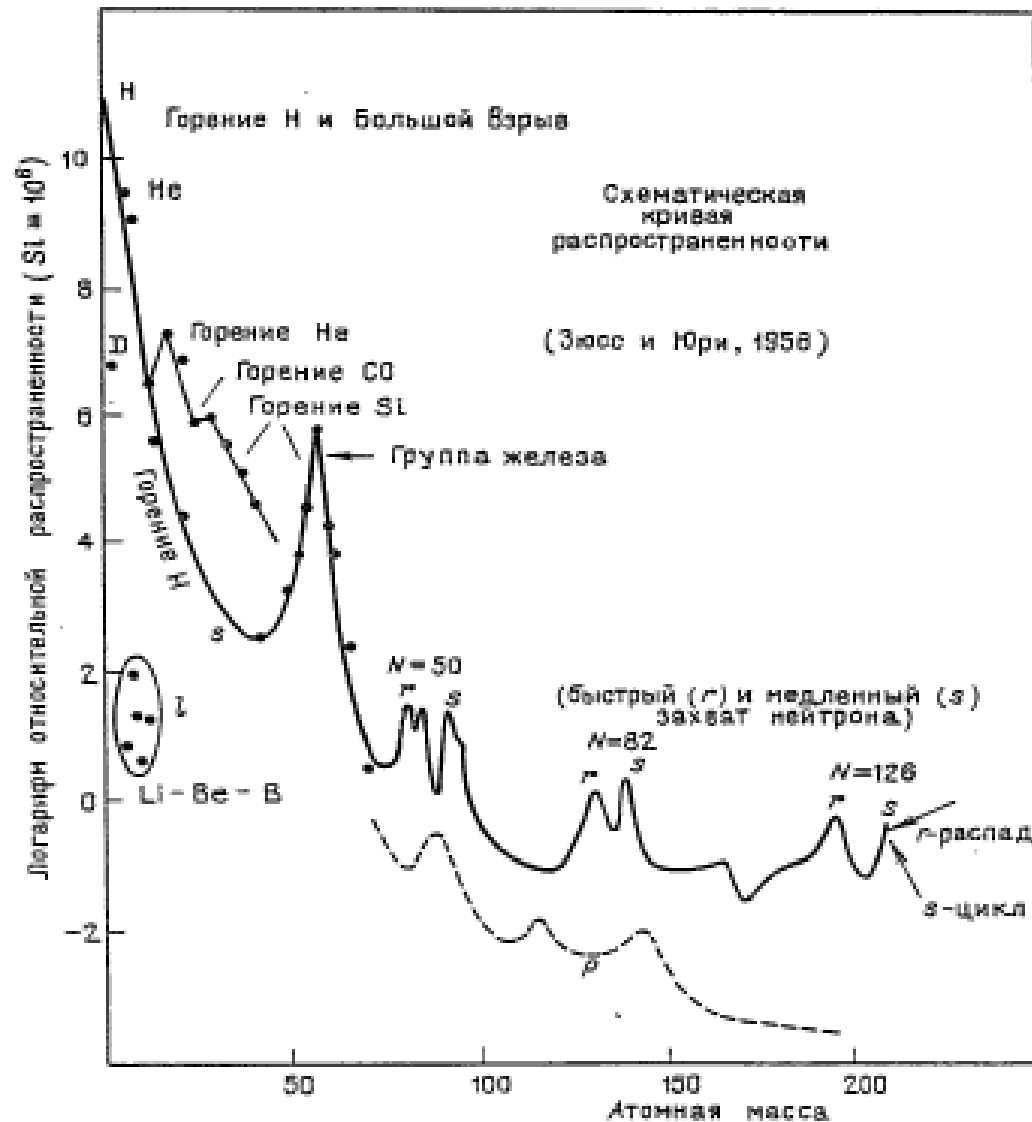


# Identification of lines $L_1$ , $L_2$ , $L_3$ ICN of $^{161}\text{Ho}$ ( $L_2$ , $L_3$ are new; known line K scales of energy)

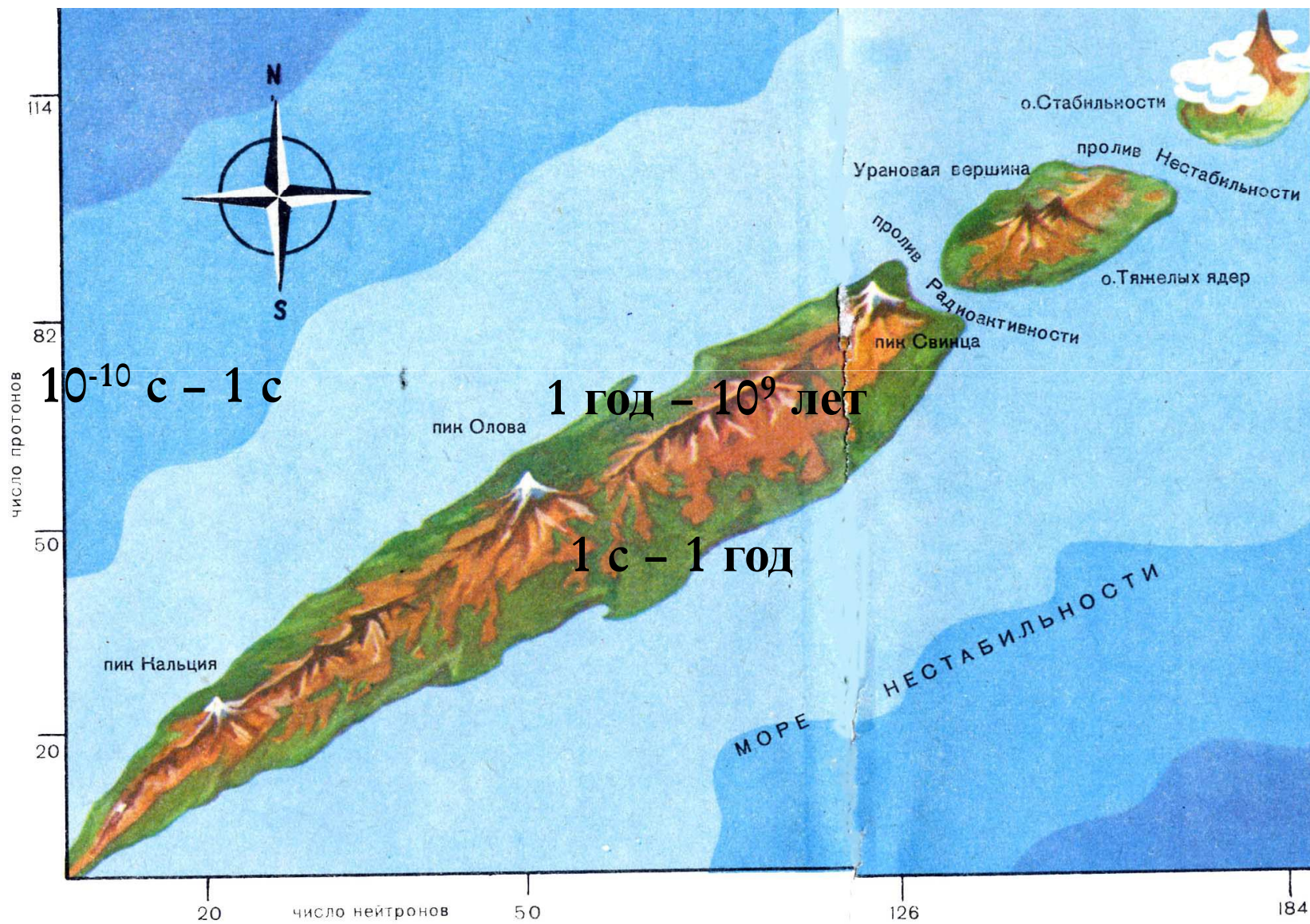


# Investigation of olivine from meteorites

## Study of charge spectrum super heavy nuclei in space



# The position of stability island (neutron-proton coordinates)



**«ОЛИМПИДА»**

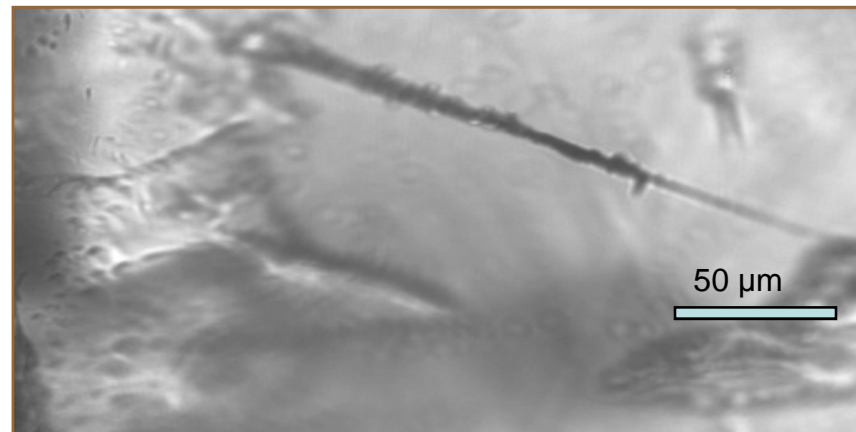
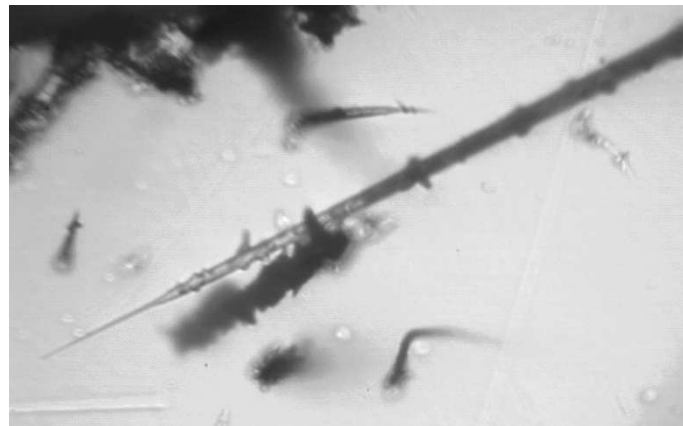
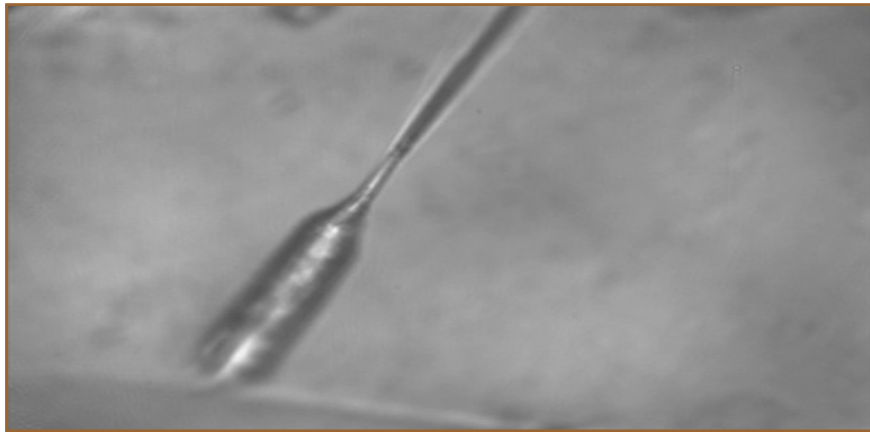
**«ОЛИПИЯ»**

(«**ОЛИ**вины из **М**етеоритов  
– **П**оиск тяжелых и  
сверхтяжелых **Я**дер»)

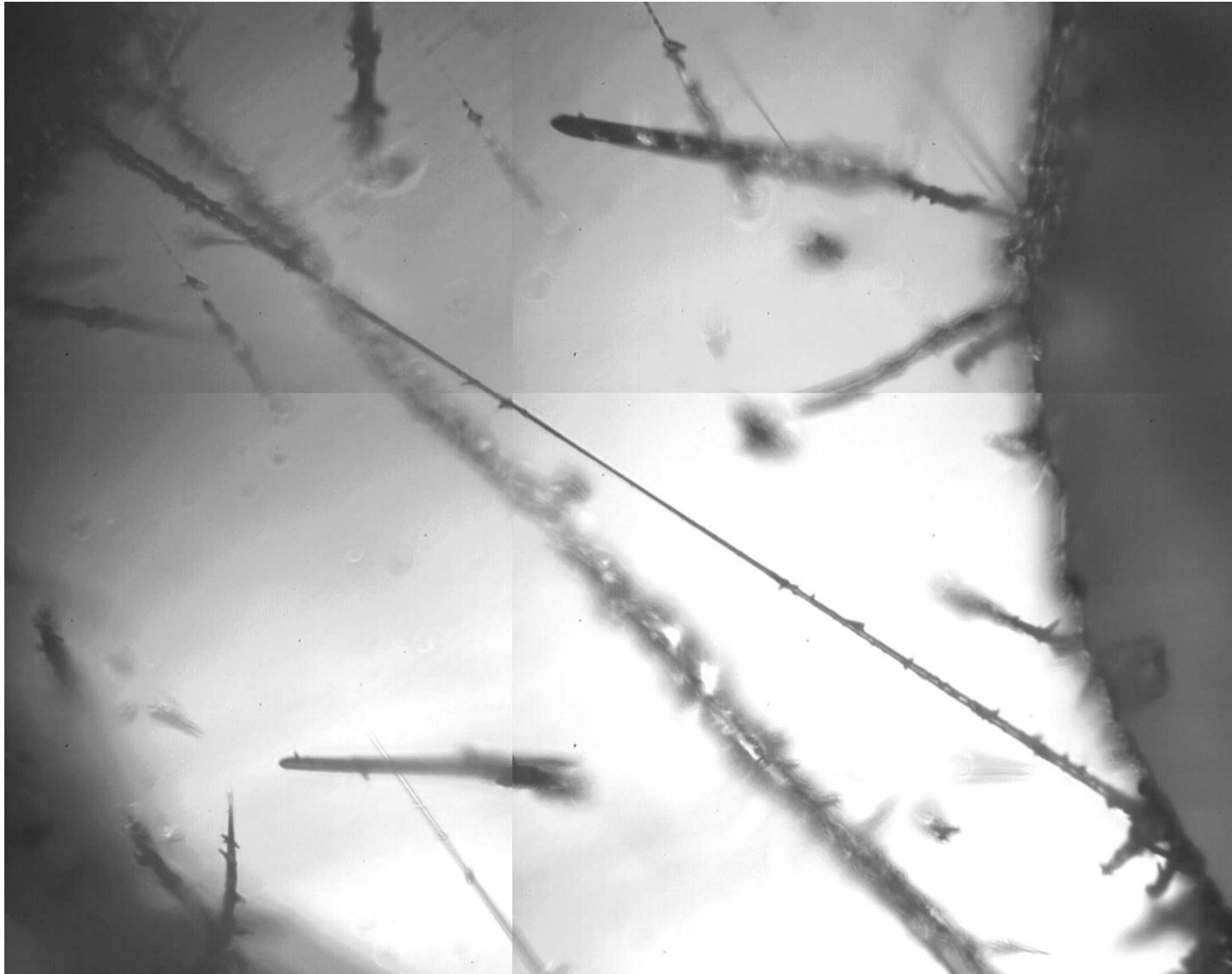


# Epoxy tablet with olivine crystals





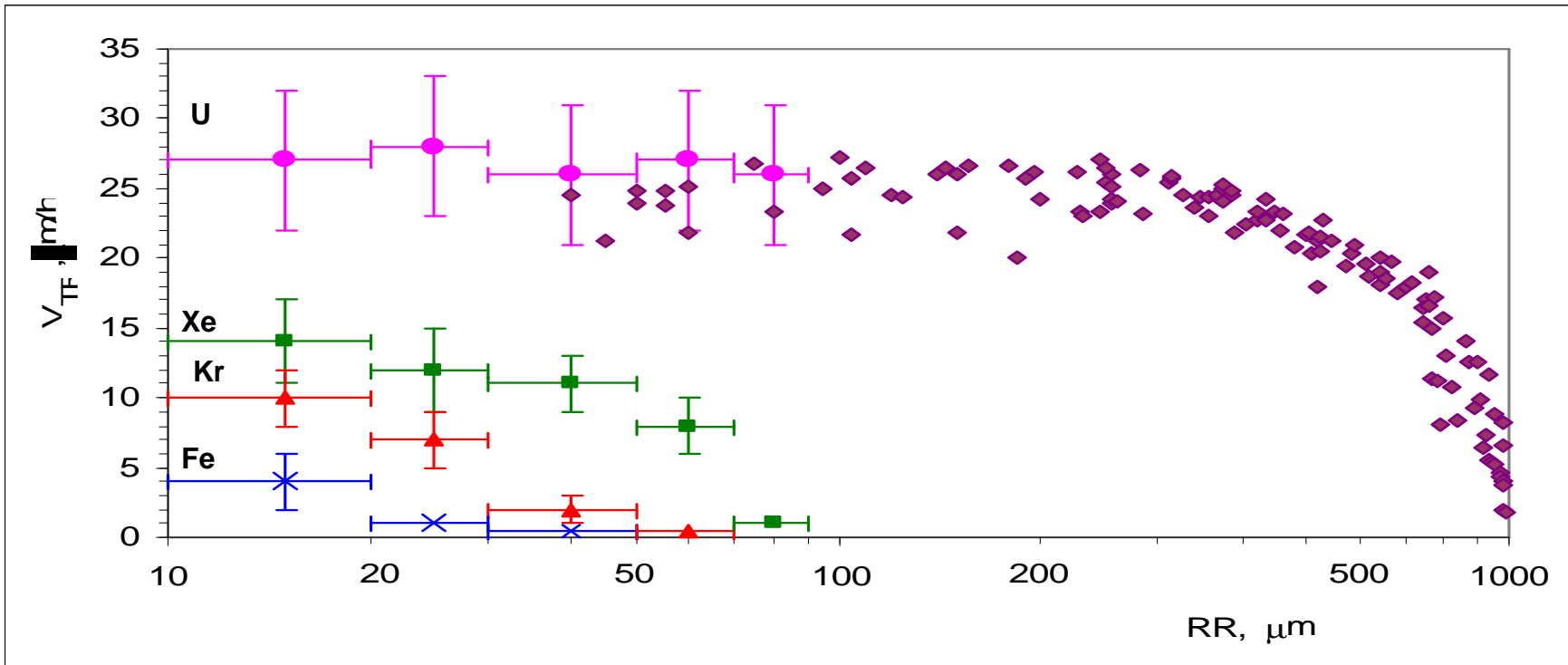
Sample of very long track (L>700 mcm; Z>70)





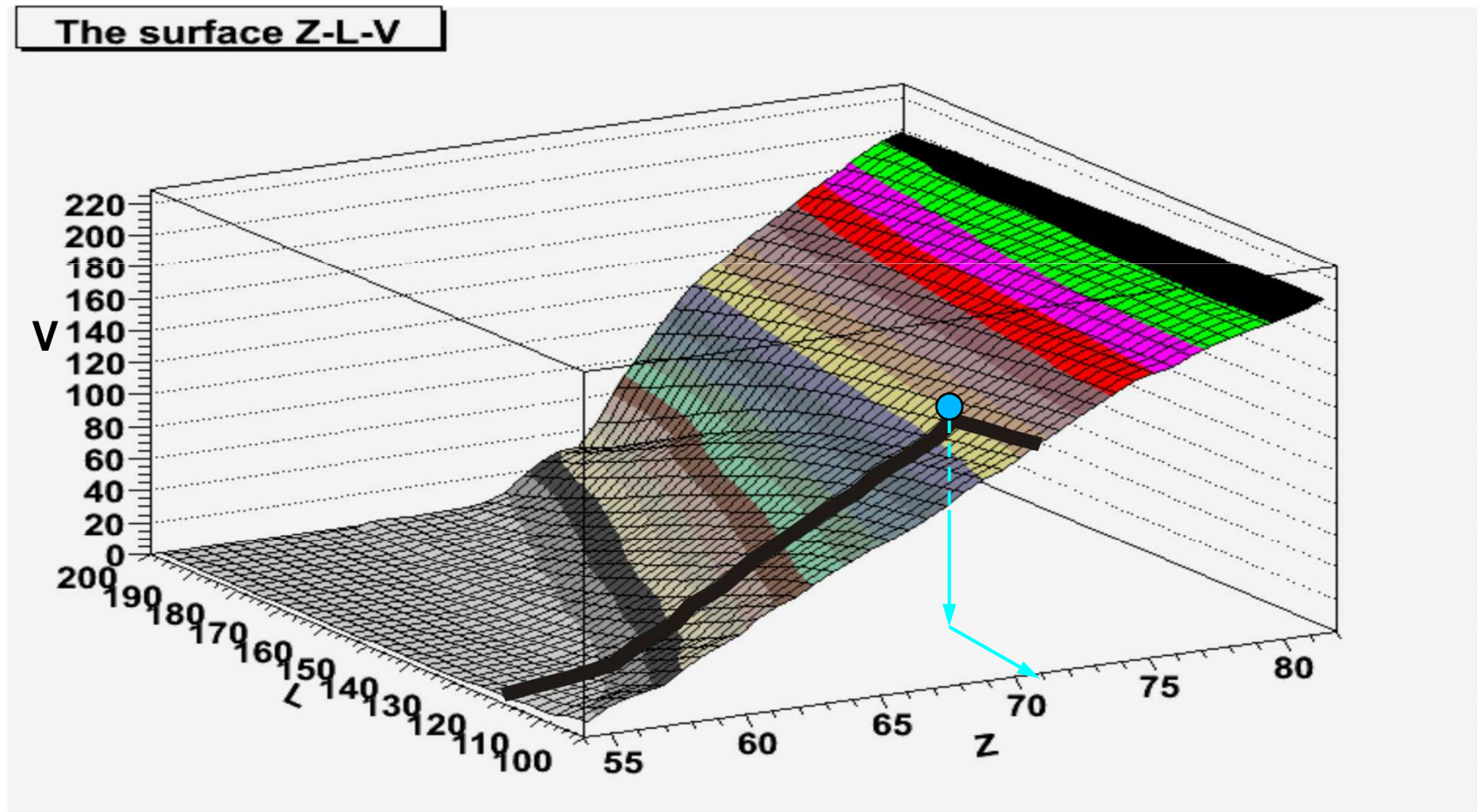
**Track-etch rate ( $V_{TR}$ ) in the pallasite Marjalahti olivine crystals from accelerated Kr, Xe and U ions and Fe cosmic ray nuclei**

Ion	E, MeV/nucl	$V_{TR}$ ( $\mu\text{m/h}$ ) at RR ( $\mu\text{m}$ )				
		$15 \pm 5$	$25 \pm 5$	$40 \pm 10$	$60 \pm 10$	$80 \pm 10$
Fe	>5	$4 \pm 2$	$\sim 0.1$	$\sim 0.05$	–	–
Kr	10.2	$10 \pm 2$	$7 \pm 2$	$2 \pm 1$	$\sim 0.1$	–
Xe	11.4	$14 \pm 3$	$12 \pm 3$	$11 \pm 2$	$8 \pm 2$	$\leq 1$
U	11.4	$27 \pm 5$	$28 \pm 5$	$26 \pm 5$	$27 \pm 5$	$26 \pm 5$

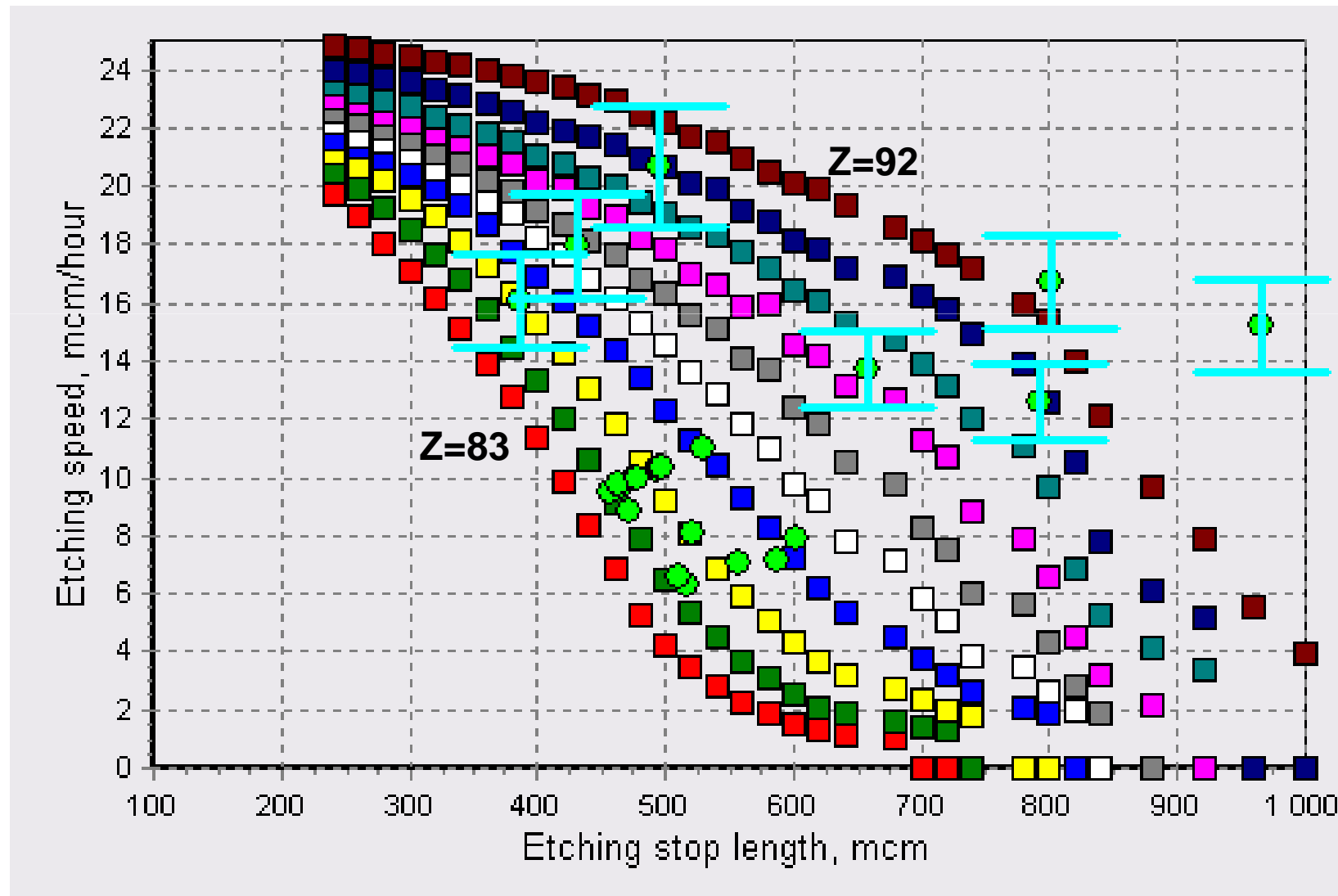


**Measured  $V_{TR}$  values at the certain residual range (RR) of accelerated Kr, Xe and U ions and GCR Fe nuclei in olivine crystals from the Marjalahti pallasite**  
**Data for U at RR > 100  $\mu\text{m}$  by Perron and Maury, 1986**

Dependence between charge  $Z$ , track length  $L$  and etch rate  $V$ .

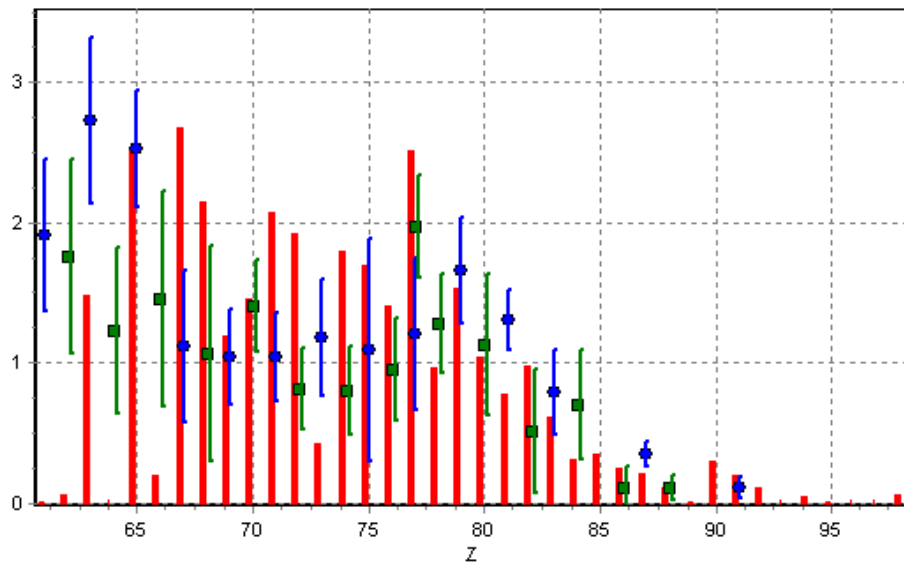


# The result of heavy nuclei processing $Z > 83$

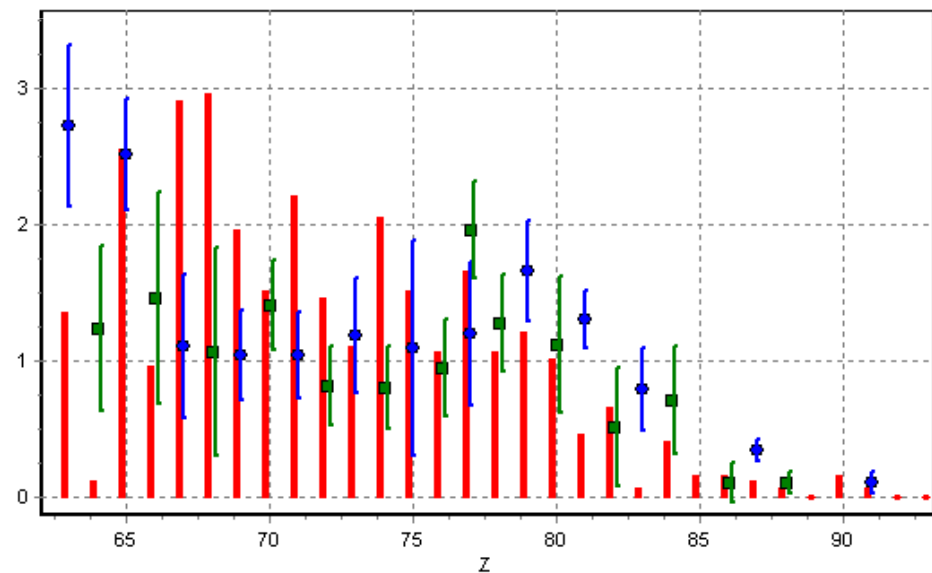


# Abundance of super heavy nuclei in comparison with Ariel 6 – o and HEAO 3 - □ data. (Abundance of Fe = $10^6$ )

The abundance of elements in olivin from meteorites.



The abundance of elements in olivin from meteorites.



Besides three nuclei where found whose charge is more 100.

Our calculations give estimation

$$106 < Z < 130$$

$$\text{Lifetime} > 3000 \text{ years}$$

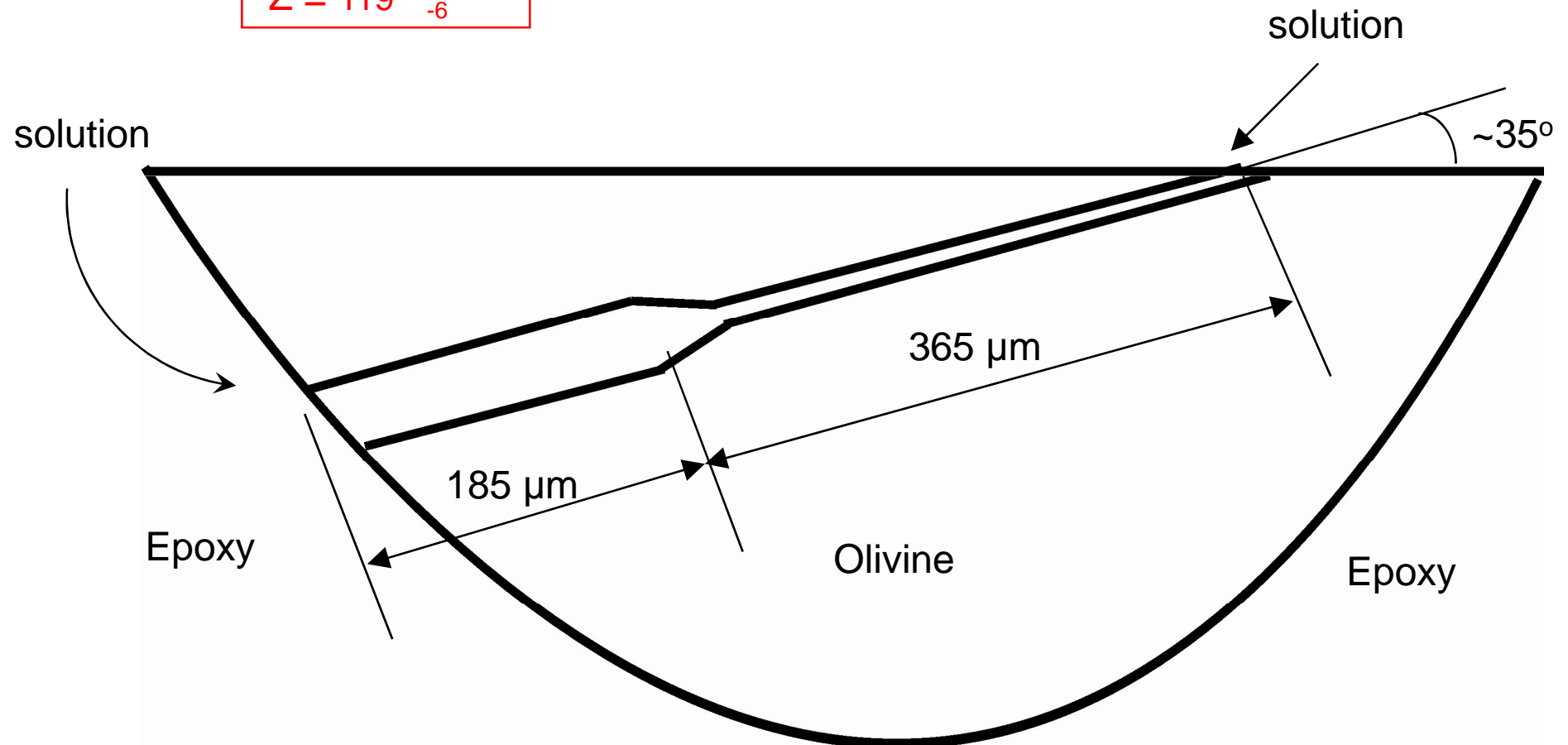
The charge of one of them is defined more exactly

The total track length in olivine is  $\sim 550 \mu\text{m}$ . The etch time is 8 hours.

But the etch solution can reach to track from both sides.

$\Rightarrow$  The minimum etch rate is about  $35 \mu\text{m/h}$

$$Z = 119^{10}_{-6}$$







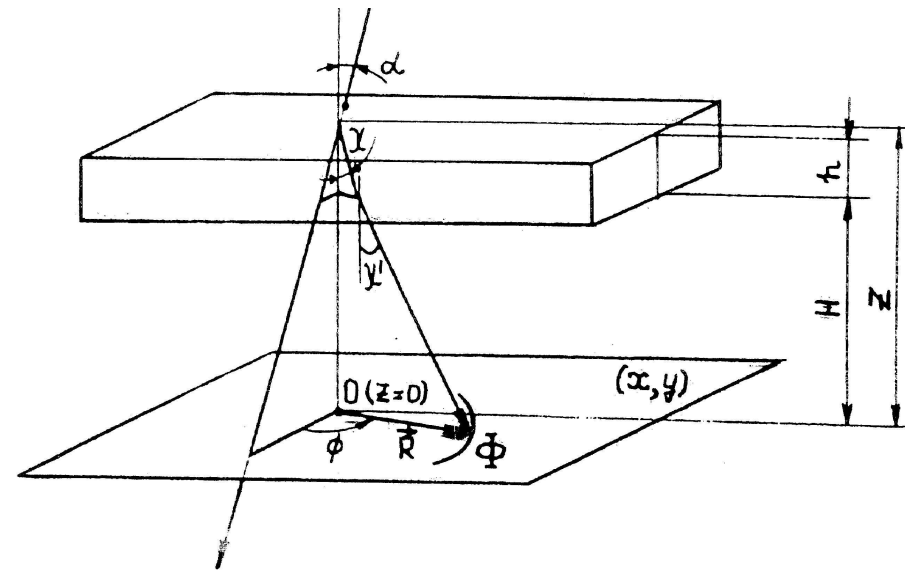
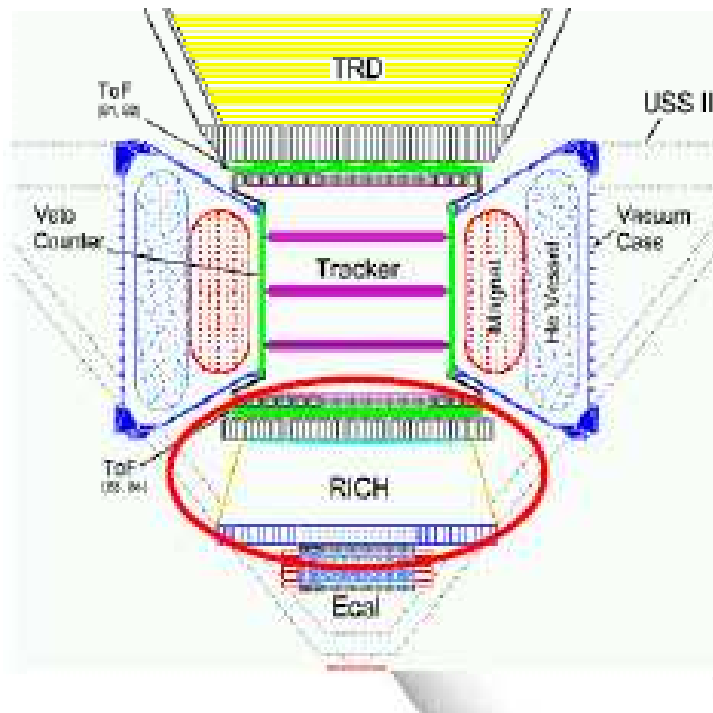


# Обработка данных RICH-детектора с помощью нейронной сети

(определение заряда и скорости)

Alpha Magnetic Spectrometer

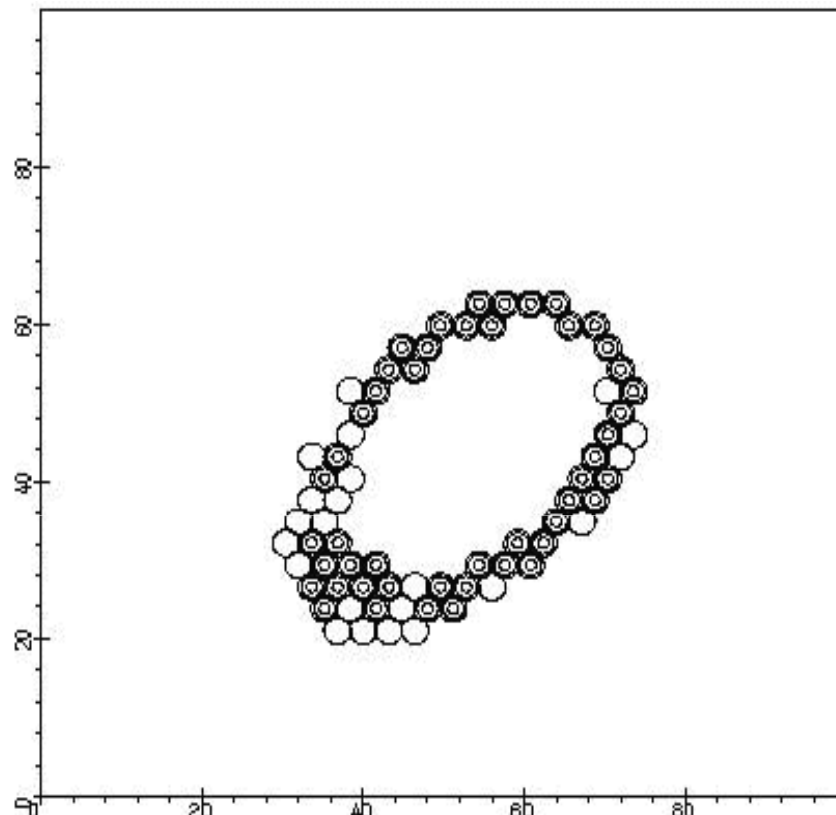
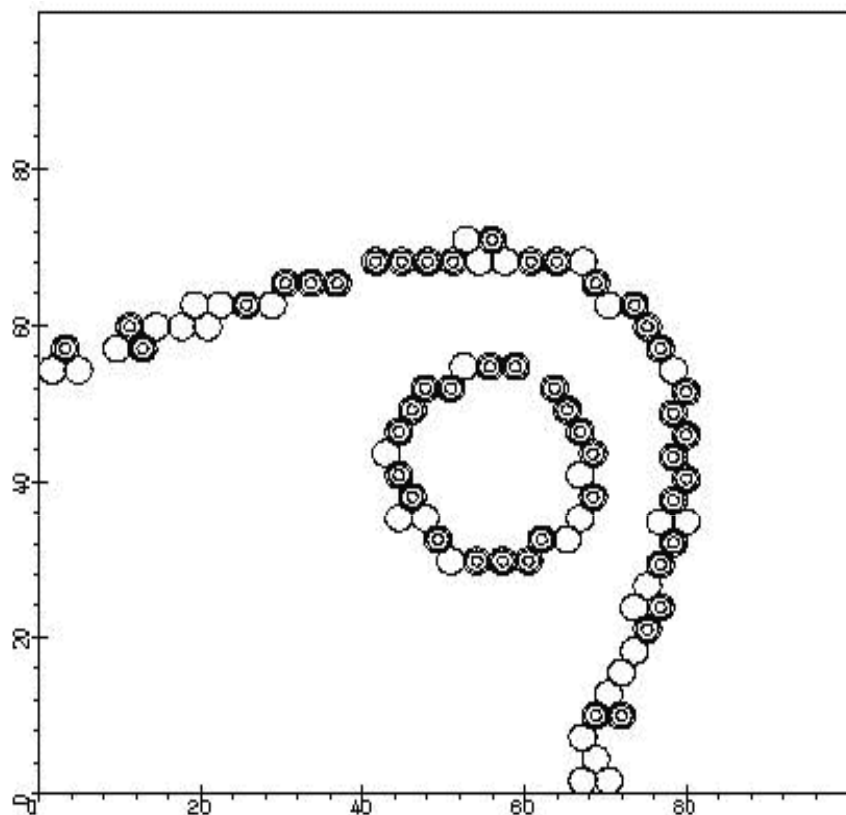
(AMS)



$$\cos(\chi) = 1/(\beta n) \quad (\beta n > 1)$$

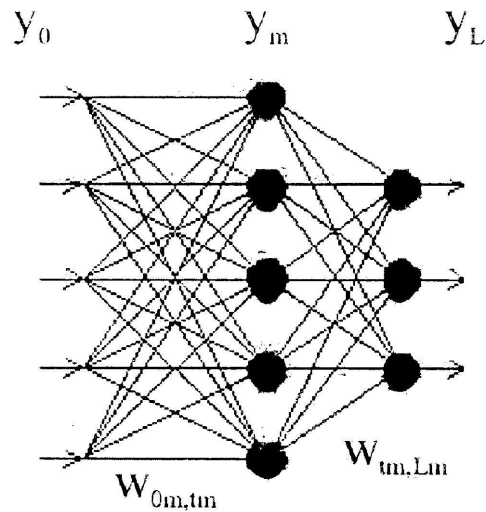
$$d^2N/d\lambda dx = (2\pi/137) \cdot (Ze)^2 (1 - 1/\beta^2 n^2(\lambda)) \cdot /\lambda^2$$

# Результат работы RICH-детектора



Необходимые данные: координаты точки входа частицы ( $x$ ,  $y$ , углы  $\theta$ ,  $\varphi$ ), амплитуда и координаты сработавших ФЭУ.

# Принцип работы нейронной сети Хопфилда (обучение с обратным распространением ошибки)



$$s_{j,m} = f \left( \sum_{i=1}^{N_{m-1}} s_{i,m-1} \cdot W_{j,m;i,m-1} - b_{j,m} \right)$$

$$f(x) = (1 + e^{-x})^{-1}$$

Входной вектор: полная амплитуда, ширина фигуры, образуемой сработавшими ФЭУ, и азимутальный угол  $\varphi$

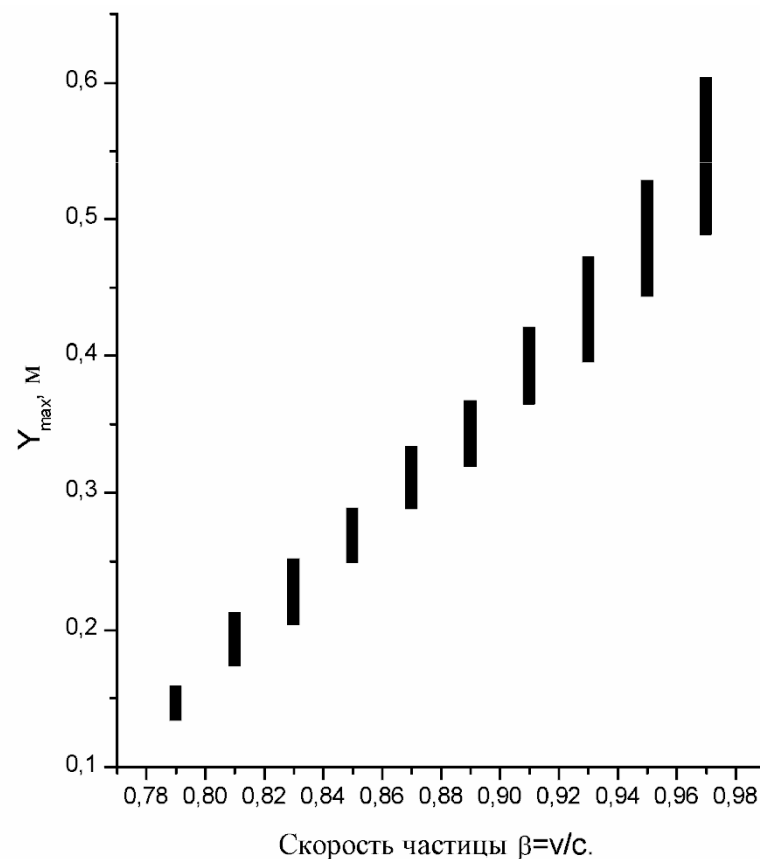
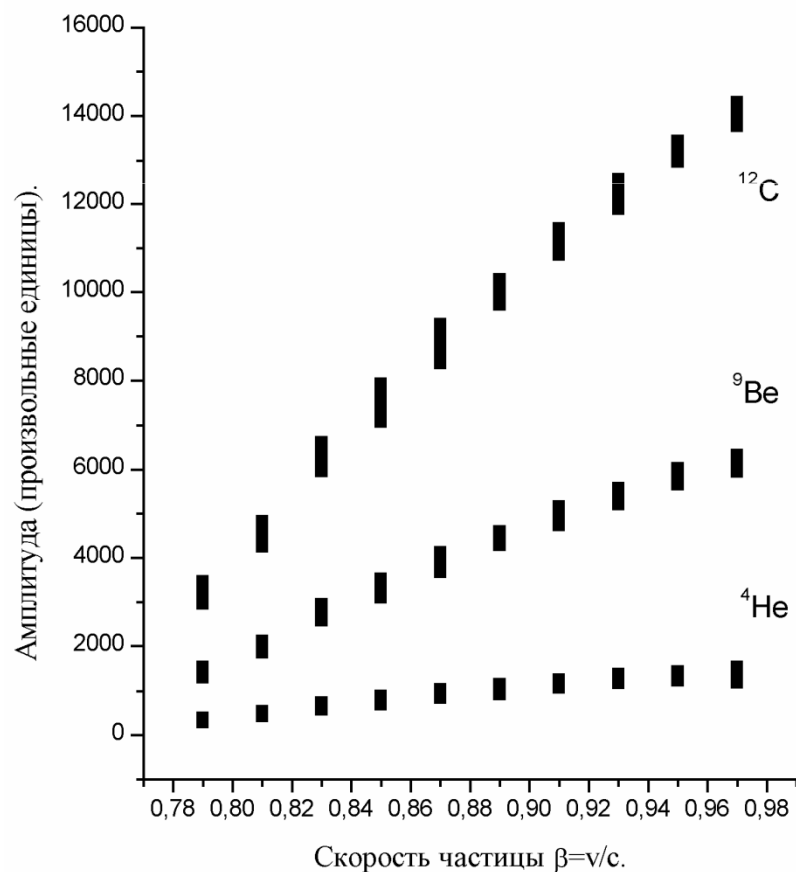
Выходной вектор: заряд, скорость

Не нужно решать обратную задачу

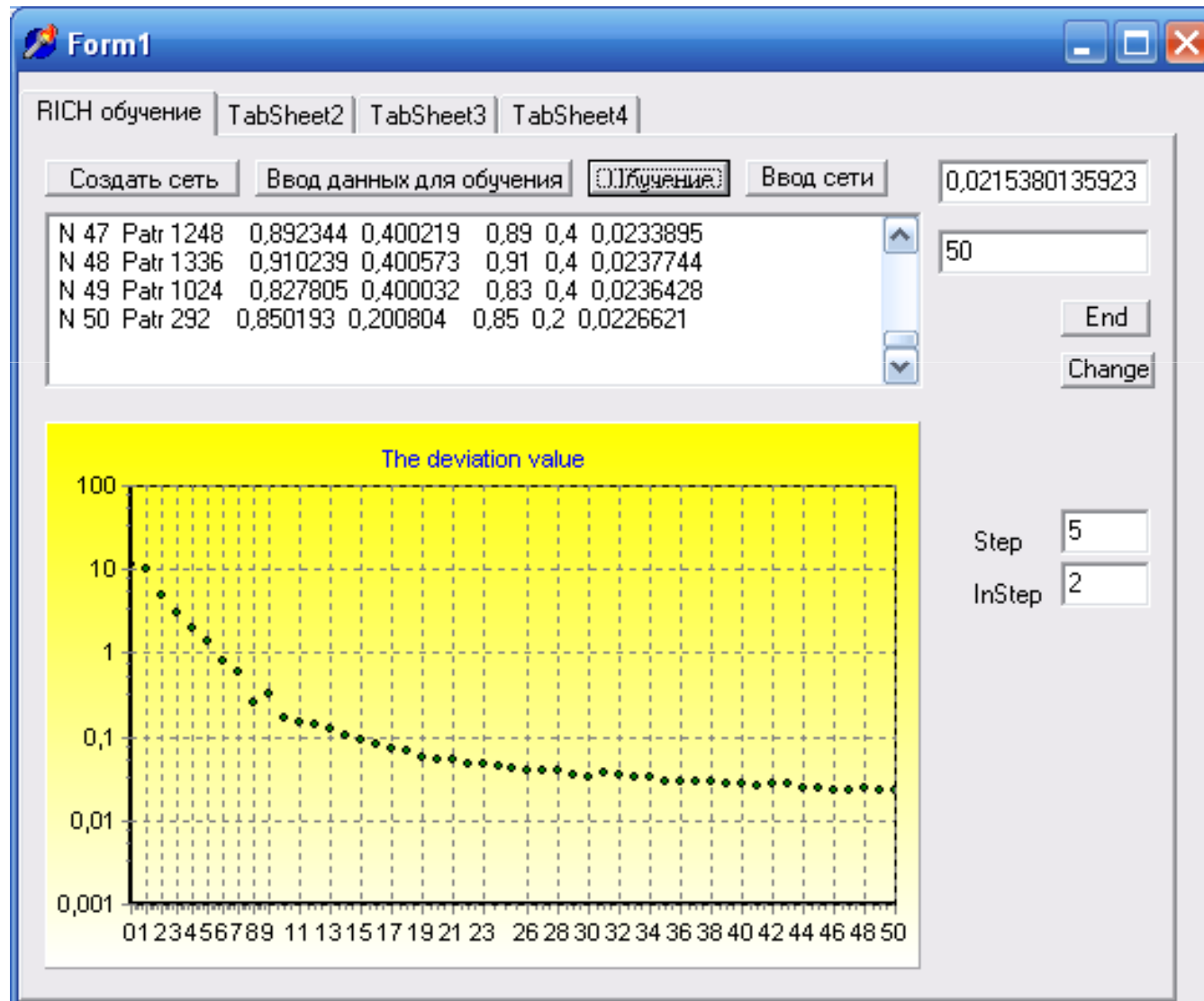
Входной вектор: полная амплитуда и координаты  
сработавших ФЭУ, азимутальный угол  $\varphi$

Выходной вектор: заряд, скорость

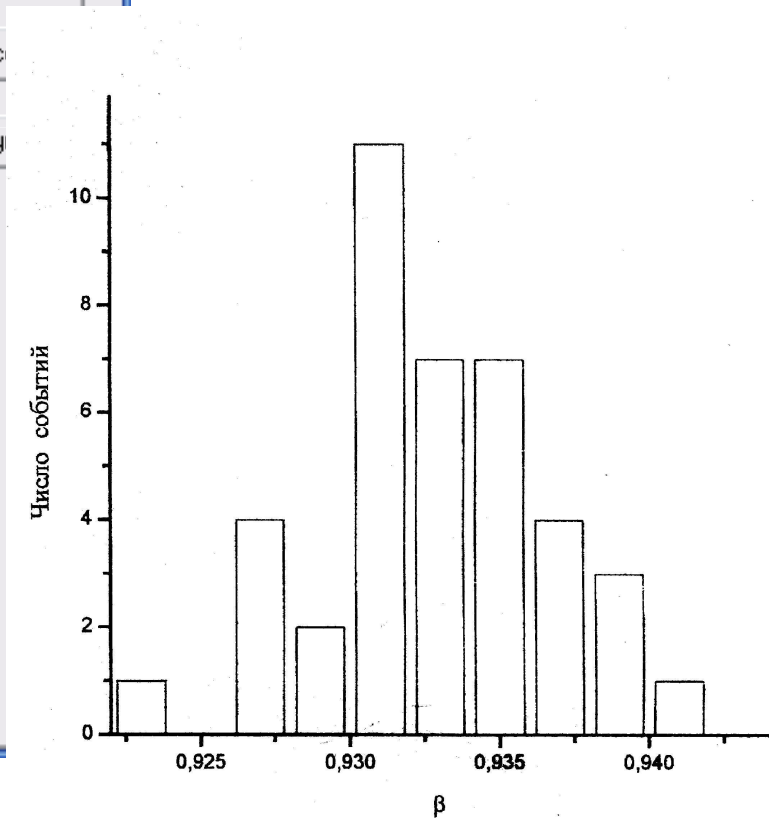
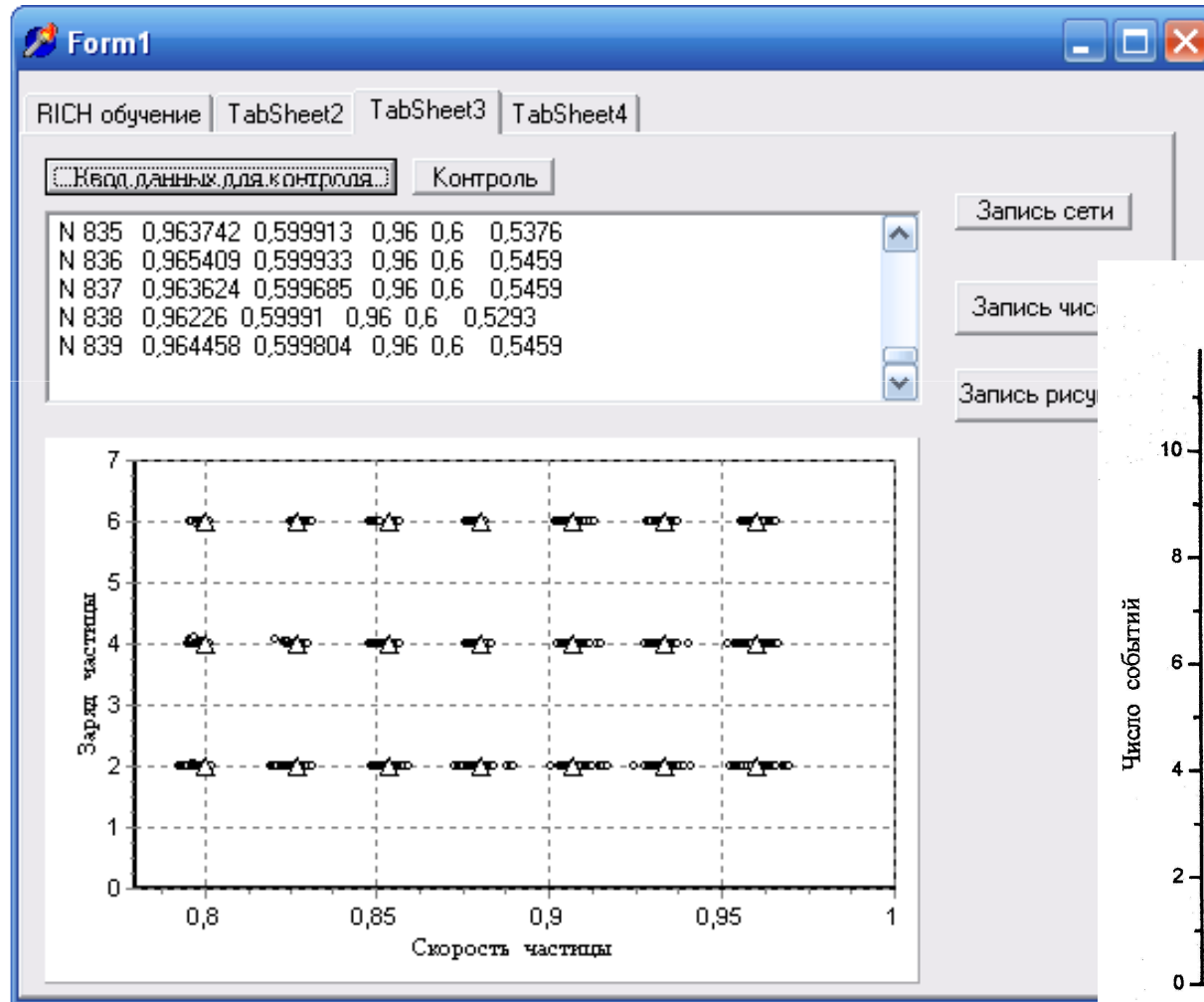
Обучение: 2400 событий (3 заряда, 10 скоростей)



# Окно пользовательской программы



# Результат распознавания 840 контрольных событий (точность по Z ~ 100 %, по $\beta$ ~ 96%)







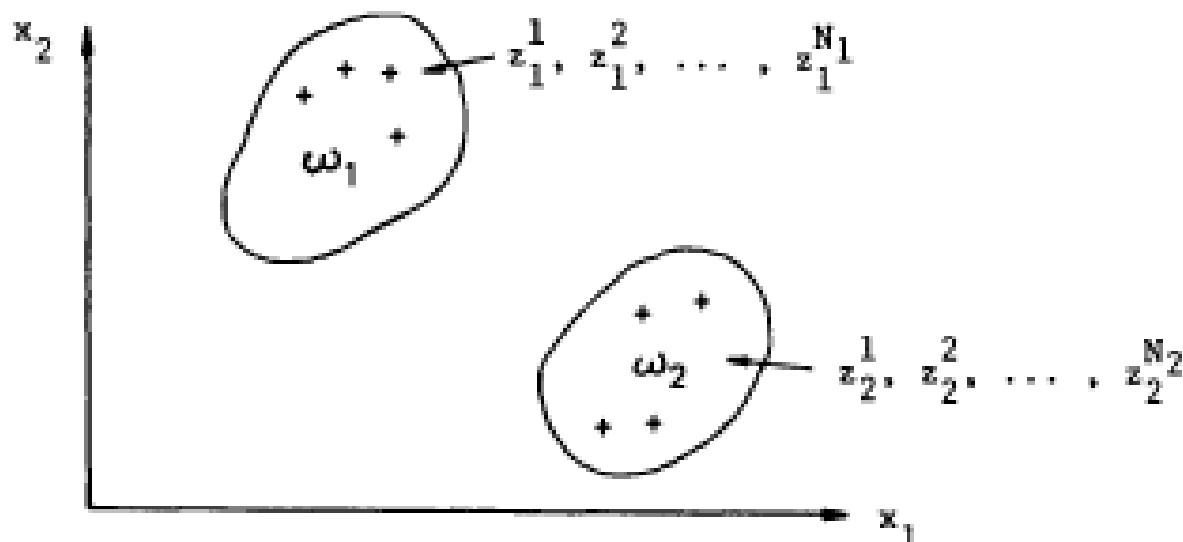




# Пространство признаков

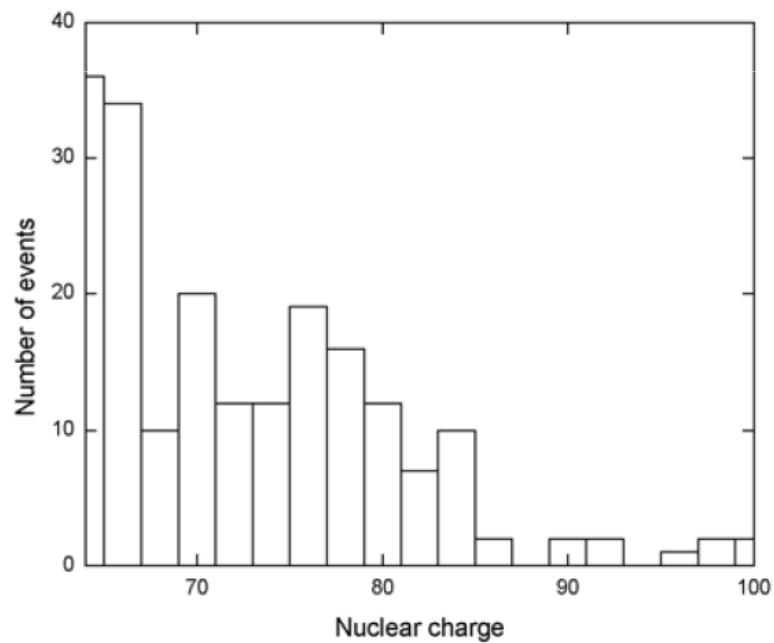
Эффективность обработки изображений решающим образом зависит от правильного выбора характеристик кластеров, подлежащих дальнейшему анализу.

Разделение двух классов  $\omega_1$  и  $\omega_2$  в пространстве признаков  $x_1$  и  $x_2$ .

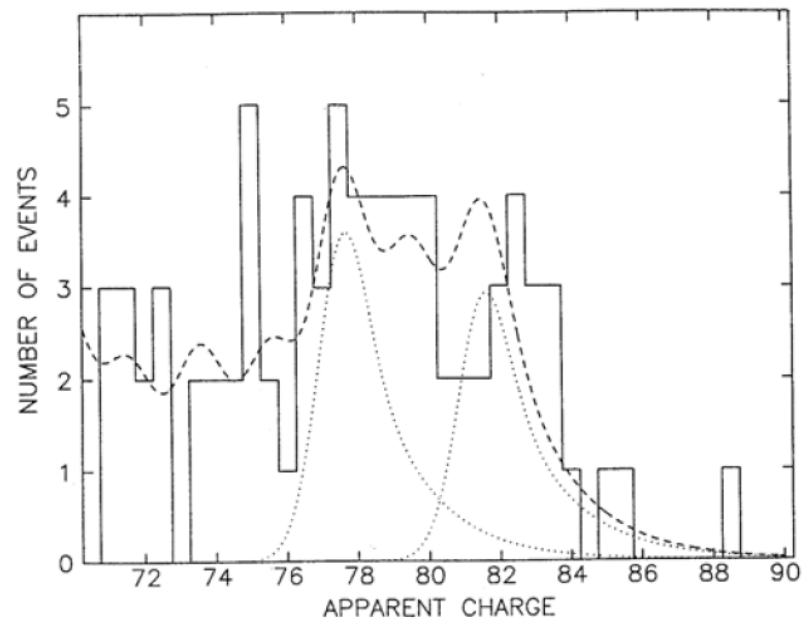


**Skylab**; E.K. Shirk, P.B. Price; 1978  
 $Z \geq 65$ . Спутник, 430 км, 230 дней,

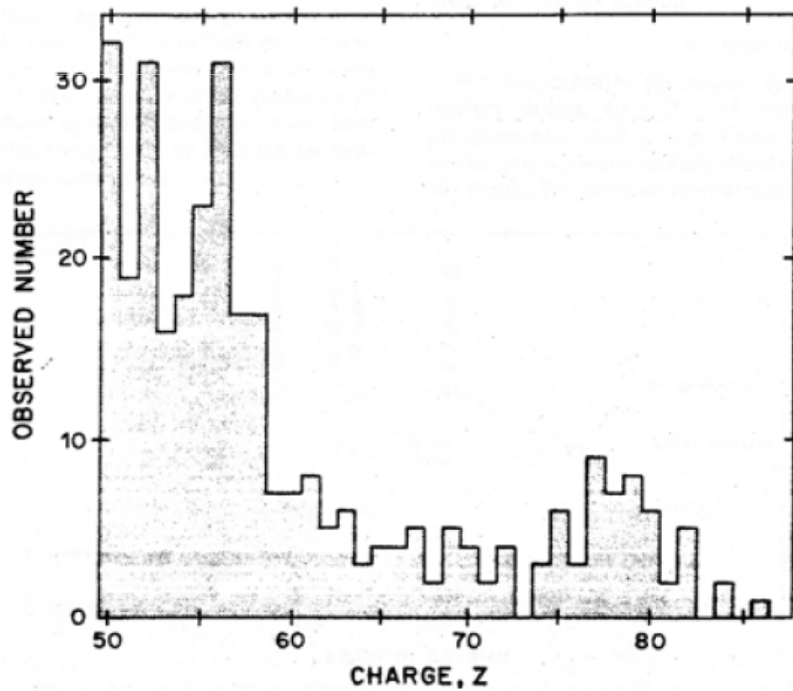
пластик (стопки лексана)



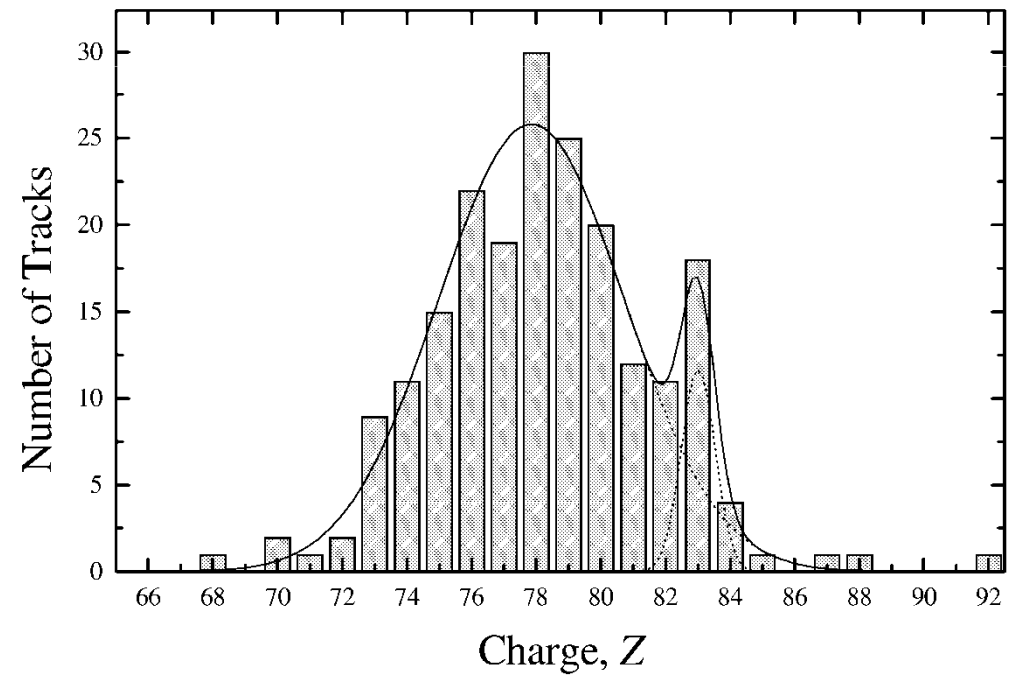
**Ariel 6**; P.H. Fowler et al. 1987  $Z \geq 70$ .  
Спутник, 625 км, 427 дней,  
электроника



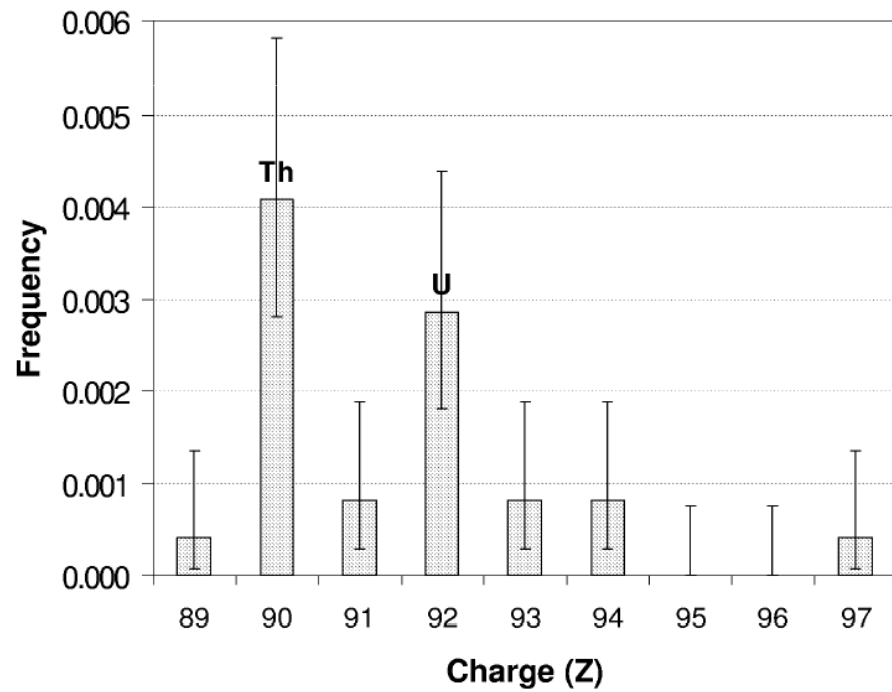
**HNE (HEAO 3);** W.R. Binns et al.  
1989  $Z \geq 50$ . Спутник, 495 км, 454  
дня, электроника



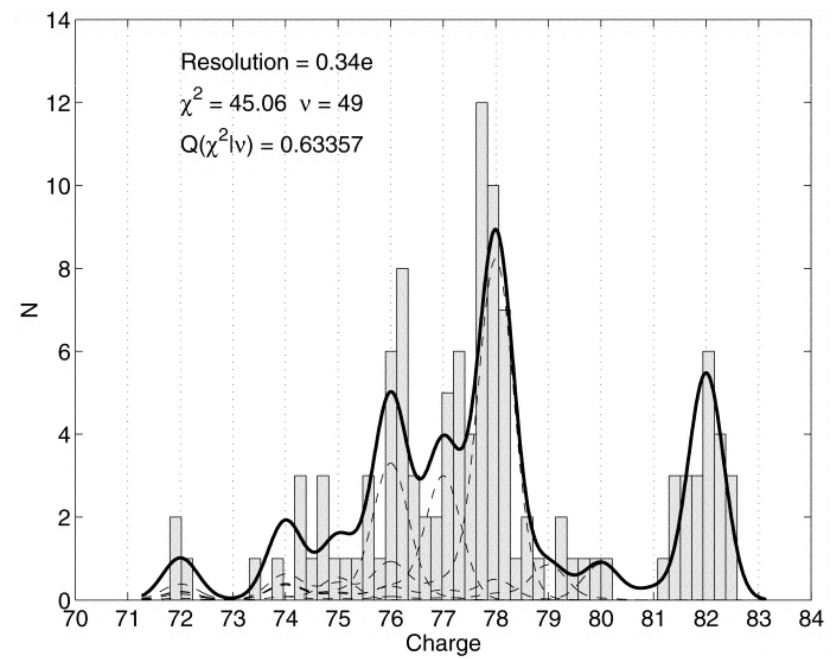
**UHCRES;** C. Domingo, J. Font et al. 1995  
 $Z \geq 68$ . Спутник, 450 км, 6 лет, пластик  
(стопки лексана)



**UHCRC**; J. Donnelly, ICRC-2001



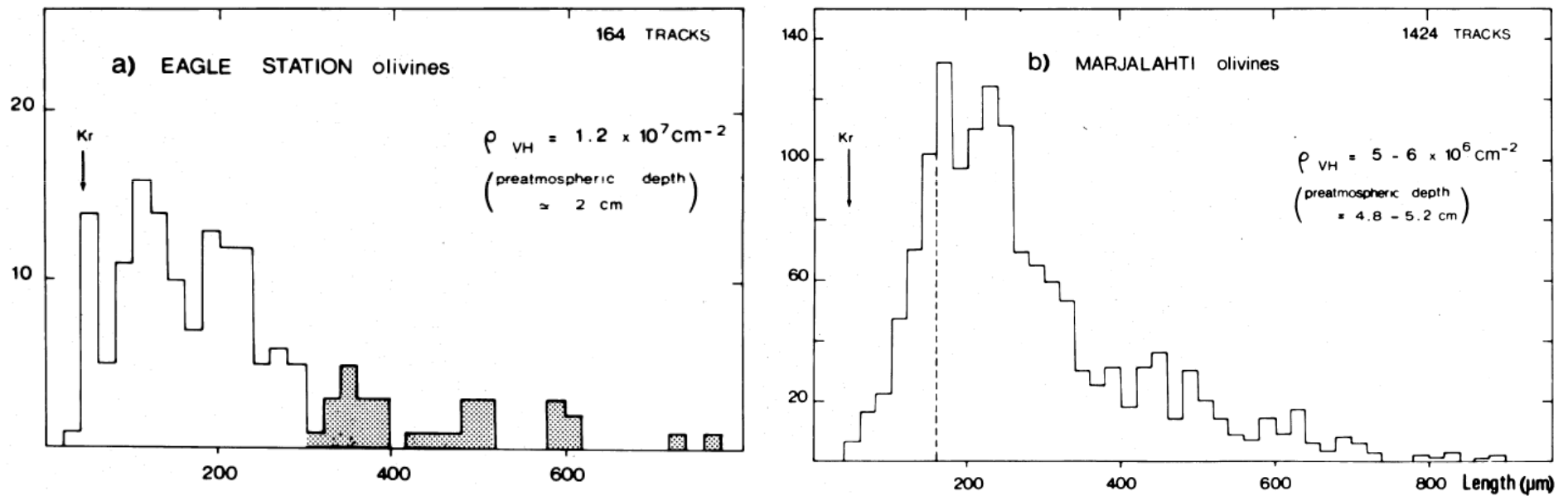
**Trek**; B. A. Weaver, A. J. Westphal.  
Станция "Мир", 5 лет, 450 км,  
СТОПКИ СТЕКОЛ



# Перелыгин (ОИЯИ): травление кристаллов оливина из метеоритов методом трек в треке.

Начало работ 1975

OTGONSUREN *ET AL.*



Позднее дополнили метод предварительным отжигом

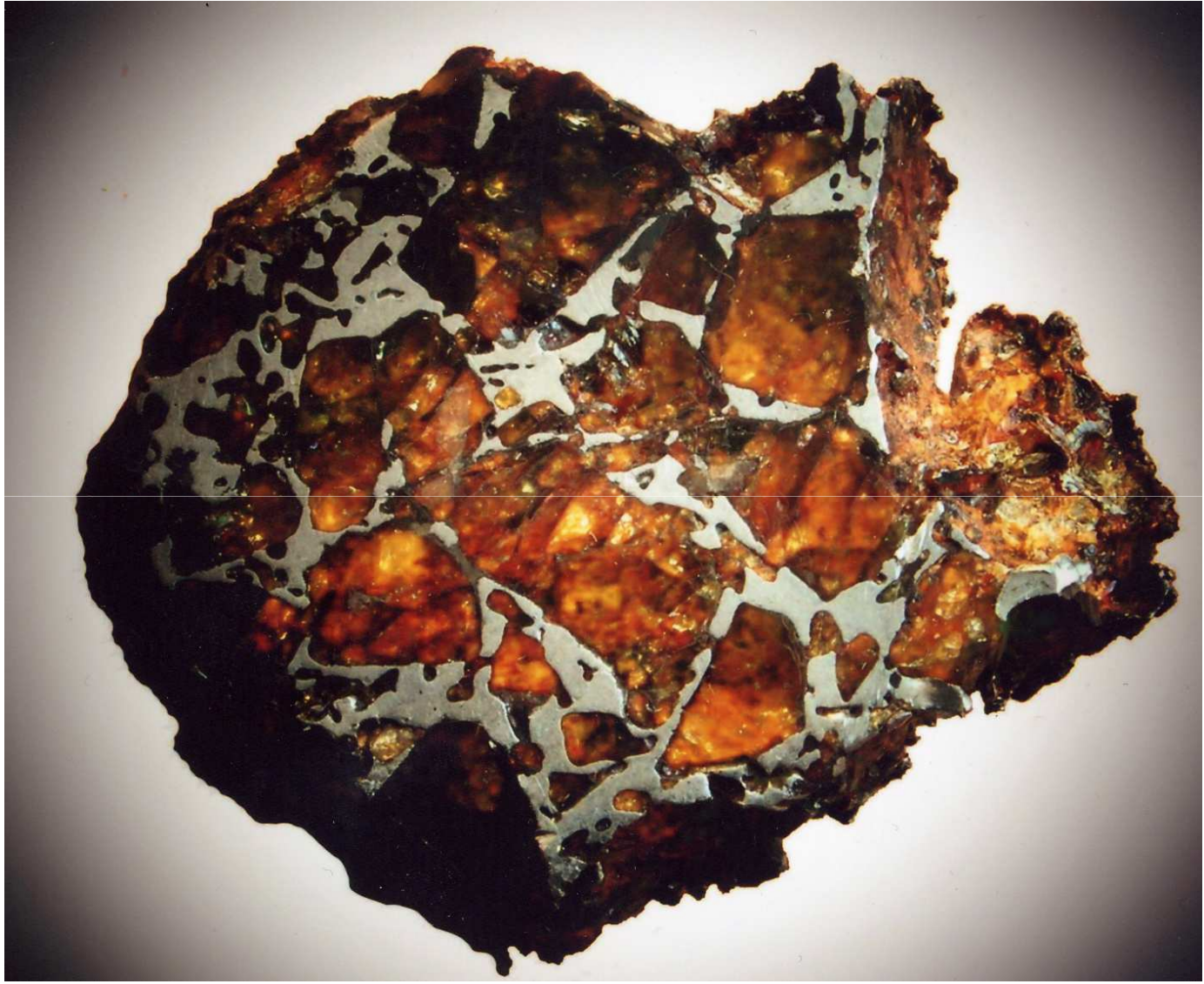
Метеорит Eagle Station (США 1880 г.).

Размер ~ 25 см. Вес ~ 38 кг. 300 млн лет

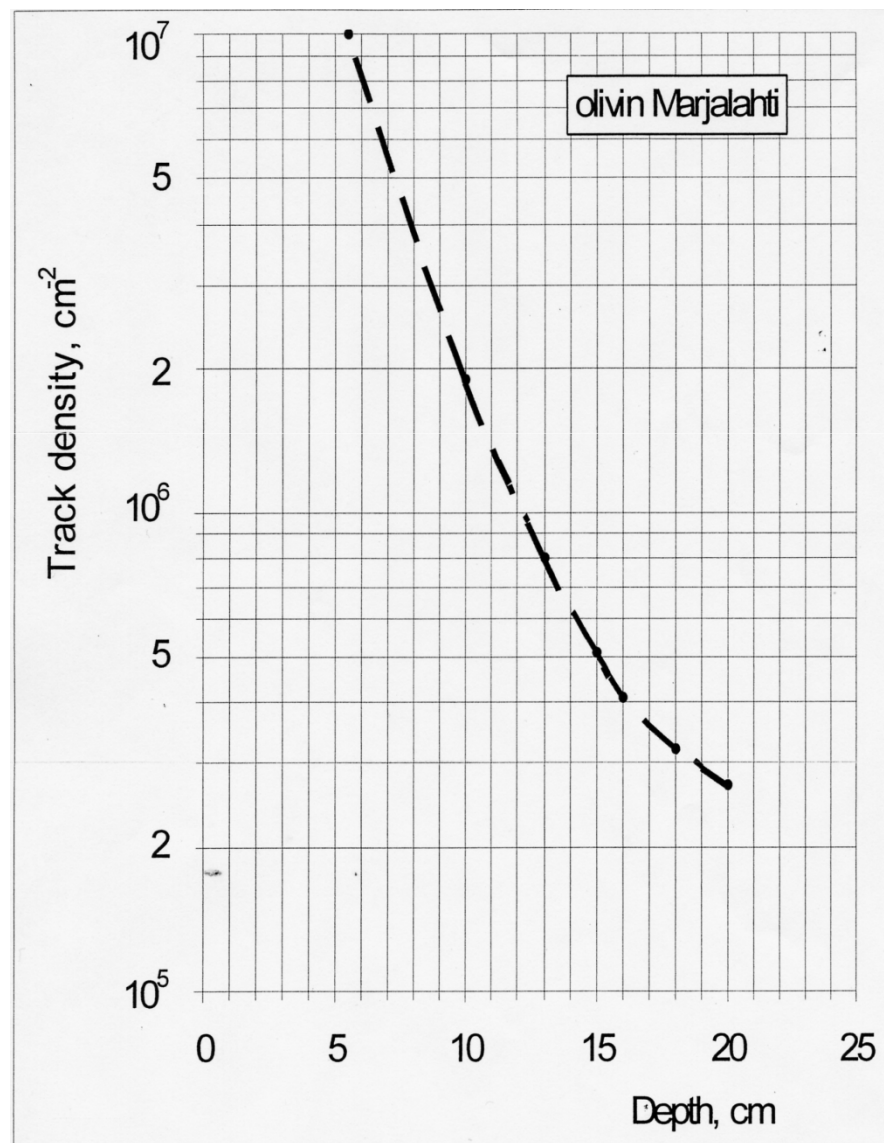
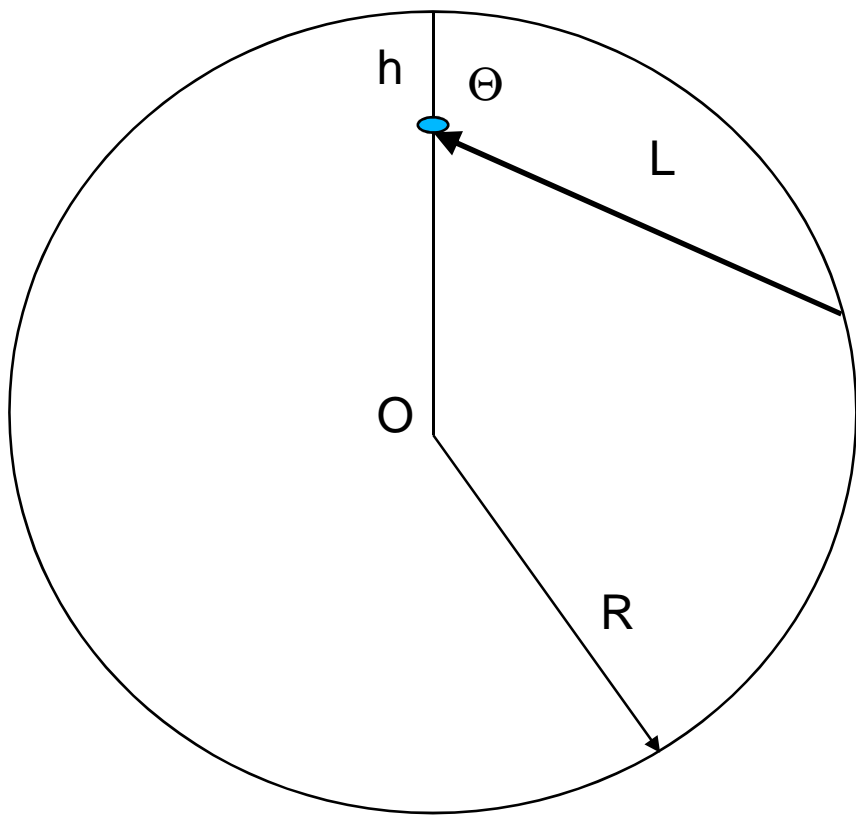








# Схема оценки энергии первичной частицы



# Распределение направлений треков тяжёлых ядер на плоскости углов $\theta$ и $\phi$ в образце оливина

